

**MORPHOLOGY, STRUCTURAL, AND SPECTRAL CHARACTERIZATIONS OF LAB-SYNTHESIZED MANGANESE OXIDES RELEVANT TO MARS.** Yanqing Xin<sup>1,2</sup>, Zongcheng Ling<sup>1\*</sup>, Haijun Cao<sup>1</sup>, Changqing Liu<sup>1</sup>, Hongchun Bai<sup>1</sup>. <sup>1</sup>Shandong Provincial Key Laboratory of Optical Astronomy and Solar-Terrestrial Environment, School of Space Science and Physics, Institute of Space Sciences, Shandong University, Weihai, Shandong, 264209, China. <sup>2</sup>Key Laboratory of Lunar and Deep Space Exploration, CAS. (zcling@sdu.edu.cn).

**Introduction:** Manganese is one of the important transition metals on Earth and Mars. On Earth, Mn is relatively abundant with an average upper crustal abundance of ~0.19 wt % MnO[1] while is ~0.4 wt % in basalts on Mars [2]. The Mn is a redox-sensitive element and has several oxidation states, such as +2, +3, +4, +6, and +7. On those Mn oxidation states, Mn<sup>2+</sup> is soluble, with Mn<sup>3+</sup> and Mn<sup>4+</sup> forming minerals in the form of Mn<sub>2</sub>O<sub>3</sub> and MnO<sub>2</sub> precipitation. Therefore, Mn is an important indicator of planetary mineral and geologic evolution because its valence change is more complex.

Mars rovers Curiosity and Opportunity have found Mn-rich minerals in multiple sedimentary outcrops at Gale and Endeavour Crater, respectively [3,4], but the exact mineral species still need to be confirmed. In the Kimberley region at Gale Crater, the content of Mn of fracture-filling materials that crosscut sandstone was found to be higher than 25%, indicating the existence of Mn-oxides rather than evaporites or other salts [5]. Furthermore, hydrated Mn-oxides were also found in Martian breccia meteorites NWA 7034 and NWA 7533. Those phases contain Mn contents up to ~32% and experimental studies show that the Mn-oxide minerals are Mn(IV)-oxides, but it is hard to identify the mineral species (birnessite or vernadite) due to its fine grain size and the aggregate nature [6].

In this work, we focus on the laboratory spectroscopic studies of seven pure Mn-oxides ( $\alpha$ -,  $\beta$ -,  $\gamma$ -,  $\delta$ -MnO<sub>2</sub>,  $\alpha$ -Mn<sub>2</sub>O<sub>3</sub>,  $\alpha$ -Mn<sub>3</sub>O<sub>4</sub>,  $\gamma$ -MnOOH). They were synthesized with hydrothermal reaction and solid-state calcination methods, using SEM, EDS to get their morphology and chemical composition, XRD to the identification of Mn-oxides. Then the spectroscopy technique such as Raman was used to characterize their spectral characteristics.

**Samples Preparation:** Five Mn-oxides ( $\alpha$ -,  $\beta$ -,  $\gamma$ -,  $\delta$ -MnO<sub>2</sub>,  $\gamma$ -MnOOH) were synthesized by hydrothermal reaction through a Teflon-lined stainless steel autoclave at different temperatures ranging from 90 to 150 °C. The two other Mn-oxides ( $\alpha$ -Mn<sub>2</sub>O<sub>3</sub>,  $\alpha$ -Mn<sub>3</sub>O<sub>4</sub>) were synthesized by solid-state calcination methods using a muffle furnace in the air at 600 °C, 1000 °C, respectively. The initial reagents were MnCl<sub>2</sub>, MnSO<sub>4</sub>, KMnO<sub>4</sub>, (NH<sub>4</sub>)<sub>2</sub>S<sub>2</sub>O<sub>8</sub>, Mn(CH<sub>3</sub>COO)<sub>2</sub>, MnCO<sub>3</sub>, etc. Detailed information on the sample synthesis is listed in Table 1.

**Sample Characterization:** The SEM (Nano SEM

450, USA) and EDS (Aztec Xmax 50, UK) were employed to check the morphology and chemical compositions of Mn-oxides. XRD patterns of Mn-oxides synthesis in the laboratory were collected on an X-ray diffractometer (Rigaku UltimaIV, Japan). The Raman spectra of the synthesized Mn-oxides were collected by inVia Raman System (Renishaw, UK). All the above experiments were conducted at Shandong University at Weihai.

Table 1 Synthesis conditions

synthesis method	objective minerals	reagents	temperature	time duration	dry temperature
hydrothermal reaction	$\alpha$ -MnO <sub>2</sub>	MnSO <sub>4</sub> KMnO <sub>4</sub>	150°C	24h	60°C
	$\beta$ -MnO <sub>2</sub>	MnSO <sub>4</sub> (NH <sub>4</sub> ) <sub>2</sub> S <sub>2</sub> O <sub>8</sub>	140°C	12h	100°C, calcination @ 400°C
	$\gamma$ -MnO <sub>2</sub>	MnCl <sub>2</sub> (NH <sub>4</sub> ) <sub>2</sub> S <sub>2</sub> O <sub>8</sub>	90°C	24h	60°C
	$\delta$ -MnO <sub>2</sub>	MnSO <sub>4</sub> KMnO <sub>4</sub>	140°C	2h	80°C
	$\gamma$ -MnOOH	Mn(CH <sub>3</sub> COO) <sub>2</sub> KMnO <sub>4</sub>	150°C	72h	85°C
solid-state calcination	$\alpha$ -Mn <sub>2</sub> O <sub>3</sub>	MnCO <sub>3</sub>	600°C	5h	
	$\alpha$ -Mn <sub>3</sub> O <sub>4</sub>	MnSO <sub>4</sub>	1000°C	17h	

**Morphology and Chemical Compositions:** The SEM images (Figure 1) show the synthesized Mn-oxides have different morphology. The  $\alpha$ -MnO<sub>2</sub> presents nanofibrous forming an irregular ball with ~8  $\mu$ m in diameter.  $\beta$ -MnO<sub>2</sub> was a nanorod with an average length of ~1.6  $\mu$ m and has good crystallinity.  $\gamma$ -MnO<sub>2</sub> presents with small poorly nanosheets and nanorods mixing in clusters with a diameter range from ~1 to ~12  $\mu$ m.  $\delta$ -MnO<sub>2</sub> has a homogeneous nanosphere with ~400-~600 nm in diameter.  $\alpha$ -Mn<sub>2</sub>O<sub>3</sub> was presented with a nanosheet structure stacked to form a porous nanosphere with ~2  $\mu$ m in diameter. Crystalline  $\alpha$ -Mn<sub>3</sub>O<sub>4</sub> with a hard dumbbell-like nanostructure.  $\gamma$ -MnOOH has presented nanorods with a length-range between ~20 to ~150  $\mu$ m. All the above morphology shows that the  $\alpha$ -Mn<sub>3</sub>O<sub>4</sub>,  $\gamma$ -MnOOH,  $\alpha$ -Mn<sub>2</sub>O<sub>3</sub> have good crystallinity compared with x-MnO<sub>2</sub> (x =  $\alpha$ ,  $\beta$ ,  $\gamma$ ,  $\delta$ ) synthesized at low temperature. EDS was employed to detect the chemistry of Mn-oxides. The results are listed in Table 2. The results reveal that the  $\beta$ -,  $\gamma$ -MnO<sub>2</sub>,  $\alpha$ -Mn<sub>3</sub>O<sub>4</sub>,  $\gamma$ -MnOOH,  $\alpha$ -Mn<sub>2</sub>O<sub>3</sub> are only composed of Mn and O without other elements. However,  $\alpha$ -MnO<sub>2</sub> and  $\delta$ -MnO<sub>2</sub> have K element in their chemical composition, which may be due to their tunnel or layered structures. Previous studies reported that the  $\alpha$ -MnO<sub>2</sub> with (2 $\times$ 2) tunnel structure and the  $\delta$ -MnO<sub>2</sub> with layer structure can accommodate ion (K, Na, Ca, Ba) or water molecule[7].

All the stoichiometric ratio of Mn-oxides are generally in accord with their nominal formula.

**Phase Identification:** The phase information of the Mn-oxides shows in Figure 2. The XRD patterns of those Mn-oxides are consistent with their PDF-2004 mineral database, indicating the purity of synthesized Mn-oxide phases. We calculated the FWHM (Full Width at Half Maxima) of their main peak. The results reveal that the  $\alpha$ - $Mn_3O_4$  (FWHM=0.098°) and  $\gamma$ - $MnOOH$  (FWHM=0.164°) have good crystallinity, while the  $\delta$ - $MnO_2$  (FWHM=1.588°),  $\alpha$ - $MnO_2$  (FWHM=0.825°),  $\gamma$ - $MnO_2$  (FWHM=0.653°) have poor crystallinity. This observation agrees well with morphology measurements by the SEM.

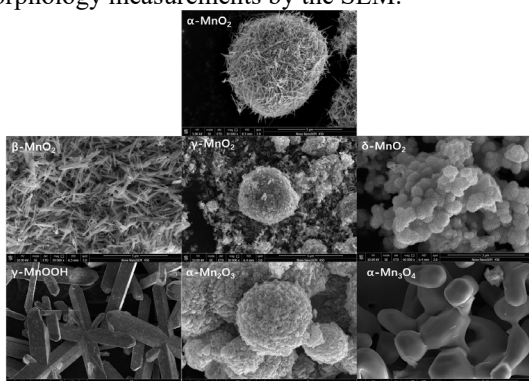


Figure 1 The Morphology of the Mn-oxides

Table 2 The Elements of Mn-oxides detected by EDS

Mn-oxides	Element	Average Atomic(%)
$\alpha$ - $MnO_2$	O	62.05
	K	3.75
	Mn	34.2
$\beta$ - $MnO_2$	O	62.68
	Mn	37.32
$\gamma$ - $MnO_2$	O	62.12
	Mn	37.88
$\delta$ - $MnO_2$	O	55.54
	K	8.03
$\alpha$ - $Mn_2O_3$	Mn	36.43
	O	50.64
$\alpha$ - $Mn_3O_4$	Mn	49.36
	O	42.68
$\gamma$ - $MnOOH$	Mn	57.32
	O	63.18
	Mn	36.82

**Raman Spectra Studies:** As shown in Figure 3, the seven Mn-oxides exhibit distinct individual peaks in their Raman spectra, which can be used as fingerprint spectra for phase identification. Generally, the range of the spectra of Mn-oxides can be divided into two parts,  $\sim 100$ - $400\text{ cm}^{-1}$ ,  $\sim 400$ - $750\text{ cm}^{-1}$ . In the range of  $500$ - $700\text{ cm}^{-1}$ , the Raman spectra can be used to distinguish the  $MnO_2$  from other Mn-oxides such as  $\gamma$ - $MnOOH$ ,  $\alpha$ - $Mn_3O_4$ , and  $\alpha$ - $Mn_2O_3$ . All the Raman spectra of  $MnO_2$  minerals exhibit major double peaks at  $\sim 570$  and  $\sim 650\text{ cm}^{-1}$ .  $\alpha$ - $Mn_2O_3$  shows double peaks at  $\sim 654$  and  $\sim 702\text{ cm}^{-1}$ , while  $\alpha$ - $Mn_3O_4$  has a single peak at  $\sim 660\text{ cm}^{-1}$ .  $\gamma$ - $MnOOH$  has triple peaks at  $\sim 531$ ,  $\sim 558$ ,  $\sim 621\text{ cm}^{-1}$ . All those strong high wavenumber peaks are derived from the stretching mode of  $MnO_6$  octahedra, making them of priority for discriminate those Mn-oxides. There are

also multiple low wavenumber peaks lower than  $\sim 400\text{ cm}^{-1}$ , originating from the bending mode of  $MnO_6$  octahedra.

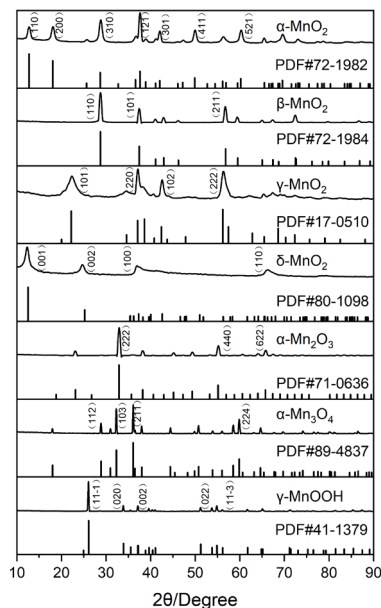


Figure 2 The XRD patterns of the Mn-oxides

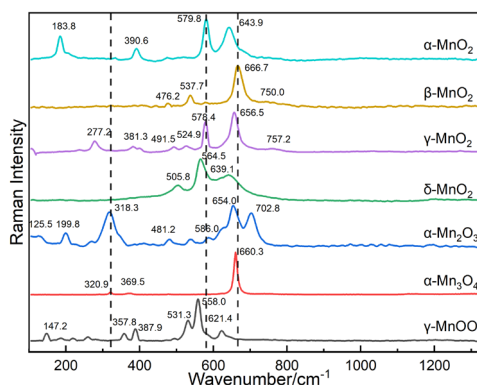


Figure 3 The Raman spectra of Mn-oxides

**Conclusion:** This work provides detailed morphology, structural, spectroscopic studies of seven Mn-Oxides synthesized in the laboratory. Future work will extend to other spectroscopic studies such as MIR and VINR. These experimental characterizations show that the Raman spectroscopy is powerful to straightforwardly identify the seven Mn-oxides. It gives a promising technique to employ the datasets acquired by future Raman instruments selected for Mars missions, i.e., SuperCam and SHERLOC on Perseverance rover and RLS on ExoMars.

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**References:** [1] Turekian K. K. et al. (1961) *Geol. Soc. Am. Bull.*, 72, 175. [2] McSween H. Y. et al. (2004) *Science*, 305, 842. [3] Arvidson R. E. et al. (2016) *Am.*, 101, 1389. [4] Lanza N. L. et al. (2014) *Geophys. Res. Lett.*, 41, 5755. [5] Lanza N. L. (2016) *Geophys. Res. Lett.*, 43, 7398. [6] Liu Y. et al. (2017) *LPSC, LPICo No. 1964*. [7] Post J. E. et al. (2020) *Am.*, 105, 1175. [8] Irvine T. N. et al. (1971) *J. Earth Sci.*, 8, 523-548. [9] Anderson R. et al. (2015) *Icarus*, 249, 2-21.