

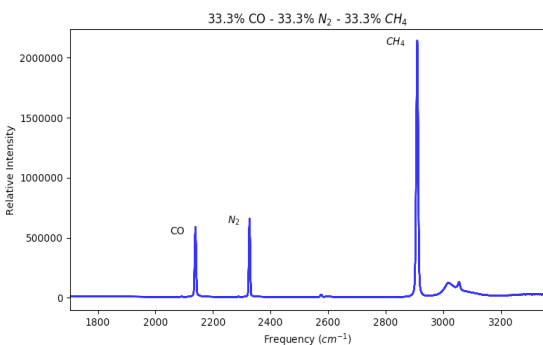
## RAMAN SHOP: EXPLORING THE CARBON MONOXIDE-NITROGEN-METHANE TERNARY SYSTEM OF SPUTNIK PLANITIA VIA RAMAN SPECTROSCOPY.

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**Introduction:** On 14 July 2015, New Horizons completed its primary mission when it flew by Pluto at a distance of 12,000 km from the surface [1]. Data sent back to Earth presented a planetary surface not wholly dead; Sputnik Planitia, the left lobe of Tombaugh Regio, displays ongoing convective processes with an average turnover rate of about 500,000 years [2]. Most likely formed by a >150 km ancestral Kuiper Belt impactor, this oval-shaped structural depression is measured to be approximately 1300 km x 900 km in size [3]. The impact basin covers roughly 5% of Pluto's surface and is characterized by polygonal convective cells surrounded by geologically diverse terrain [4]. This area of convection is composed predominantly of N<sub>2</sub> ice with smaller quantities of CO and CH<sub>4</sub> ices.

It is this N<sub>2</sub>-rich ice that we are most interested in describing by means of Raman spectroscopy. Pure N<sub>2</sub> freezes at 63.15 K and shifts from its hexagonal  $\beta$ -phase to a cubic  $\alpha$ -phase at 35 K [5]. CO is fully miscible, in N<sub>2</sub> and encounters its phase transition at 61 K. The addition of CO to N<sub>2</sub> increases the temperature at which the  $\alpha$ - $\beta$  transition occurs, making it possible for the process to take place on and below Pluto's ~40 K surface. That being said, it is of particular interest to find how the CO-N<sub>2</sub> system is perturbed when CH<sub>4</sub> is introduced. The exploration of this ternary system may provide insight into whether the occurrence of the  $\alpha$ - $\beta$  transition is plausible within the Sputnik Planitia region.

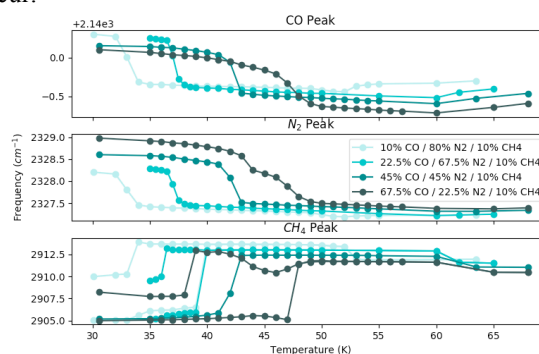
**Analysis:** Raman scattering provides a means of detecting phase transitions by displaying a downshift in frequency when the transition occurs. Each of the three components used in the mixture exhibit a prominent peak in a given region of the spectrum, as shown in figure 1.



**Figure 1: Spectrum of equal mixing at 45 K.** Each prominent peak is described with the corresponding molecule. As the temperature is decreased the peak shifts downward.

By using a Gaussian fit we are better able to track the movements of the peaks as they shift in wave-number. In the case of the CH<sub>4</sub> peak, we implemented a double Gaussian fit as it displays a double peak feature as the temperature is lowered. This is caused by the solubility limit, and indicates that CH<sub>4</sub> is crystallizing out of the liquid. When the CH<sub>4</sub> is dissolved in the mixture, it peaks predominantly near 2912 cm<sup>-1</sup>. As the solubility limit is reached, a second peak forms at 2905 cm<sup>-1</sup>, which is the region of the spectrum where pure CH<sub>4</sub>'s most prominent peak appears.

Figure 2 shows how the peaks shift in frequency as a function of temperature. The double peak feature is displayed in the temperature-frequency diagram as a split in the max peaks where the feature begins to occur.

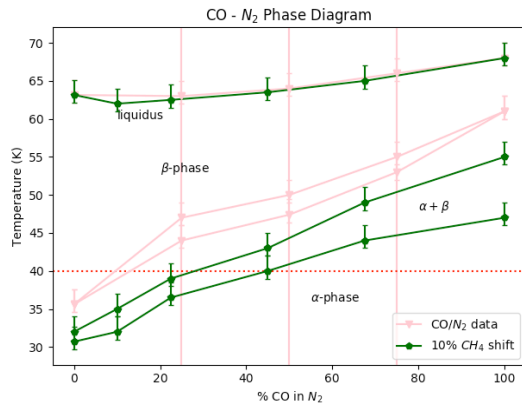


**Figure 2: Temperature-Frequency Diagram of CO-N<sub>2</sub>-CH<sub>4</sub>.** Each point represents a measurement taken at a specific temperature and the resulting shift in frequency.

While the N<sub>2</sub> and CO peaks mostly agree with one another, the CH<sub>4</sub> peak tells a different story. CH<sub>4</sub> has a strong Raman feature, and quickly overshadows the peaks formed by N<sub>2</sub> and CO, even when it is only 10% of the overall solution. This indicates that it may have the ability to display more subtle changes occurring in the system.

**Results:** A multitude of spectra were collected first as a binary sequence of CO-N<sub>2</sub> and then as a ternary system by adding 10% CH<sub>4</sub> to the CO-N<sub>2</sub> mixtures. Previous research has shown that the combination of CO and N<sub>2</sub> raises the temperature at which the solid-solid  $\alpha$ - $\beta$  phase transition occurs in N<sub>2</sub> alone. The elevation in temperature implies that a phase transition should be apparent in the remotely sensed New Horizons data, and yet, it is not. Figure 3 demonstrates that inserting CH<sub>4</sub> into the system depresses the temperature at which the transition from  $\beta$  to  $\alpha$  occurs, indicating that the convective cells of Sputnik Planitia may

not actually be at appropriate temperatures for solid phase transitions to take place.



**Figure 3: Phase diagram of the apparent shift caused by the addition of  $\text{CH}_4$  into the  $\text{CO-N}_2$  system.** Measurements indicate that the introduction of  $\text{CH}_4$  into the system causes a depression of when the solid-solid transition occurs. The red dotted line represents the average temperature of Pluto's surface.

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**References:** [1] Protopapa, S., *et al.* (2017) *Icarus*, 287, 218–228. [2] McKinnon, W.B., *et al.* (2016) *Nature*, 534, 82–85. [3] McKinnon, W.B., *et al.* (2017) *LPS XLVIII*, Abstract #2854. [4] Trowbridge, A.J., *et al.* (2016) *Nature*, 534, 79–81. [5] Angwin, M.J. & Wasserman, J. (1966) *The Journal of Chemical Physics*, 44, 417.