

NORTHWEST AFRICA 8535: UNIQUE DUNITIC ANGRITE. C.B. Agee, H.M. Miley, K. Ziegler, and M.N. Spilde. Institute of Meteoritics and Department of Earth and Planetary Sciences, University of New Mexico, Albuquerque, NM 87131.

Introduction: We report here the discovery of a new type of angrite, NWA 8535, a dunite with approximately 90-95% zoned olivine, which now expands the range of known lithologies for the angrite parent body beyond the current bounds of “volcanic” and “plutonic” types.



Fig. 1. Hand sample, the NWA 8535 main mass, broken surface reveals green olivine crystals and prominent cavity (dark shadow).

Physical characteristics: NWA 8535 is a 149-gram single stone purchased in Morocco in March 2014. The exterior of the sample is slightly weathered retaining patches of black fusion crust, coarse green olivine crystals up to several millimeters are visible on the surface as well as some smoothed mm-sized vugs. The broken surface that yielded the deposit sample revealed a fresh, bright green, polycrystalline texture with a prominent cavity -- presumably a vug or large vesicle (figure 1).

Mineralogy & Petrology: Electron microprobe examination of several polished mounts shows approximately 90-95% olivine, ~3% fassaite, ubiquitous aluminous spinel, chromium spinel, troilite, kamacite, taenite, and trace amounts of kirschsteinite -- no plagioclase was detected. Olivines are all zoned with Fe-enriched rims, many of the grains probed are in the size range of 500-1000 μm , however in hand sample

several larger single crystals up to ~1 cm were observed. The most magnesian olivine core is Fa10.2, the most Fe-rich rim measured is Fa38.8, with the mean value $\text{Fa}20.6 \pm 5.8$ ($n=148$). Because olivine dominates the mineralogy of NWA 8535, the bulk composition is

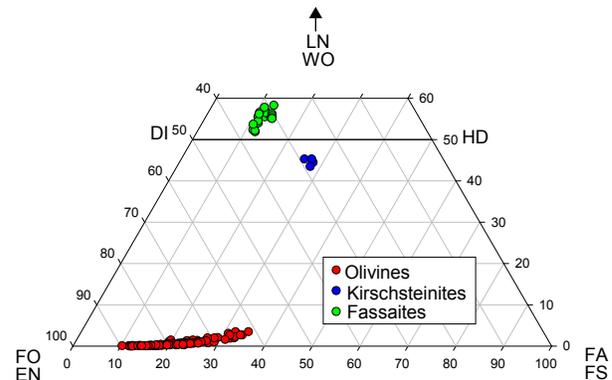


Fig 2. Combined olivine and pyroxene quadrilateral showing the olivine, kirschsteinite, and fassaite compositions in NWA 8535.

by far the most magnesian and refractory of the angrites. Olivines have a wide range of CaO contents ranging from 0.02-2.49 wt%, positively correlating with Fa-content (figure 2), and also have relatively

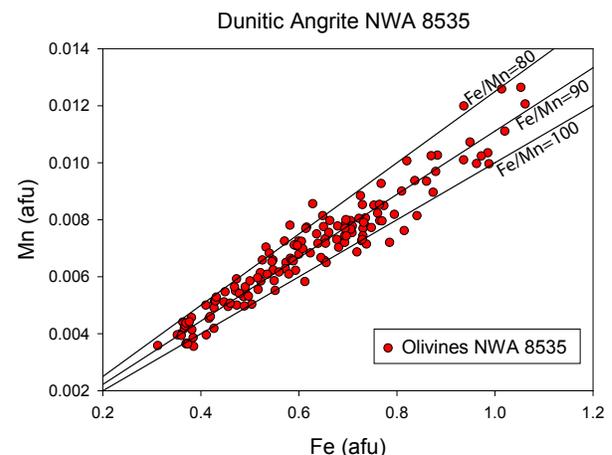


Fig 3. Fe versus Mn showing the range of values for olivines in NWA 8535.

high $\text{Cr}_2\text{O}_3=0.26 \pm 0.18$ wt%. Olivines in NWA 8535 have typical angritic values of $\text{Fe}/\text{Mn}=90 \pm 7$, which was an early diagnostic determined during the reconnaissance and classification phase of this meteorite (figure 3). Fassaite (Al-Ti-bearing diopside), another

common angritic mineral occurs throughout NWA 8535, primarily as small grains (~50-100 μm) along olivine grain boundaries and occupying triple junctions. The mean values of fassaite are $\text{Fs}_{12.4\pm 1.7}\text{Wo}_{55.3\pm 1.8}$, $\text{Fe}/\text{Mn}=92\pm 24$, $\text{Al}_2\text{O}_3=12.9$ wt%, $\text{TiO}_2=1.9$ wt%, $n=20$ (figure 2).

Oxygen Isotopes: NWA 8535 is classified as an angrite meteorite based on oxygen isotopes, Fe/Mn of olivine, and presence of fassaite and kirschsteinite. Figure 4 shows our results for laser fluorination analyses of seven acid-washed portions of bulk sample with values of $\Delta^{17}\text{O} = -0.070, -0.138, -0.109, -0.053, -0.069, -0.085, -0.105$ (all linearized, permil). All of the measured values fall within the $\Delta^{17}\text{O}$ range for angrite meteorites, however one sample gave a $\delta^{18}\text{O}$ value that extends the lower range of angrites by nearly 1%.

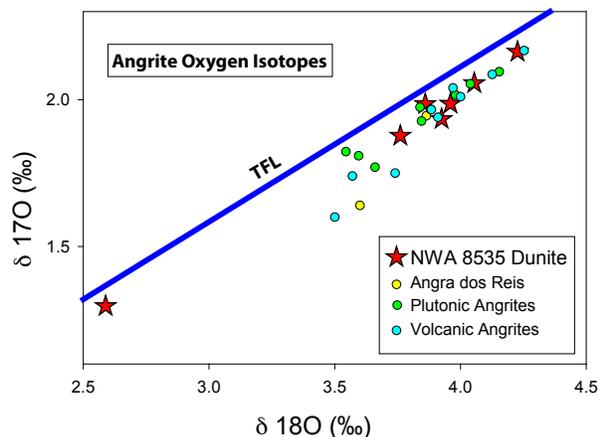


Fig 4. Triple oxygen isotope diagram showing data from NWA 8535 (red stars) compared to other angrite meteorites [1]. TFL=terrestrial fractionation line.

Discussion: NWA 8535 expands the range of known lithologies from the angrite parent body, however there remain numerous questions about its origin, formation and genetic relationship to the other angrites. Although NWA 8535 is an ultramafic rock (dunite), it bears little textural resemblance to the so-called plutonic angrites, in that its olivines are strongly zoned (figure 5), suggesting rapid cooling during formation. Therefore, it is possible that NWA 8535 has a more direct link to the so-called volcanic angrites, most of which appear to be rapidly cooled basalts. Interestingly, the proposed “xenocrystic” olivines in volcanic angrites Asuka 881371 and D’Orbigny [1] overlap in composition with the olivines in NWA 8535, however further work is needed to establish any genetic relationship to these meteorites. Future work will also include age-dating, trace element and isotopic studies, magnetics, and the nature of the vugs and numerous vesicle-like inclusions in NWA 8535.

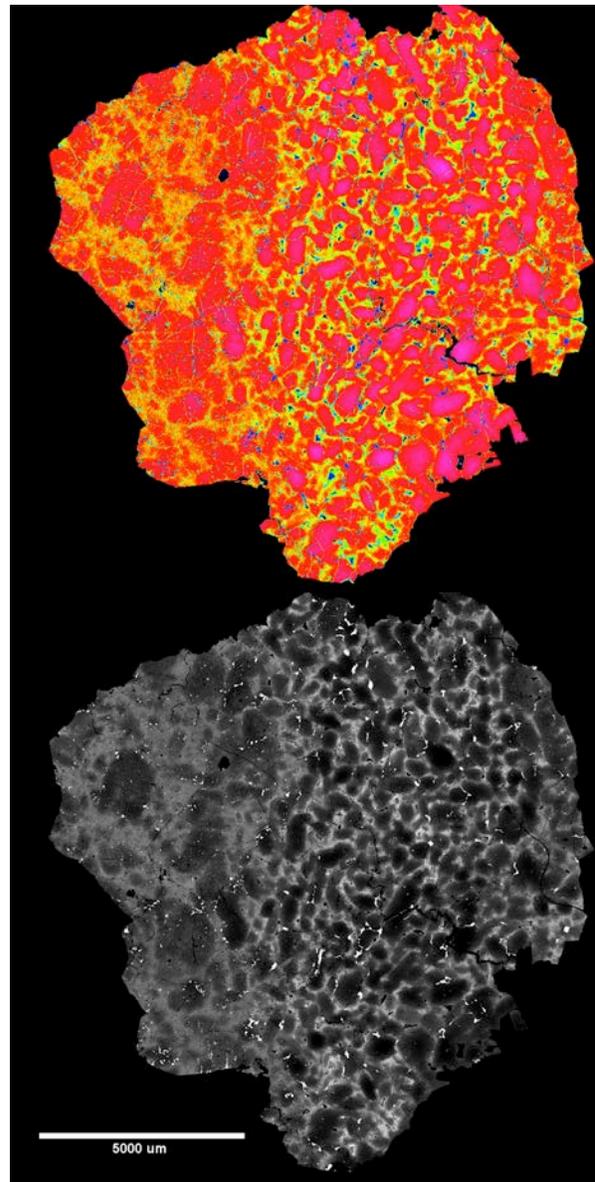


Fig 5. An example of a NWA 8535 deposit sample probe mount used in this study. Top image is a false color magnesium map, bottom is the same image in BSE, scale bar 5000 microns.

References: [1] Keil K. (2012) *Chemie der Erde*, 72, 191–218 (and referenced conference abstracts therein).