

HAYABUSA2: CURRENT SUMMARY.

T. Nakamura¹, S. Watanabe^{2,3}, M. Hirabayashi⁴, N. Hirata⁵, N. Hirata⁶, R. Noguchi³, Y. Shimaki³, H. Ikeda⁷, E. Tatsumi⁸, M. Yoshikawa³, S. Kikuchi³, H. Yabuta⁹, S. Tachibana^{8,2}, Y. Ishihara^{3†}, T. Morota², K. Kitazato⁵, N. Sakatani³, K. Matsumoto^{10,11}, K. Wada¹², H. Senshu¹², C. Honda⁵, T. Michikami¹³, H. Takeuchi³, T. Kouyama¹⁴, R. Honda¹⁵, S. Kameda¹⁶, T. Fuse¹⁷, H. Miyamoto⁸, G. Komatsu^{18,12}, S. Sugita⁸, T. Okada³, N. Namiki^{10,11}, M. Arakawa⁶, M. Ishiguro¹⁹, M. Abe³, R. Gaskell²⁰, E. Palmer²⁰, O. S. Barnouin²¹, P. Michel²², A. S. French²³, J. W. McMahon²³, D. J. Scheeres²³, P. A. Abell²⁴, Y. Yamamoto³, S. Tanaka³, K. Shirai³, M. Matsuoka³, M. Yamada¹², Y. Yokota³, H. Suzuki²⁵, K. Yoshioka⁸, Y. Cho⁸, S. Tanaka⁶, N. Nishikawa⁶, T. Sugiyama⁵, H. Kikuchi⁸, R. Hemmi⁸, T. Yamaguchi^{3††}, N. Ogawa³, G. Ono⁷, Y. Mimasu³, K. Yoshikawa⁷, T. Takahashi³, Y. Takei³, A. Fujii³, C. Hirose⁷, T. Iwata^{3,11}, M. Hayakawa³, S. Hosoda³, O. Mori³, H. Sawada³, T. Shimada³, S. Soldini³, H. Yano³, R. Tsukizaki³, M. Ozaki^{3,11}, Y. Iijima^{3‡}, K. Ogawa⁶, M. Fujimoto³, T.-M. Ho²⁶, A. Moussi²⁷, R. Jaumann²⁸, J.-P. Bibring²⁹, C. Krause³⁰, F. Terui³, T. Saiki³, S. Nakazawa³, Y. Tsuda³, ¹Tohoku University, Sendai 980-8578, Japan, ²Nagoya University, Nagoya 464-8601, Japan, (seicoro@eps.nagoya-u.ac.jp), ³Institute of Space and Astronautical Science, JAXA, Japan, ⁴Auburn University, Auburn, AL 36849, USA, ⁵University of Aizu, Aizu-Wakamatsu 965-8580, Japan, ⁶Kobe University, Kobe 657-8501, Japan, ⁷Research and Development Directorate, JAXA, Sagami-hara 252-5210, Japan, ⁸University of Tokyo, Tokyo 113-0033, Japan, ⁹Hiroshima University, Higashi-Hiroshima 739-8526, Japan, ¹⁰National Astronomical Observatory of Japan, Mitaka 181-8588, Japan, ¹¹SOKENDAI (The Graduate University for Advanced Studies), Hayama 240-0193, Japan, ¹²Chiba Institute of Technology, Narashino 275-0016, Japan, ¹³Kindai University, Higashi-Hiroshima 739-2116, Japan, ¹⁴National Institute of Advanced Industrial Science and Technology, Tokyo 135-0064 Japan, ¹⁵Kochi University, Kochi 780-8520, Japan, ¹⁶Rikkyo University, Tokyo 171-8501, Japan, ¹⁷National Institute of Information and Communications Technology, Kashima 314-8501, Japan, ¹⁸Università d'Annunzio, 65127 Pescara, Italy, ¹⁹Seoul National University, Seoul 08826, Korea, ²⁰Planetary Science Institute, Tucson, AZ 85710, USA, ²¹Johns Hopkins University Applied Physics Laboratory, Laurel, MD 20723, USA, ²²Université Côte d'Azur, Observatoire de la Côte d'Azur, CNRS, Laboratoire Lagrange, 06304 Nice, France, ²³University of Colorado, Boulder, CO 80309, USA, ²⁴NASA Johnson Space Center, Houston, TX 77058, USA, ²⁵Meiji University, Kawasaki 214-8571, Japan, ²⁶DLR (German Aerospace Center), Institute of Space Systems, 28359 Bremen, Germany. ²⁷CNES (Centre National d'Etudes Spatiales), 31401 Toulouse, France, ²⁸DLR, Institute of Planetary Research, 12489 Berlin-Adlershof, Germany, ²⁹Institute d'Astrophysique Spatiale, 91405 Orsay, France, ³⁰DLR, Microgravity User Support Center, 51147 Cologne, Germany. †Current affiliation: National Institute for Environmental Studies, Tsukuba 305-8506, Japan, ††Current affiliation: Mitsubishi Electric Corporation, Kamakura 247-8520, Japan, ‡Deceased. (email: tomoki@m.tohoku.ac.jp)

Hayabusa2 is the 2nd Japanese sample return mission from C-type asteroid 162173 Ryugu. The purpose of the mission is to uncover the early solar system evolution and to obtain crucial information for the origin of life and water in the solar system. So far (till 2019 April), the science operations of the spacecraft went very well: we succeeded spectroscopic, morphological, and geological observation of an entire region of the asteroid for global mapping from Box-A (20km high) and -C (5-7km high), first touchdown on a part of the equatorial ridge after several time rehearsals, deployment of the small landers such as MASCOT, and injection of a metallic Cu plate (Small Carry-on Impactor) onto the asteroid surface. Characteristic spinning-top shape and very low density (1.2 g/cc) indicate that Ryugu is a rubble pile asteroid formed by a catastrophic impact and deformed by high spin rates sometime after the impact [1]. Spectroscopic observation shows that Ryugu is very dark object (< 2% reflectance at 0.55 μ m with standard-measurement condition [1, 2]) and has a small 2.7 μ m (~10% absorption) band suggestive of a global distribution of phyllosilicates [3]. The closest analogue would be partially-dehydrated carbonaceous chondrites or C-rich interplanetary dust particles [2,3]. Ryugu's surface shows slight color variations: bluer areas are located at regions such as equatorial ridge and poles [2], although we do not fully understand what is responsible for the color variation. On the day of Feb 22th 7:29am (JST), HY2 spacecraft succeeded to touchdown on the surface area of L08-E1 that is very close to the planned touchdown point (deviation is ~1m), suggesting that the spacecraft was very well controlled. A metallic projectile was fired when the sample horn touched the surface and numerous rock fragments were despersed from the surface and came to the horn, which we observed by CAM-H camera. Therefore, it is expected to some rock fragments to reach a catcher in the spacecraft. The sample recovery capsule will be coming back to the Earth in December 2020 and the initial analysis of the returned samples will start from spring 2021.

References: [1] Watanabe S. et al. (2019) *Science* 364, 10.1126/science.aav8032. [2] Sugita S. et al. (2019) *Science* 364, 10.1126/science.aaw0422. [3] Kitazato et al. (2019) *Science* 364, 10.1126/science.aav7432.