

FE-SULFIDES IN RYUGU PARTICLE C0025_01: A COMPARISON WITH CI AND CY CHONDRITESC. S. Harrison^{1,2}, A. J. King¹, R. H. Jones², S. S. Russell¹, T. Nakamura³, H. Yurimoto⁴, T. Noguchi⁵, R. Okazaki⁶, H. Yabuta⁷, H. Naraoka⁶, K. Sakamoto⁸, S. Tachibana⁹, S. Watanabe¹⁰ and Y. Tsuda⁸¹Natural History Museum, London. E-mail: c.harrison@nhm.ac.uk. ²Univ. of Manchester. ³Tohoku Univ. ⁴Hokkaido Univ. ⁵Kyoto Univ. ⁶Kyushu Univ. ⁷Hiroshima Univ. ⁸ISAS, JAXA. ⁹The Univ. of Tokyo. ¹⁰Nagoya Univ.

Introduction: Remote observations of C-type asteroid 162173 Ryugu show a weak absorption band at $\sim 2.72 \mu\text{m}$ suggesting abundant Mg-rich phyllosilicates and a close affinity to the aqueously altered CI and/or thermally metamorphosed CY carbonaceous chondrites [1]. In December 2020, JAXA's Hayabusa2 mission successfully returned $>5 \text{ g}$ of material from the surface and sub-surface of Ryugu to Earth. Initial characterization of the pristine samples indicates that they are most similar to the CI chondrites [2]. The morphology, composition, and abundance of Fe-sulfides are distinctive features of the CI and CY carbonaceous chondrites, reflecting different degrees of aqueous and thermal processing on their parent bodies [3-6]. Here, we compare the coarse ($>10 \mu\text{m}$) Fe-sulfide grains in a Ryugu particle with those found within the matrices of CI and CY chondrites.

Methods: The ten largest Fe-sulfide grains in polished sections of Ryugu particle C0025_01, the CI chondrites Ivuna and Orgueil, and the CY chondrites Yamato (Y)-82162 (Stage III, $\sim 500 - 750^\circ\text{C}$) and Y-86720 (Stage IV, $>750^\circ\text{C}$) were characterized using a ZEISS EVO 15LS scanning electron microscope (SEM) equipped with an energy dispersive X-ray spectrometer (EDS). Backscattered electron (BSE) images, element maps, and spot analyses were acquired at 20 kV with a beam current of 1.5 nA. Additional BSE images of Fe-sulfide grains in Ryugu particle C0025_01 were obtained using an FEI Quanta field emission SEM at 9 kV.

Results: The ten largest Fe-sulfide grains in Ryugu particle C0025_01, and the CI and CY chondrites, are typically euhedral to subhedral and lath-like in shape, although some irregular grains are also present (Fig. 1). The size range of the largest Fe-sulfide grains in C0025_01 ($\sim 30 - 70 \mu\text{m}$) is similar to Ivuna ($\sim 40 - 80 \mu\text{m}$) and Orgueil ($50 - 150 \mu\text{m}$, where 70 % of the grains are $<80 \mu\text{m}$). In comparison, the Fe-sulfide grains are larger in Y-82162 ($\sim 75 - 130 \mu\text{m}$) and Y-86720 ($\sim 190 - 330 \mu\text{m}$). The abundance of Fe-sulfide grains $>10 \mu\text{m}$ in size in Ryugu ($\sim 1 \text{ area}\%$) differs to that of both the CI chondrites ($\sim 0.13 - 0.3 \text{ area}\%$) and the CY chondrites ($\sim 1.5 - 2.0 \text{ area}\%$). The Fe-sulfide grains in Ryugu particle C0025_01 are pyrrhotite $[(\text{Fe},\text{Ni})_{1-x}\text{S}]$ with a composition close to $(\text{Fe},\text{Ni})_7\text{S}_8$, where the average value of x is 0.13. The pyrrhotite grains have an average Ni content of $1.80 \pm 0.86 \text{ wt.}\%$ and an Fe/S ratio of $0.85 \pm 0.01 \text{ at.}\%$ (no. of analyses = 6). This is consistent with the composition of the Fe-sulfide grains in Ivuna ($n = 18$) and Orgueil ($n = 31$), where the average value of x is 0.12, the Ni content is $\sim 1.15 \text{ wt.}\%$, and the Fe/S ratio is ~ 0.86 (similar to the results of [6]). Three of the Fe-sulfide grains in Ryugu particle C0025_01 contain $<10 \mu\text{m}$ sized irregular inclusions of pentlandite $[(\text{Fe},\text{Ni})_9\text{S}_8]$ that were not observed in either Ivuna or Orgueil. However, Bullock et al. [3] reported rare $\sim 10 \mu\text{m}$ sized pentlandite inclusions in Ivuna (plus the CI chondrites Alais and Tonk). The composition of Fe-sulfides in CY chondrites is clearly distinct from Ryugu and the CIs; the laths are composed of stoichiometric troilite (FeS), and all grains contain $<12 \mu\text{m}$ sized inclusions of pentlandite. In Y-82162 the pentlandite inclusions are oriented laths, whereas in Y-86720 they are unoriented irregular blebs that often occur adjacent to metal inclusions.

Discussion: The largest Fe-sulfide grains in Ryugu particle C0025_01 are compositionally and morphologically very similar to those in the CI chondrites, providing further evidence for a close relationship between these materials. The Fe-sulfides in Ryugu and the CI chondrites most likely formed during near-complete low temperature (100°C) aqueous alteration. However, the presence of pentlandite inclusions in three of the ten largest Fe-sulfide grains suggests that hydration of Ryugu particle C0025_01 was less extensive than Ivuna and Orgueil [3]. The Fe-sulfides in Ryugu differ in size, composition, and morphology from those in the thermally metamorphosed CY chondrites, indicating that they were never heated to temperatures $>500^\circ\text{C}$ post-hydration [4,5,7].

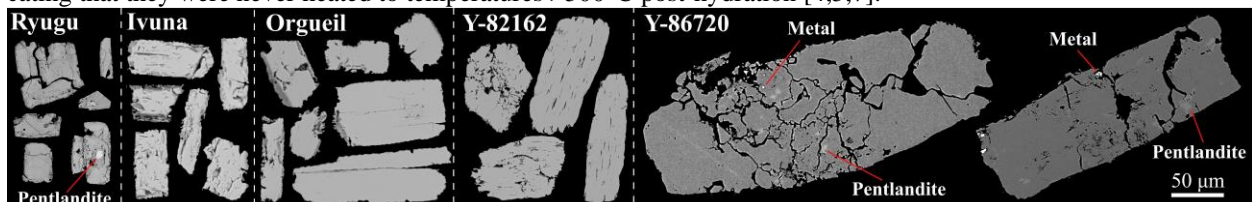


Figure 1. Selected Fe-sulfide grains from each section: scale bar of $50 \mu\text{m}$ is the same for all panels.

References: [1] Kitazato A. et al. (2019) *Science*. 364:272. [2] Yada T. et al. (2022) *Nat. Astro.* 6:214. [3] Bullock E. S. et al. (2005) *Geochim. et Cosmochim. Acta.* 69:2687. [4] Kimura M. et al. (2011) *Meteorit. & Planet. Sci.* 46:431. [5] King A. J. et al. (2019) *Geochem.* 79:125531. [6] Schrader D. L. et al. (2021) *Geochim. et Cosmochim. Acta.* 303:66. [7] Nakamura T. et al. (2005) *J. of Mineralog. & Petrolog. Sci.* 100:260.