GENERATION OF ULTRAHIGH SPATIAL RESOLUTION DIGITAL TERRAIN MODELS FOR A MARTIAN LAVA FLOW ANALOG FROM KILAUEA VOLCANO, HAWAII. S. P. Scheidt[1] and C. W. Hamilton, [1]Lunar and Planetary Laboratory, University of Arizona, 1629 E University Blvd, Tucson, AZ 85721 (scheidt@lpl.arizona.edu).

Introduction: Sinuous channels are common on the surface of many planets and moons, but on the Earth and Mars, these channels may either have been formed by fluvial or volcanic processes. Distinguishing between these formation mechanisms is therefore critical for correctly interpreting their origin and geological significance. The December 31, 1974, lava flow from Kīlauea Caldera, Hawai`i, exhibits fluvial-like morphologies, and was selected as a locale for field campaigns as a terrestrial analog for volcanic sinuous channels on Mars, such as those located within the Cerberus Fossae 2 unit in Elysium Planitia. The sheet-like lava flow [1] was rapidly emplaced [2, 3, 4] from a series of en echelon fissures southwest of Kīlauea Caldera. The resulting flow thickness is ~1-m-thick in the proximal regions and transition to much thicker channelized `a`ā flows ~2 km from the vent [1].

Topographic data of the flow were needed to test the hypothesis that sinuous channels can form by lava flow emplacement processes involving an initial constructional stage, followed by lava drainage; and to make direct comparisons with morphological and topographic features observed in Mars Reconnaissance Obiter (MRO) High Resolution Imaging Science Experiment (HiRISE) imagery and stereo-derived Digital Terrain Models (DTMs). Here, we describe preliminary results obtained using Multi-View Stereo-Photogrammetry (MVSP) techniques applied to aerial photographs acquired from a kite-based platform.

Methods: Digital phototographs were aquired during the Summer of 2014 for a 0.5 km^2 region of the December 1974 flow using low-altitude, Kite Aerial Photography (KAP) methods in the Hawai`i Volcanoes National Park (supported by NASA Planetary Geology and Geophysics grant NNX13AQ05G, under the authorization of National Park Service permit #HAVO-2012-SCI-0025). The KAP technique allowed the capture of high-resolution images at low cost from a unique aerial vantage point. These images where then used to produce contiguous orthographic images and topographic data for the lava flow surface. Incidence angles (θ) were 0° ± 10° for images collected using a Canon Powershot S110 loaded with a firmware script to capture 12 megapixel images automatically every 3 seconds in flight from a customized KAP platform. A second camera (Pentax Optio-WG2 with geotagging) was mounted on a platform with an automatically rotating servomotor and infrared shutter release to acquire off-nadir images (9 megapixels) every 5 seconds (θ < 20° ± 10°). KAP rigs were attached to the kite line approximately 10 m below the kite using a Picavet suspension system. Coverage of the area was accomplished by hiking across the lava flow during eight separate flights. The images were then used as input for a custom MVSP processing workflow based on [5, 6]. This software simultaneously calculates the aerial positions of the camera for each image, refines interior and exterior camera orientation parameters, and generates a 3-Dimensional (3D) point cloud of the surface (Fig. 1). In general, feature detection and matching of multiple overlapping images results in a unique solution of 3D tie points. A dense reconstruction of the scene is then processed from the sparse tie points [e.g., 6] and post-processing (e.g., statistical outlier removal and smoothing) are applied to create color-textured, georeferenced orthoimages and DTMs. Black and white ground targets (1.3 × 1.3 m) were placed throughout the lava flow surface, geolocated with differential GPS, and used for geolocation of the 3D data.

Results: The preliminary 3D model covers approximately 5×10^6 m^2 and contains 5×10^8 points, resulting in a grid resolution of < 2 cm.

Figure. 1. The image shows two rendered 3D point clouds seen from an oblique view. The upper cloud shows positions of 11,649 aerial images captured by a camera lofted by a kite. The lower portion represent a half billion topographic points calculated using MVSP of the December 1974 flow on Kīlauea Volcano, Hawai`i.
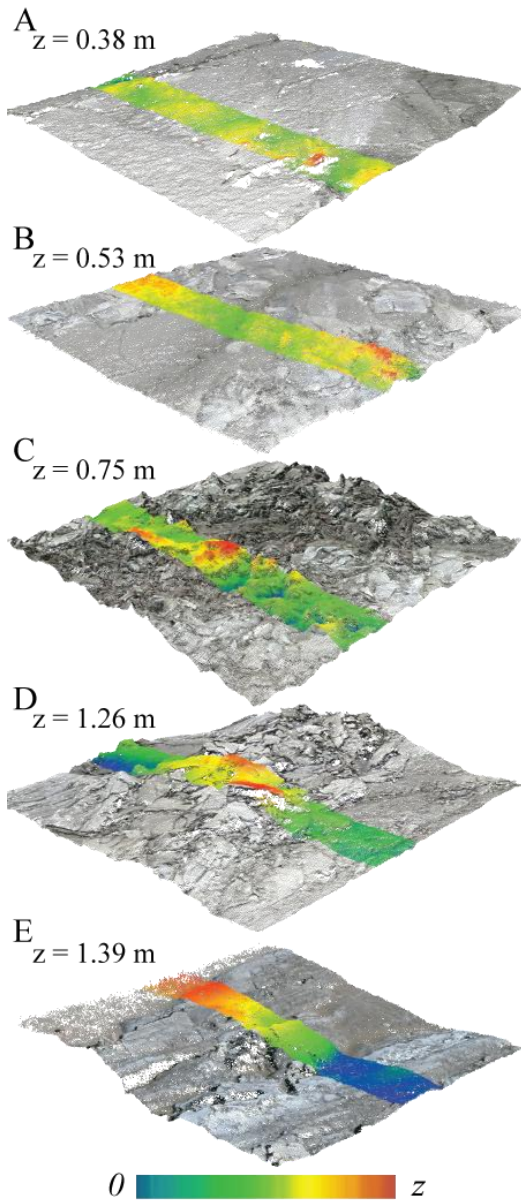
Figure 2. Rendered 3D point clouds (10 × 10 m area, 3 mm resolution) exhibit varied flow surface textures. The color overlay illustrates the maximum elevation difference within the colored region (z) for five transitional surface textures: (A) smooth platy pāhoehoe, (B) a lava spreading zone, (C) rubbly pāhoehoe, (D) slabby pāhoehoe, and (E) and overbank flows.

Highly oblique images (θ ≈ 60°) provide a synoptic view of the flow (Fig. 3A). Dark- and light-albedo areas highlight rafted lava slabs and plates, sinuous channels, and streamlined islands. Islands with a tan color are high-standing outcrops (kīpuka) composed of older lava flows, which contrast starkly with the silver-gray and black December 1974 lava [8]. Vertices in the 3D point cloud are rendered with the average color of matched image pixels (Fig. 3B). Figure 3 also shows other features of interest, such as older lava units exposed along the northern flow margin, and a prominent fault scarp of Kīlauea's Southwest Rift Zone, which confines the southern margin of the December 1974 flow. Data are sparse along the edge of the point cloud and in low-texture areas (e.g., sand) where the MVSP algorithms could extract little to no information.

Subsequent data processing will generate orthoimages and merge the dense 3D point clouds with ground-based photometric data and Light Detection and Ranging (LiDAR) measurements to enable detailed facies mapping and terrain analysis in ArcGIS.
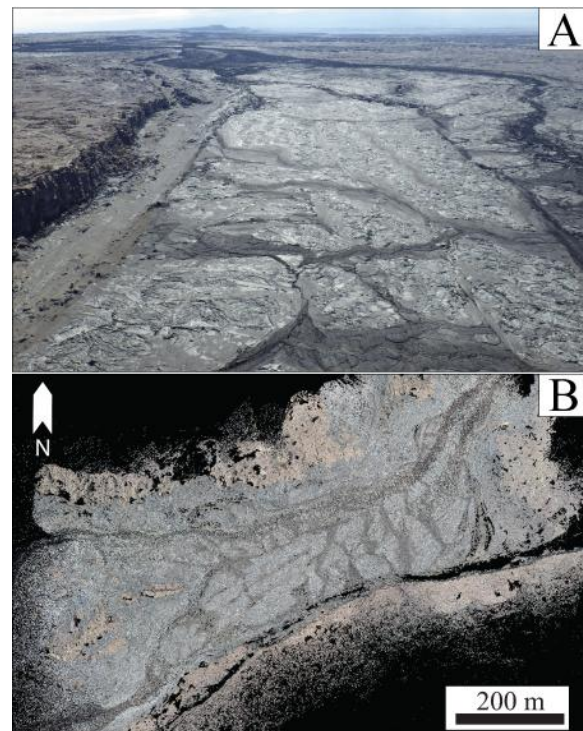

Figure 3. (A) Oblique aerial image of the study area within the December 1974 flow, facing south-southwest.(B) 3D point cloud viewed from nadir.

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