

WHAT CAN THE ORIENTATIONS OF BENNU'S BOULDERS TELL US ABOUT ITS EVOLUTION?

S.R. Schwartz^{1,2,*}, R.-L. Ballouz^{1,3}, E.I. Asphaug¹, O.S. Barnouin⁴, C. Bennett¹, K.N. Burke¹, Y. Cho⁵, H.C. Connolly⁶, C.D. d'Aubigny¹, D. DellaGiustina¹, M. Hayakawa³, C. Honda⁷, R. Honda⁸, E.R. Jawin⁹, M. Jutzi¹⁰, S. Kameda¹¹, T. Kouyama¹², D.S. Lauretta¹, M. Matsuoka³, P. Michel², H. Miyamoto⁵, J.L. Molaro¹³, T. Morota¹⁴, M. Pajola¹⁵, A.C. Quillen¹⁶, B. Rizk¹, N. Sakatani³, H. Sawada³, D.J. Scheeres¹⁷, S. Sandford¹⁸, S. Sugita⁵, H. Suzuki¹⁹, E. Tatsumi⁵, K.J. Walsh²⁰, M. Yamada²¹, Y. Yokota³, K. Yoshioka⁵, and the Hayabusa2 and OSIRIS-REX Teams.
¹Lunar and Planetary Laboratory, University of Arizona (*1629 University Blvd., Tucson, AZ 85721, srs@lpl.arizona.edu), ²UCA-CNRS-Observatoire de la Côte d'Azur, ³ISAS/JAXA, ⁴The Applied Physics Laboratory, Johns Hopkins University, ⁵University of Tokyo, ⁶Rowan University, ⁷University of Aizu, ⁸Kochi University, ⁹Smithsonian National Institute, ¹⁰University of Bern, ¹¹Rikkyo University, ¹²National Institute of Advanced Industrial Science and Technology, ¹³Planetary Science Institute, ¹⁴Nagoya University, ¹⁵INAF-Astronomical Observatory of Padova, ¹⁶University of Rochester, ¹⁷University of Colorado, Boulder, ¹⁸NASA Ames Research Center, ¹⁹Meiji University, ²⁰Southwest Research Institute, ²¹Chiba Institute of Technology.

Introduction: Boulder populations on Solar System bodies are the net result of the evolutionary processes that have taken place on their surfaces to date. Catastrophic impacts and high-energy events will organize surface boulders more randomly than a series of small-scale events like local micrometeorite impacts, distant larger impacts, thermal quakes, and local avalanching. The power that such small-scale events have to reshape a surface depends on their frequency, seismic efficiency, and how often they are erased by high-energy events. Here we look for boulder orientations as signatures of small-scale events, which may relate to the seismic efficiency and thus the internal structure of the subsurface.

Boulder identification and characterization has been underway since October 2018 and will continue until sample site selection in late 2019. A preliminary analysis of boulder orientations was conducted on PolyCam images (33 cm/px) of asteroid Bennu taken on December 1, 2018, over a region that spans 20° of longitude and about 100° of latitude (around 5% of the surface). This region was chosen to include some specific features of interest and may or may not be representative of the surface as a whole. In this work, boulders are identified with ellipses.

This preliminary dataset shows evidence of a trend for boulders to be oriented with their long ends along the north-south direction (**Fig. 1**, cyan), which corresponds to the global sloping direction, also aligned north-south [1]. Further, this figure shows that if we weight the “value” of each boulder by its elongation, the case for this preferential boulder orientation becomes stronger (**Fig. 1**, purple). We compare this result to the global north-south distribution of boulder orientations. In addition, using the local dynamical slopes from the OSIRIS-REx Radio Science Working Group based upon shape models from the Altimetry Working Group [2-3], we show how boulders align themselves in relation to the local dynamic slope [4-6].

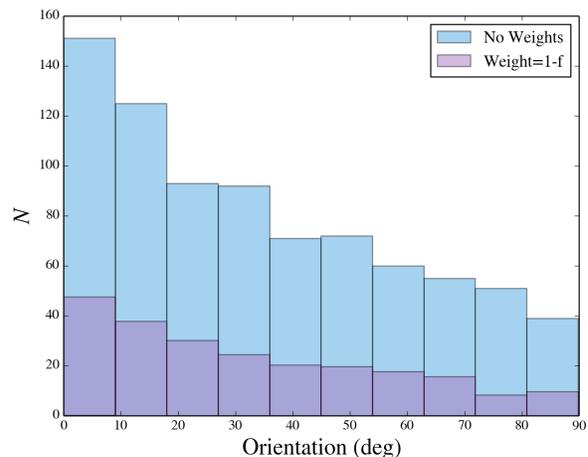


Figure 1: Distribution of boulder orientations relative to their local lines of longitude (cyan). Weighting the boulders by their elongation ($f = \text{short axis}/\text{long axis}$) is also shown (purple).

References: [1] Schwartz, S.R. et al. (2018), *AGU P21A-11*, [2] Barnouin O.S. et al., *AGU P33C-3835*, [3] Scheeres, D.J. et al. (2018), *AGU P22A-05*, [4] Schwartz S.R. et al. (2012), *Granular Matter* 14, 363. [5] Stadel J.G. (2001), *Ph.D. thesis, U. Washington*. [6] Richardson D.C. et al. (2000), *Icarus* 143, 45.

Acknowledgements: This material is based upon work supported by NASA under Contract NNM10AA11C issued through the New Frontiers Program and from Grant no. 80NSSC18K0226 as part of the OSIRIS-REx Participating Scientist Program. S.R.S. and P.M. acknowledge support from the Complex Systems and Space, Environment, Risk and Resilience Academies of the Initiative d'EXcellence “Joint, Excellent, and Dynamic Initiative” (IDEX JEDI) of the Université Côte d'Azur. M.P. was supported by the Italian Space Agency (ASI) under agreement no. 2017-37-H.0. P.M. acknowledges support from the French space agency (CNES). Additional support provided by JSPS International Planetary Network.