

THERMAL PROPERTIES OF CHELYABINSK METEORITE.

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Introduction: Knowledge of bulk density and porosity of meteorites enables one to predict thermal properties of meteorites and their parent bodies [1-5]. Mineralogy, petrology and selected physical properties of Chelyabinsk chondrite (LL 5, S4, W0) have been studied since its fall on February 15th, 2013 [6, 7]. The aim of the study was to determine thermal properties of Chelyabinsk meteorite: specific heat capacity, volumetric heat capacity, thermal diffusivity, and thermal conductivity at room and low temperatures.

Methods: Specific heat capacity C_p was determined using relationship between C_p and bulk density ρ_{bulk} of meteorites [1], and using abundance and composition of constituent minerals of Chelyabinsk chondrite [6,7]. Thermal diffusivity D , and thermal conductivity K were determined using relationships: between D and ρ_{bulk} [2], between K and ρ_{bulk} [3], and between K and porosity P [4,5]. Literature data on mean bulk density $\rho_{bulk} = (3.32 \pm 0.09) \cdot 10^3 \text{ kg/m}^3$, mean porosity $P = (6.0 \pm 3.2)\%$ of the Chelyabinsk meteorite [6], and on the normative composition of the chondrite [7] were used in calculations.

Results: Specific heat capacity of Chelyabinsk meteorite predicted by $C_p(\rho_{bulk})$ dependence is $C_p = 701 \text{ J/(kg}\cdot\text{K)}$ at 300K. When the correcting factor c ($c=1.33 \pm 0.03$) responsible for the low temperature decrease in C_p values is taken into considerations we have $C_p = (527 \pm 12) \text{ J/(kg}\cdot\text{K)}$ at 200 K. Using data on the normative composition of Chelyabinsk chondrite [7], and constituent minerals thermal data gives: $C_p = 684 \text{ J/(kg}\cdot\text{K)}$ at 300 K, and $(514 \pm 12) \text{ J/(kg}\cdot\text{K)}$ at 200 K. The main contribution to Chelyabinsk C_p comes from olivine (60%), pyroxenes give 18%, feldspar 13%, troilite 5%, kamacite 2%, and the rest of the constituent minerals gives 2% contribution. The volumetric heat capacity of Chelyabinsk $C_{volumetric} = 2.33 \pm 0.03 \cdot 10^6 \text{ J/(m}^3\cdot\text{K)}$ at RT, and is close to the value characteristic of stony meteorites ($2.5 \cdot 10^6 \text{ J/(m}^3\cdot\text{K)}$) [1]. Thermal diffusivity D of Chelyabinsk is $(0.77 \pm 0.20) \cdot 10^{-6} \text{ m}^2/\text{s}$ at 200 K, and $(0.80 \pm 0.50) \cdot 10^{-6} \text{ m}^2/\text{s}$ at 300 K. Thermal conductivity K of Chelyabinsk is $4.5 \pm 0.4 \text{ W m}^{-1} \text{ K}^{-1}$ at 200 K, and $(4.1 \pm 1.4) \text{ W m}^{-1} \text{ K}^{-1}$ at 300 K. Calculations show that a significant spread noted in ρ_{bulk} , and P values of various Chelyabinsk samples ($\rho_{bulk} = (3.14-3.60) \cdot 10^3 \text{ kg/m}^3$, $P = 1.5-11.4\%$ [6]) leads to the range of values: $C_p = (670-723) \text{ J/(kg}\cdot\text{K)}$, $K = (1-8) \text{ W m}^{-1} \text{ K}^{-1}$, $D = (0.1-1.8) \cdot 10^{-6} \text{ m}^2/\text{s}$ at 300 K, i.e. to the range of values expected for direct measurements.

Conclusions: Estimated thermal properties of Chelyabinsk LL5 chondrite are comparable with thermal properties of Kilabo (LL6), NWA 4560 (LL3.2), and Softmany (L6) chondrites.

References: [1] Szurgot M. 2011. Abstract #1150.pdf. 42nd Lunar & Planetary Science Conference. [2] Szurgot M. and Wojtatowicz T.W. 2011. Abstract #5036. 74th Annual Meeting of the Meteoritical Society. [3] Szurgot M. 2011. Abstract #5074. 74th Annual Meeting of the Meteoritical Society. [4] Opeil C.P. et al. 2012. *Meteoritics & Planetary Science* 47:319-329. [5] Szurgot M. et al. 2012. *Meteorites* 2:53-65. [6] Kohout T. et al. 2014. *Icarus* 228:78-85. [7] Galimov E. M. et al. 2013. *Geochemistry International* 51:522-539.