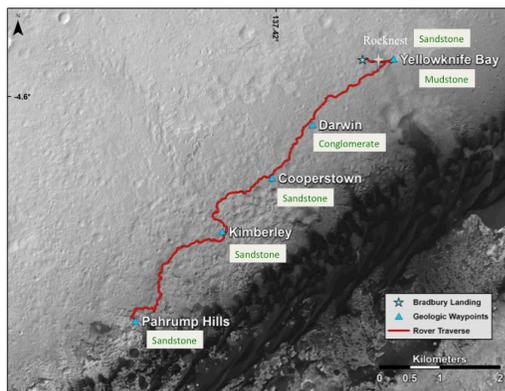


**THE DIVERSITY OF SEDIMENTS AT GALE CRATER FROM CHEMCAM OBSERVATIONS: EVIDENCE FOR MULTIPLE SEDIMENT SOURCE CHEMISTRIES, DIVERSE ALTERATION HISTORIES, AND MULTIPLE DIAGENETIC EPISODES.** D. L. Blaney<sup>1</sup> Roger Wiens<sup>2</sup>, Sylvestre Maurice<sup>3</sup>, Ryan Anderson<sup>4</sup>, John Bridges<sup>5</sup>, Sam Clegg<sup>2</sup>, Laetitia Le Deit<sup>6</sup>, Martin Fisk<sup>7</sup> Olivier Forni<sup>3</sup>, Olivier Gasnault<sup>3</sup>, Linda Kah<sup>8</sup>, Nina Lanza<sup>2</sup>, Jeremie Lasue<sup>3</sup>, Nicholas Mangold<sup>6</sup>, Marion Nachon<sup>6</sup>, Horton Newsom<sup>9</sup>, Agnes Piller<sup>2</sup>, Violaine Sautter<sup>10</sup>, and the MSL Science Team.<sup>1</sup>NASA Jet Propulsion Laboratory, California Institute of Technology, 4800 Oak Grove Drive, MS 264-527, Pasadena, CA 91109, <sup>2</sup>Los Alamos National Laboratory, Los Alamos, NM 87544 USA, <sup>3</sup> Université Paul Sabatier; UPS-OMP; Institute de Recherche en Astrophysique et Planetologie (IRAP), F-31400 Toulouse, France. CNRS; IRAP; 9 Av. Colonel Roche, BP 44346, F-31028 Toulouse cedex 4, France, <sup>4</sup>U.S. Geological Survey, Astrogeology Science Center, Flagstaff, AZ 86001-1698, USA, <sup>5</sup> Space Research Centre, Dept. of Physics & Astronomy, University of Leicester, LE1 7RH, UK, <sup>6</sup>LPG Nantes, CNRS, UMR6112, Université Nantes, Nantes, France, <sup>7</sup>College of Earth, Ocean, and Atmospheric Sciences, Oregon State University, Corvallis, OR 97331, <sup>8</sup>University of Tennessee, Knoxville, TN 37996, SA, <sup>9</sup>Institute of Meteoritics, University of New Mexico, Albuquerque, NM 87106, USA, <sup>10</sup>Museum National d'Histoire Naturelle, Laboratoire de Minéralogie et Cosmochimie du Muséum, France.

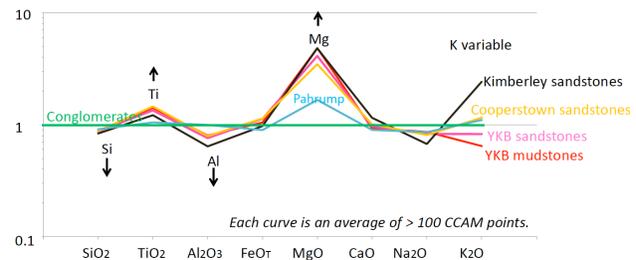
**Summary:** Since landing in Gale Crater the Curiosity Rover has been investigating sedimentary materials using ChemCam. To date >5,400 separate chemical rock and soil measurements have been made, which enable us to investigate the range of sediment chemistry along the traverse (Figure 1). Data reveals that the Gale sediments have multiple initial starting chemistries, and have experienced different alteration histories, including multiple episodes of diagenesis.



**Figure 1: Curiosity traverse showing key locations of ChemCam measurements.**

**Data:** ChemCam is a Laser Induced Breakdown Spectrometer (LIBS) with an integrated Remote Microscopic Imager (RMI) to provide context of where each LIBS spectra is collected. ChemCam LIBS works by firing a laser focused to a 350-550  $\mu\text{m}$  diameter spot that produces plasma from the rock. Spectra of elemental emission lines are recorded from 240-850 nm and used to determine the elemental composition of the rock [1,2, for more details on ChemCam and data analysis]. Chemical compositions were generated from individual spectra using the ChemCam team standard Partial Least Squares 1 analysis to produce elemental oxide abundance for  $\text{SiO}_2$ ,  $\text{Fe}_2\text{O}_3$ ,  $\text{CaO}$ ,  $\text{MgO}$ ,  $\text{Al}_2\text{O}_3$ ,  $\text{TiO}_2$ ,  $\text{Na}_2\text{O}$ , and  $\text{K}_2\text{O}$  [e.g. 3,4, 5]. These

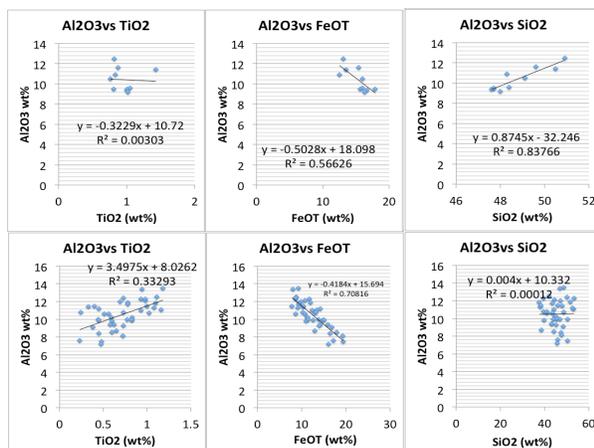
data can be used to examine the compositional make-up of lithified outcrops. Elemental variation within individual targets can then be used to infer mineralogy by looking at elemental correlations.



**Figure 2. Key sedimentary unit compositions ratioed over conglomerate mean chemistry showing differences in sediment chemistry.**

**Sediment Source Chemistry:** Sediments and sedimentary rocks within Gale crater have three primary chemical source components (Figure 2). First is a basaltic component, which dominates the composition at locations such as Yellowknife Bay and Rocknest [6,7]. Second is a distinctly plagioclase feldspar-enriched composition, which is a contributor to coarser-grained facies at Darwin [8] and at Pahrump [9]. Third is a potassium feldspar-bearing component that is present in the Cooperstown and Kimberly outcrops [10]. Basalt-dominated compositions are present in both mudstones at Yellow Knife Bay [6,11] and in sandstones [7, 12]. The differences in chemistry may be the result of: 1) different source regions for the sediments either originating from different physical locations or sampling deeper units in the crust as deposition continued; 2) differences in transport / sorting during deposition; or 3) a combination a source region variability and transport. However since the sandstones have indications of all three-source components the differences in chemistry is not primarily due to grain size effects and some change in the source chemistry is likely.

**Degree of Alteration:** The sediments at Gale range from having strong preservation of primary igneous components, to having substantial contribution from secondary alteration minerals. CheMin observations at Yellowknife Bay show large amounts of amorphous material and smectite clays [13] while the Kimberly outcrop is much less altered [14]. At locations where there is no CheMin mineralogical data, ChemCam can be used to assess the amount of alteration via a combination of sedimentary texture and elemental patterns that can be used to infer mineralogy. In conglomerates such as those at Darwin, pebbles frequently have igneous textures [13] and ChemCam patterns support a plagioclase phase being present [8]. At Rocknest, Fe and Mg patterns suggest both the presence of iron oxides and depletion of primary olivine materials [7]. These data indicate that, while likely not as altered as the Yellowknife Bay mudstones, some chemical alteration had taken place in the formation. At Pahrump, different elemental patterns are seen in the  $\text{Al}_2\text{O}_3$ - $\text{TiO}_2$ - $\text{FeOT}$ - $\text{SiO}_2$  (Figure 3) in the adjacent Book Cliff and Pink Cliff regions suggesting that although the overall chemistry of the outcrop is similar for both facies [7,14], the Pink Cliff (lower) part of the outcrop is much more altered.



**Figure 3. Elemental relationships between  $\text{Al}_2\text{O}_3$  and  $\text{TiO}_2$ ,  $\text{FeOT}$ , and  $\text{SiO}_2$  in the Book Cliffs Region (top, target Goblin Valley) and in Pink Cliffs region (bottom, targets Delta (no vein), San Rafael Swell (no rim around crack), Castle Valley, Ibex Pass (no vein), Saddle Peak). The correlation of  $\text{Al}_2\text{O}_3$  with  $\text{TiO}_2$  and lack of correlation with  $\text{SiO}_2$  suggest that altered mineral phases are present at Pink Cliffs. This contrasts with the Book Cliffs region observation where positive  $\text{Al}_2\text{O}_3$  with  $\text{SiO}_2$  suggest plagioclase feldspar is present.**

**Diagenetic Features:** There is also evidence for widespread and variable diagenetic processes. Post deposition aqueous activity at Gale continued to be widespread with variable chemistry including fluids rich in calcium and magnesium sulfate, Mn-oxides, and fluorine.

**Calcium Sulfate** Calcium sulfate veins have been observed in all of the sandstones / mudstones observed by ChemCam at Gale [e.g. 6, 7, 8, 9, 10, 15]. These veins fill fractures that cross cut the sandstones and are likely the most recent aqueous event in the area.

**Mg Loss and Mg Enhancements:** Raised Mg-rich ridges are seen at Yellowknife Bay and Pahrump by ChemCam [16] and at Pahrump APXS determined that they were also enriched in S, N, Cl, and Br suggesting that they were caused by the emplacement of  $\text{MgSO}_4$  rich fluids [17]. Sediments depleted in MgO have also been observed, most notably at Rocknest [7] suggesting that diagenetic alteration of olivine may be a source for the Mg [7, 16].

**MnO Coatings/Resistant Ridges:** At Kimberly, MnO coatings and resistant ridges indicate the presence of a highly oxidizing fluid [18]. Enhanced MnO is spatially restricted and is one of the most localized diagenetic features.

**CaF:** CaF has been detected in various locations, including Pahrump concentrated in resistant ridges, indicating a fluvial origin for the phase [9,19].

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