

**Monday, March 20, 2017**  
**CERES: FEATURES, COMPOSITION, AND EVOLUTION**  
**8:30 a.m. Montgomery Ballroom**

[M105]

**Chairs: Debra Buczkowski**  
**Ernesto Palomba**

- 8:30 a.m. Raymond C. A. \* Castillo-Rogez J. C. Ermakov A. Park R. S. Marchi S. et al.  
[Large-Scale Heterogeneity of Ceres: Clues to Interior Evolution](#) [#1506]  
 Ceres is nearly hydrostatic with compensated topography. Correlation of residual gravity anomalies with other data provide clues to its evolution.
- 8:45 a.m. Buczkowski D. L. \* Sizemore H. G. Jozwiak L. M. Schenk P. M. Scully J. E. C. et al.  
[Floor-Fractured Craters on Ceres: Implications for Internal Composition and Processes](#) [#2117]  
 We present a geomorphic and topographic analysis of the Ceres floor fractured craters and propose hypotheses for their formation.
- 9:00 a.m. Thangjam G. \* Nathues A. Platz T. Mengel K. Cloutis E. A. et al.  
[Understanding Cere's Compositional Heterogeneity from Bright and Dark Materials](#) [#2235]  
 The bright and dark materials on dwarf planet Ceres is intended to study using both FC and VIR data to shed light into the surface compositional heterogeneity.
- 9:15 a.m. De Sanctis M. C. \* Ammannito E. McSween H. J. Marchi S. Raponi A. et al.  
[Discovery of Organics Rich Terrains on Ceres](#) [#1493]  
 Here we report about the discovery of organic rich areas on Ceres by the Dawn VIR instrument. The spectra show a strong 3.2–3.5  $\mu\text{m}$  band.
- 9:30 a.m. Pieters C. M. \* Nathues A. Thangjam G. Hoffman H. De Sanctis C. et al.  
[Context of Unusual Red Organic-Rich Areas on Ceres and Geologic Constraints for Their Origin](#) [#1296]  
 Organic material recently identified on Ceres is found to be associated with a family of small fresh craters, providing an important constraint for its origin.
- 9:45 a.m. Jaumann R. \* Presuker F. Krohn K. von der Gathen I. Stephan K. et al.  
[Topography and Geomorphology of the Interior of Occator Crater on Ceres](#) [#1440]  
 The crater floor of Occator on Ceres contains plain material is about 500 m thick and due to impact melt hydrothermal alteration or cryovolcanic crater filling.
- 10:00 a.m. Stein N. \* Ehlmann B. Ammannito E. Palomba E. De Sanctis M. C. et al.  
[Characteristics, Formation, and Evolution of Faculae \(Bright Spots\) on Ceres](#) [#2547]  
 Faculae (bright spots) on Ceres are comprehensively mapped and categorized. Potential mechanisms for their formation and evolution are evaluated.
- 10:15 a.m. Palomba E. \* Longobardo A. De Sanctis M. C. Stein N. T. Ehlmann B. et al.  
[Bright Spots on Ceres: Occator, Oxo, and the Others](#) [#1566]  
 By using the VIR data, a catalogue of 98 Ceres Bright Spots was obtained and analyzed to reveal if they have a common or a different origin and evolution.
- 10:30 a.m. Sizemore H. G. \* Platz T. Schorghofer N. Crown D. A. Prettyman T. H. et al.  
[Ceres' Pitted Terrains: Morphological Context and Implications for Ground Ice](#) [#2033]  
 Ceres is the third terrestrial body known to host pitted crater materials. We discuss cerean pit morphology, development, and implications for ground ice.

- 10:45 a.m. Hughson K. H. G. \* Russell C. T. Schmidt B. E. Chilton H. Combe J.-P. et al.  
[Layered and Low-Aspect-Ratio Ejecta on Ceres: Probing the Effect of Ground Ice on Fluidized Ejecta Deposits](#) [#1609]  
Ceres supports a multitude of flows similar to martian fluidized ejecta. We analyze and model these flows to assess the effect of ground ice on their genesis.
- 11:00 a.m. Platz T. \* Nathues A. Schorghofer N. Preusker F. Mazarico E. et al.  
[Surface Water-Ice Deposits in the Northern Shadowed Regions of Ceres](#) [#2447]  
This study describes how permanent shadow maps are generated and how water-ice deposits were detected (including processes leading to water-ice deposition).
- 11:15 a.m. Prettyman T. H. \* Yamashita N. Toplis M. J. McSween H. Y. Castillo-Rogez J. et al.  
[Elemental Measurements of Ceres by Dawn: The Search for Nickel](#) [#1677]  
Elemental mapping data acquired by Dawn's Gamma Ray and Neutron Detector provide new insights into the hydrothermal evolution of Ceres.
- 11:30 a.m. Lawrence D. J. \* Peplowski P. N. Feldman W. C. Prettyman T. H. Russell C. T. et al.  
[High-Energy Gamma Rays from Dawn's Gamma Ray and Neutron Detector: A Measure of Ceres' Hydrogen Concentration](#) [#2098]  
High-energy gamma rays measured by Dawn/GRaND provide a measurement of Ceres' hydrogen concentration, and may also be sensitive to Fe variations.
- 11:45 a.m. Castillo-Rogez J. C. \* Raymond C. A. Prettyman T. H. McSween H. Y. Ruesch O. et al.  
[Evolution of Large Volatile-Rich Bodies: New Insights from Ceres](#) [#2172]  
We study the feedbacks between chemistry and geophysics at Ceres observed by Dawn. This knowledge is transferred to other large volatile-rich bodies.