

**AN ION MICROPROBE STUDY OF BE-B ISOTOPE SYSTEMATICS IN MELILITE-RICH CAIS BASED ON NEWLY DETERMINED Be/B RELATIVE SENSITIVITY FACTORS FOR MELILITIC GLASS STANDARDS.**

K. Fukuda<sup>1</sup>, W. Fujiya<sup>2</sup>, H. Hiyagon<sup>1</sup>, N. Sugiura<sup>1</sup>, N. Takahata<sup>3</sup>, and Y. Sano<sup>3</sup>. <sup>1</sup>Department of Earth and Planetary Science, The University of Tokyo (k.fukuda@eps.s.u-tokyo.ac.jp), <sup>2</sup>College of Science, Ibaraki University, <sup>3</sup>Atmosphere and Ocean Research Institute, The University of Tokyo.

**Introduction:** Beryllium-10, which decays to <sup>10</sup>B with a half-life of 1.4 Myr [1], is produced almost exclusively by spallation reactions induced by solar or galactic cosmic rays. Recent ion microprobe studies have demonstrated that CAIs in several types of carbonaceous chondrites record a wide range of initial <sup>10</sup>Be/<sup>9</sup>Be ratios [2-4]. This suggests that variable irradiation processes and/or conditions have existed in the early solar system. However, in the ion microprobe analyses, a Be/B relative sensitivity factor (RSF) for a melilite composition has not been well established due to the absence of suitable standards. In order to overcome this problem, we synthesized glass standards with melilitic compositions with different Be/B ratios. Here we report preliminary results of RSF and Be-B systematics of melilite-rich CAIs in CO, CH, and ungrouped C chondrites.

**Synthesis of glass standards with variable Be/B ratios:** We synthesized three glasses with melilitic compositions, which were doped with Be and B with variable Be/B ratios. Their major element concentrations were determined using FE-EPMA (JEOL-8530F, The Univ. of Tokyo). Furthermore, Be concentrations were measured using LA-ICPMS (ICAP Q, The Univ. of Tokyo). Unfortunately, B concentrations could not be well determined at this time. In this abstract, therefore, we estimated B concentrations by SIMS (see below).

**SIMS analysis:** B concentrations and <sup>9</sup>Be/<sup>11</sup>B ratios were determined by a NanoSIMS 50 at Atmosphere and Ocean Research Institute (AORI), The Univ. of Tokyo. A primary O<sup>-</sup> beam of 1 nA was focused to ~5 μm in diameter and rastered over 10 x 10 μm<sup>2</sup>. <sup>9</sup>Be<sup>+</sup>, <sup>10</sup>B<sup>+</sup>, <sup>11</sup>B<sup>+</sup>, <sup>30</sup>Si<sup>+</sup>, and <sup>43</sup>Ca<sup>+</sup> ions were detected simultaneously in the multi-collection mode. B concentrations of synthetic glasses were estimated using <sup>11</sup>B/<sup>30</sup>Si ratios in comparison with that of NIST 610 glass. Previous studies showed that matrix effects on the B sensitivity are small for minerals with diverse chemical compositions [5,6]. Therefore, in our preliminary results, we estimated B concentrations of synthetic glasses by SIMS measurements using NIST 610.

**RSF:** The RSF for melilite-like glasses was estimated using calibration curves determined from three melilite-like glasses with different Be/B ratios (<sup>9</sup>Be/<sup>11</sup>B ~7, 18, and 44). The RSF (<sup>9</sup>Be/<sup>11</sup>B)/(<sup>9</sup>Be/<sup>11</sup>B) is  $2.36 \pm 0.12$  (2σ) in favor of Be ions relative to B ions. Under the same analytical condition, we also obtained the RSF for NIST 610 glass ( $2.73 \pm 0.25$ ; 2σ), which is ~16 % higher than that of melilite-like glasses.

**Initial <sup>10</sup>Be/<sup>9</sup>Be in melilite-rich CAIs:** We also investigated Be-B systematics of melilite-rich CAIs in Y81020 (CO3.05), SaU290 (CH3), and Y82094 (ungrouped C3.2) chondrites using the NanoSIMS 50 at AORI. Melilites in Y81020 and Y82094 show excesses in <sup>10</sup>B and the inferred initial <sup>10</sup>Be/<sup>9</sup>Be ratios ranged from 0.9 to  $2.0 \times 10^{-3}$ , slightly higher than those of CV CAIs (e.g.,  $0.88 \times 10^{-3}$  [8]). If these CAIs formed contemporaneously, the variation in the <sup>10</sup>Be/<sup>9</sup>Be ratio observed in this study supports <sup>10</sup>Be production by local irradiation processes [e.g., 4, 7]. Although analytical uncertainties are large, melilites in SaU290 have the inferred initial <sup>10</sup>Be/<sup>9</sup>Be ratios ranged from 0.4 to  $1.1 \times 10^{-3}$ . Recently, Gounelle et al. [3] found that the highest initial <sup>10</sup>Be/<sup>9</sup>Be ratio has been reported for an Isheyevo (CB/CH) CAI. Dunham et al. [4] found that the lowest one has been reported for a NWA5028 (CR2) CAI [4]. Given our and these previous results, CB, CH, and CR CAIs may represent more a wide range of initial <sup>10</sup>Be/<sup>9</sup>Be ratios than CV CAIs.

**References:** [1] Korschinek G. et al. 2010. *Nuclear Instruments and Methods in Physics Research B* 268:187-191. [2] Wielandt D. et al. 2012. *The Astrophysical Journal Letters* 748: L25 (7pp). [3] Gounelle M. et al. 2013. *The Astrophysical Journal Letters* 763: L33 (5pp). [4] Dunham E. et al. 2016. Abstract #2723. 47th Lunar and Planetary Science Conference. [5] Chaussidon M. and Libourel G. 1993. *Geochimica et Cosmochimica Acta* 57:5053-5062. [6] Sugiura N. et al. 2001. *Meteoritics & Planetary Science* 36:1397-1408. [7] McKeegan K. D. et al. 2000. *Science* 289: 1334-1337. [8] Chaussidon M. et al. 2006. *Geochimica et Cosmochimica Acta* 70:224-245.