Pb ISOTOPIC AGE OF ALM-A – A FELDSPAR-RICH VOLCANIC ROCK FROM THE CRUST OF THE UREILITE PARENT BODY.

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**Introduction:** ALM-A (Almahata Sitta trachyandesitic meteorite) is a 24.2 gram SiO<sub>2</sub>-rich volcanic rock from the Almahata Sitta fall [1]. It is composed mainly of feldspar and Ca-rich and Ca-poor pyroxene [2,3]. Its affinity to ureilites is shown by O isotopic composition and mineralogy [3] and, therefore, this rock is thought to derive from the crust of the ureilite parent body.

A <sup>26</sup>Al-<sup>26</sup>Mg age of ALM-A was determined from multiple ion microprobe analysis of feldspar grains, which yielded an initial <sup>26</sup>Al/<sup>27</sup>Al=1.1±0.4×10<sup>-7</sup> [3], corresponding to an age of 6.5 Ma after CAI formation. This age is meaningful only if, at the time of CAI formation, the initial <sup>26</sup>Al/<sup>27</sup>Al of the ureilite precursor material was the same as that of CAIs. Furthermore, the feldspar grains from ALM-A analysed for <sup>26</sup>Al-<sup>26</sup>Mg show no correlated variations in Al/Mg and <sup>26</sup>Mg/<sup>24</sup>Mg ratios [3], hence the <sup>26</sup>Al-<sup>26</sup>Mg age is a model date, and as such is not as reliable as isochron ages.

Here we use Pb-isotopic data to independently constrain the crystallization age of ALM-A. Additional U-Pb analyses, as well as U isotopic and <sup>182</sup>Hf-<sup>182</sup>W studies are in progress.

**Techniques:** A 3.3 mg pyroxene-rich fraction and a 9.5 mg whole rock fraction with fragment size between 100 and 250 µm were leached in 0.5M HNO<sub>3</sub> (W1), 7M HNO<sub>3</sub> (W2), 6M HCl (W3) and 0.5M HF (W4), and dissolved in concentrated HF+HNO<sub>3</sub> (R). The leachates and residues were spiked, dissolved and analysed following [4]. The accuracy of analyses was verified by analyses of the synthetic EarlyTime ET1x solution that simulates isotopic composition of ancient radiogenic Pb [5,6]. This report is based on Pb isotopic analyses of HCl and HF washes and residues. The analyses of the HNO<sub>3</sub> washes and U concentration measurements are in progress.

**Results:** The HCl and HF washes from pyroxene and whole rock fractions contain moderately radiogenic Pb with <sup>206</sup>Pb/<sup>204</sup>Pb between 32 and 42. The residues contain less radiogenic Pb with <sup>206</sup>Pb/<sup>204</sup>Pb of 27.3 (pyroxene) and 11.9 (whole rock). A <sup>207</sup>Pb/<sup>206</sup>Pb vs. <sup>204</sup>Pb/<sup>206</sup>Pb isochron regression through five data points (all except pyroxene residue) yields an age of 4562.0±3.4 Ma (MSWD=5.1), assuming <sup>238</sup>U/<sup>235</sup>U=137.79. The isochron passes through the isotopic composition of the primordial Pb, signifying the absence of terrestrial contamination.

**Discussion:** The Pb-isotopic isochron age of ALM-A corresponds to formation of this rock at 5.3±3.4 Ma after CAIs [7]. It confirms the <sup>26</sup>Al-<sup>26</sup>Mg age of the same rock [3], and agrees with <sup>26</sup>Al-<sup>26</sup>Mg and <sup>53</sup>Mn-<sup>53</sup>Cr ages of other ureilites [8]. The consistency with the <sup>182</sup>Hf-<sup>182</sup>W ages of ureilites [9] suggests that crust formation on the ureilite parent body occurred shortly after mantle differentiation.

References: [1] Horstmann M. and Bischoff A. 2014. *Chem. Erde-Geochemistry* 74:149-184. [2] Bischoff A. et al. 2013. *Meteoritics Planet. Sci.* 48:A60. [3] Bischoff A. et al. 2014. *Proc. Natl. Acad. Sci.* 111:12689-12692. [4] Amelin Y. et al. 2010. *Earth Planet. Sci. Lett.* 300:343-350. [5] Connelly J.N. and Condon D.J. 2014. *Goldschmidt Conference*, abstract #448. [6] Huyskens M.H. et al. 2015. 46<sup>th</sup> LPSC, abstract #2340. [7] Connelly J.N. et al. 2012. *Science* 338:651–655. [8] Goodrich C.A. et al. 2010. *Earth Planet. Sci. Lett.* 295:531-540. [9] Budde G. et al. 2015. 46<sup>th</sup> LPSC, abstract #2293.