

**COORDINATED CHEMICAL AND ISOTOPIC IMAGING OF THE BELLS (CM2) METEORITE.**

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**Introduction:** Organic matter in carbonaceous meteorites preserves complex records of interstellar, nebular, and asteroidal chemistry [1]. Traditional geochemical methods have yielded detailed information on chemical properties of bulk samples. However, observations of large *H*, *C* and *N* isotopic anomalies in  $\mu\text{m}$ -scale organic nanoglobules suggest that organics from nebular/interstellar environments co-exist with organics formed or altered within the meteorite parent body [4]. To decode this record requires high spatial resolution chemical and isotopic analysis. We are performing *in situ* correlated chemical and isotopic mapping of the matrix of the Bells (CM2) meteorite. This meteorite was chosen because it contains abundant <sup>15</sup>N- and *D*-rich organic nanoglobules [2-4].

**Methods:** Bells (CM2) matrix grains were pressed into *Au* foil and imaged by optical and low-dose ultraviolet (UV) fluorescence microscopy. UV fluorescence reveals the distribution of aromatic and conjugated organic moieties and organic nanoglobules [5]. Organic chemical mapping of a 50×100  $\mu\text{m}$  region was performed at a spatial resolution of ~ 5-7  $\mu\text{m}$  using the JSC two-step laser mass spectrometer ( $\mu\text{-L}^2\text{MS}$ ). We used a vacuum UV (118 nm) laser ionization source to enable detection of virtually all classes of organics. SEM-EDX was used to establish the mineral composition of the mapped matrix fragments.

**Results & Discussion:** The matrix showed heterogeneous fluorescence emission from clusters and/or veins of sub- $\mu\text{m}$  globular features interpreted as organic nanoglobules. Organic maps of nanoglobule-poor and nanoglobule-rich matrix were dramatically distinct. In nanoglobule-rich regions the overall abundance of organic species was several fold higher than the surrounding matrix and was dominated by low weight *O*, *N* & *S* containing molecular species. The most prominent species include  $\text{NH}_3$ ,  $\text{H}_2\text{S}$ ,  $\text{CH}_3\text{SH}$ ,  $\text{CH}_2\text{O}$ ,  $\text{CH}_3\text{CHO}$  &  $\text{CH}_3\text{COCH}_3$ , and  $\text{CH}_3\text{CH}_2\text{CHO}$ . In contrast, nanoglobule-poor regions were an order of magnitude depleted in these hetero-organics, being composed of minimally alkylated 3- to 5-ring aromatics and higher molecular weight  $\text{C}_{12}\text{-C}_{28}$  hydrocarbons with variable degrees of unsaturation. The strong microscale chemical dichotomy provides evidence for preservation of pre-accretional chemical processes. The high abundance of  $\text{NH}_3$  in concert with simple aldehydes and ketones also supports the hypothesis that nanoglobules formed through formose-type condensation reactions [6]. Previously observed isotopic anomalies in organic nanoglobules suggests that this occurred by radiation processing of organic grains in nebular or interstellar environments [4].

**References:** [1] Sephton (2002) *Phil. Trans. R. Soc. A* 363:2729-2742. [2] Nakamura et al. (2002) *IJA* 1:179-189. [3] Garvie & Buseck *MAPS* 41:633-641. [4] Nakamura-Messenger et al. (2006) *Science* 314:1439-. [5] Clemett et al. (2009) *MAPS* 44:A52. [6] Cody et al. (2011) *PNAS* 108, 19171-19176.