

Mn-Cr dating of enstatite chondrites

J. Hopp^{1,2}, M. Trierloff^{1,2}, J.-C. Storck^{1,3}, T. Ludwig¹, H.-P. Meyer¹, R. Altherr¹ and A. El Goresy⁴, ¹Institut für Geowissenschaften, Universität Heidelberg, Im Neuenheimer Feld 234-236, D-69120 Heidelberg, jens.hopp@geow.uni-heidelberg.de, ²Klaus-Tschira-Labor für Kosmochemie, Universität Heidelberg, Im Neuenheimer Feld 234-236, D-69120 Heidelberg, Germany, ³Institut für Geochemie und Petrologie, ETH Zürich, Sonneggstrasse 5, 8092 Zürich, Switzerland, ⁴Bayerisches Geoinstitut, Universität Bayreuth, Universitätsstrasse 30, D-95440 Bayreuth, Germany.

Introduction: The ^{53}Mn - ^{53}Cr system (half-life 3.7 Ma) is a short-lived decay system that can be used as a relative chronometer of early solar system processes. Ideally, if $^{53}\text{Cr}/^{52}\text{Cr}$ -ratios in a sample are linearly correlated with the associated ratio of (stable) $^{55}\text{Mn}/^{52}\text{Cr}$ the excess ^{53}Cr can be interpreted to be the result of in situ decay of ^{53}Mn . The respective slope of this isochron corresponds to the initial $^{53}\text{Mn}/^{55}\text{Mn}$ of the sample. This initial $^{53}\text{Mn}/^{55}\text{Mn}$ ratio can be translated into an absolute age by comparing it with the initial $^{53}\text{Mn}/^{55}\text{Mn}$ ratio of a standard meteorite of known absolute age. Commonly used is the LEW 86010 angrite Mn-Cr age standard with an absolute U-Pb age of 4557.8 ± 0.5 Ma and an initial $^{53}\text{Mn}/^{55}\text{Mn}$ ratio of $(1.25 \pm 0.07) \cdot 10^{-6}$ (errors are 1σ -uncertainties) [1].

Mn and Cr are both rather abundant elements and assumed to have a relatively low diffusivity. Application of the Mn-Cr dating method is limited to few carrier phases with a low partitioning behavior of Cr and high Mn/Cr ratios and hence, besides other problematic issues, is still rarely applied in early solar system chronology. Enstatite chondrites contain a large variety of exotic mineral phases formed at low $f(\text{O}_2)$, in particular Mn-bearing sulfides, which might be suited for Mn-Cr dating. The aim of our project is to add more cosmochronological constraints on the early evolution of the enstatite chondrite parent bodies.

Methods: In this study we screened several distinct mineral phases for potential ^{53}Cr -excesses in enstatite chondrites at the new Cameca 1280 HR ion probe facility in Heidelberg, Germany. We used a focussed beam mode and the primary beam current commonly was in the range of 2 to 3 nA (equivalent to a spot diameter of about 3-5 μm). The field aperture was adjusted to values down to about 500 in order to avoid potentially nearby Cr-bearing phases. Element ratios were obtained by electron microprobe analyses with a Cameca SX51 device.

Samples and results: We analyzed the enstatite chondrites Sahara 97072 (EH 3), Indarch (EH 4), EET 96135 (EH 4/5), LAP 02225 (EH impact melt), MAC 88136 (EL 3) and Neuschwanstein (EL 6). For some of these chondrites age data from different chronometers already exist (e.g., Mn-Cr [1], [2], [3] I-Xe [4], [5], [6], Ar-Ar [7]). The major Mn-bearing phases in EH and EL chondrites are niningerite and alabandite, respectively. However, Mn/Cr ratios in both phases always were below 400, i.e., far below required ratios to detect ^{53}Cr -excesses. Hence, $^{53}\text{Cr}/^{52}\text{Cr}$ -ratios are indistinguishable from the $^{53}\text{Cr}/^{52}\text{Cr}$ -ratio determined by analyses of Cr-rich phases troilite and daubréelite (Mn/Cr < 1). Results of troilite and daubréelite analyses were intended for determination of the relative sensitivity factor (rsf) for Mn/Cr (= ratio of Mn/Cr compositions determined by electron microprobe versus ion probe signals). However, the fine-grained nature of daubréelite in most meteorites and the relatively low amount of Mn in troilite made a consistent determination of rsf-values challenging. In case of LAP 02225 a reproducible rsf-value of 1.43 ± 0.01 (two analyses of daubréelite) could be obtained. Only in sphalerite we found significant excess ^{53}Cr associated with high $^{55}\text{Mn}/^{52}\text{Cr}$ -ratios (up to 10^7). However, commonly data do not define a distinct isochron but rather show some scatter. Most sphalerite data were obtained for the impact melt LAP 02225 for which a range in initial $^{53}\text{Mn}/^{55}\text{Mn}$ of $(1.5-3.5) \cdot 10^{-7}$ could be derived, corresponding to an absolute age range of 4546.5-4551.0 Ma (rsf=1.43) which agrees with I-Xe and Ar-Ar ages. This is consistent with fast cooling associated with simultaneous closure of all isotope chronometers, as expected for an impact melt rock. The initial $^{53}\text{Mn}/^{55}\text{Mn}$ derived for a single sphalerite datum of MAC 88136 is $(2.55 \pm 0.30) \cdot 10^{-7}$, equivalent to an absolute age of 4549.3 ± 0.6 Ma (for rsf=1). This is relatively young and may point to a weak disturbance of the Mn-Cr system by a thermal event. For Indarch we obtained an initial $^{53}\text{Mn}/^{55}\text{Mn} = (4.2 \pm 1.1) \cdot 10^{-6}$ (rsf=1), which corresponds to an absolute age of 4564.1 ± 1.5 Ma, in broad agreement with previous results [1], [2], [3]. Application of a rsf=1.43 as for LAP 02225 would result in 1.9 Ma younger sphalerite ages for both meteorites. A better determination of the Mn/Cr rsf-value for sphalerite is an important aspect of future work. In addition, more Mn-Cr analyses of EL chondrites are planned.

References: [1] Shukolyukov A. and Lugmair G. W. 2004. *Geochimica et Cosmochimica Acta* 68:2875-2888. [2] El Goresy et al. 1992. Abstract #1165. 23th Lunar & Planetary Science Conference. [3] Wadhwa M. et al. 1997. *Meteoritics & Planetary Science* 32:281-292. [4] Kennedy B. M. et al. 1988. *Geochimica et Cosmochimica Acta* 52:201-211. [5] Busfield A. et al. 2008. *Meteoritics & Planetary Science* 43:883-897. [6] Hopp J. et al. 2016. *Geochimica et Cosmochimica Acta* 174:196-210. [7] Hopp J. et al. 2014. *Meteoritics & Planetary Science* 49:358-372.