

Ground-based Magnetoseismology Observations for the Next Decade

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Solar and Space Physics Decadal Survey White Papers
Workshop #3:

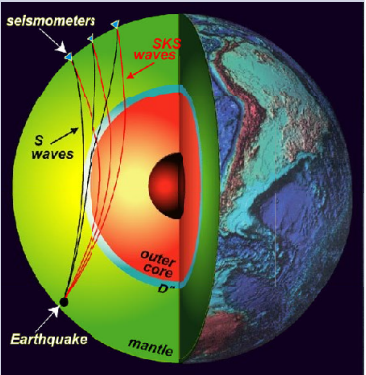
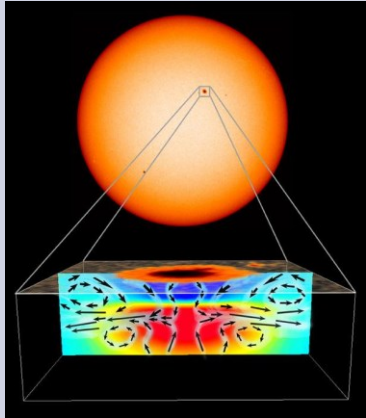

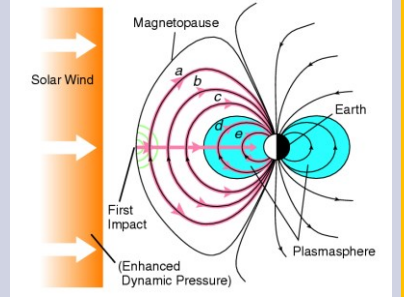
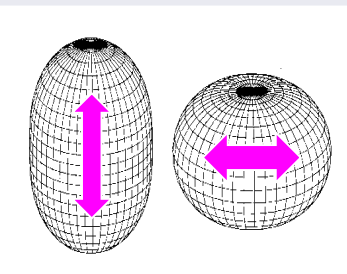
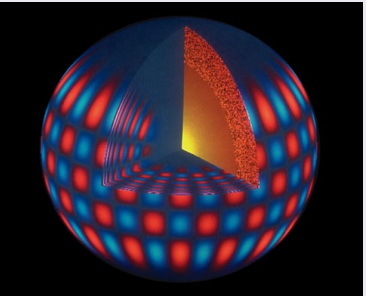
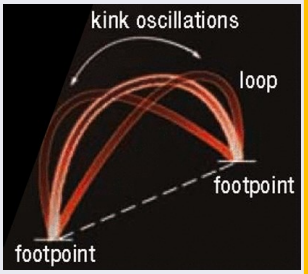
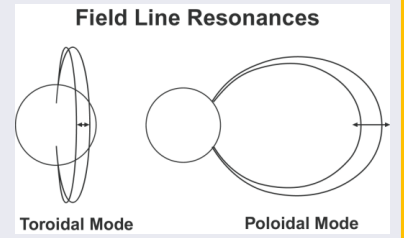
The Future of Ground-Based Research for
Magnetospheric and ITM Physics

March 15–16, 2022

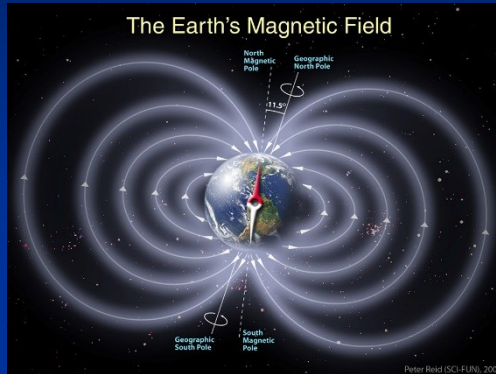


ISSI Magnetoseismology Team 2017-2018

Two Seismic Methods in Four Disciplines

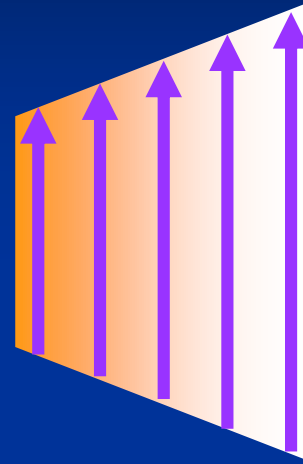
	<i>Seismology</i>	<i>Helio-/Astero- Seismology</i>	<i>Coronal Seismology</i>	<i>Magneto- seismology</i>
<i>Travel-time Method</i>				
<i>Normal- mode Method</i>				<p>Field Line Resonances</p>  <p>Toroidal Mode Poloidal Mode</p>

Normal-mode Magneto-seismology



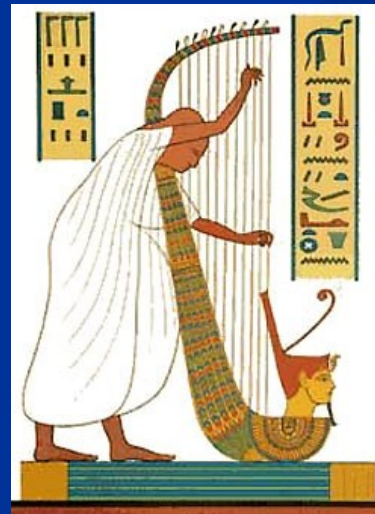
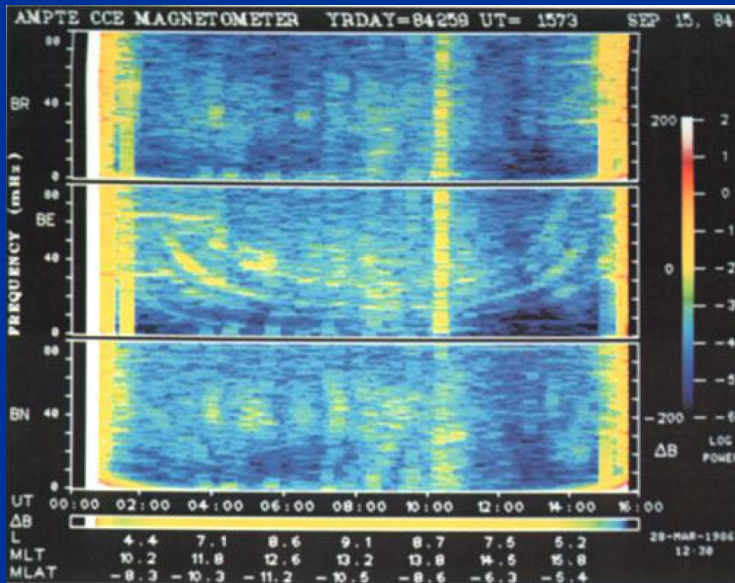
Satellite observations of resonant frequencies

\vec{B}



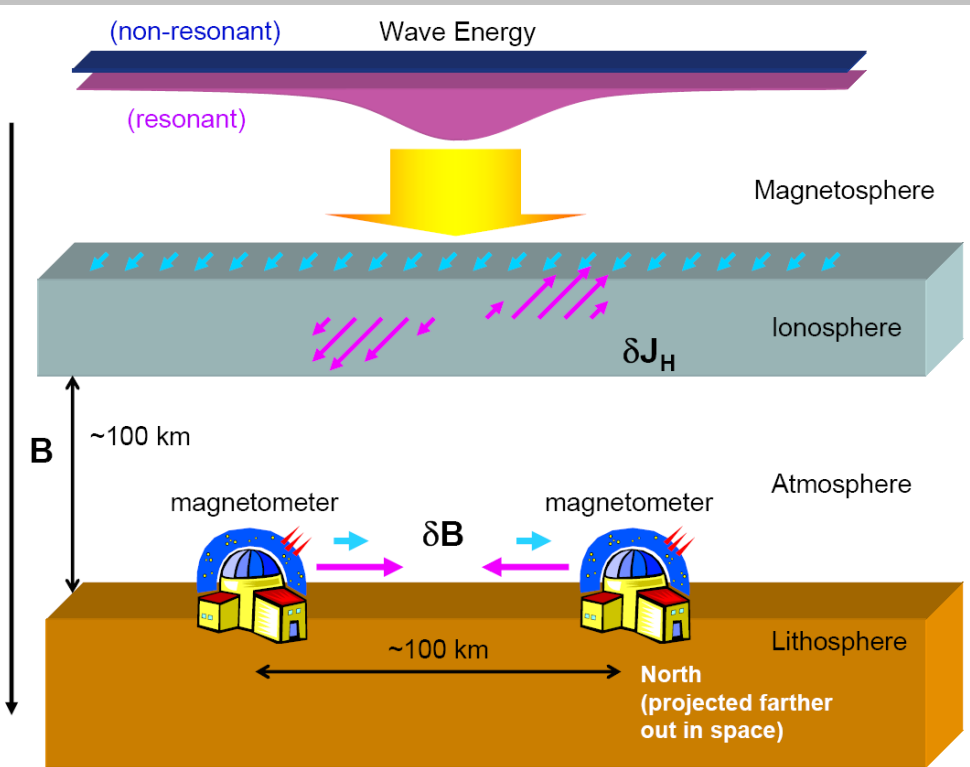
Resonant frequency:

$$f = \frac{V_A}{L} \propto \frac{B}{L\sqrt{\rho_m}}$$



- B : magnetic field (well modeled)
- L : length of field line (well modeled)
- ρ_m : plasma mass density (less known)

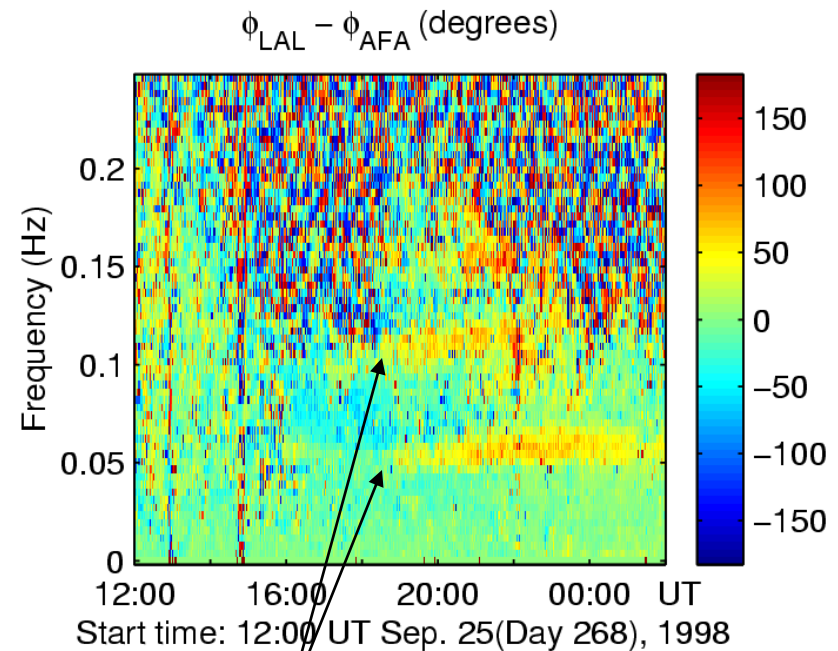
Detecting FLR frequencies from the Ground



distance bet'n stations \sim ionospheric height

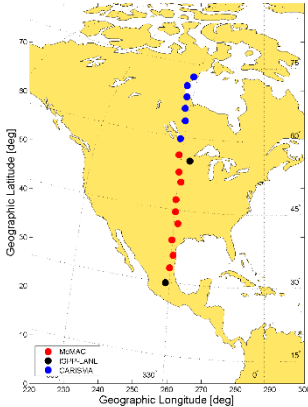
“The gradient technique”
(including “Cross-phase”)

Phase-Difference Spectrogram
(Between two stations in N-S direction)

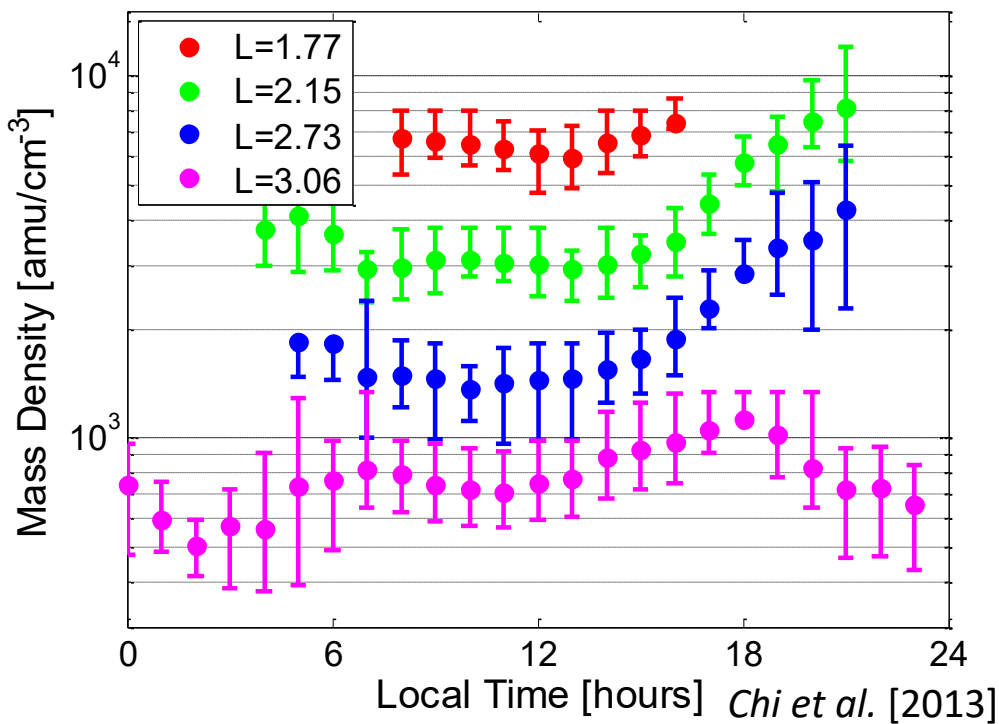


- The gradient technique can detect FLR frequencies for $> 80\%$ daytime hours

Thermospheric Effect on Plasmaspheric Density

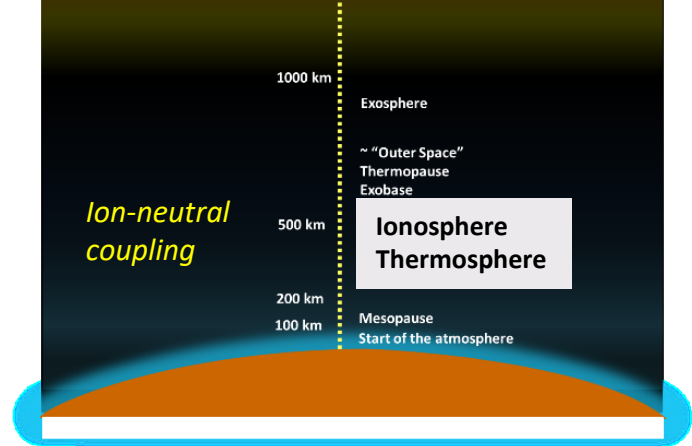
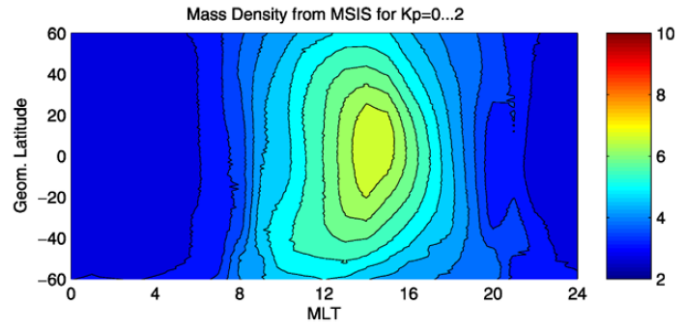
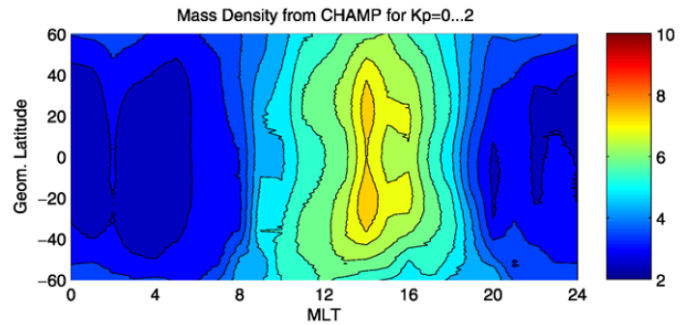


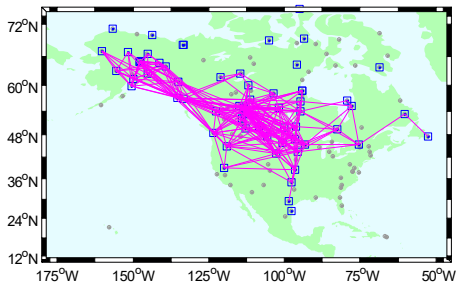
Plasmaspheric density estimated by normal-mode method



- Plasmasphere: Density increases in afternoon/evening hours are not predicted by models.

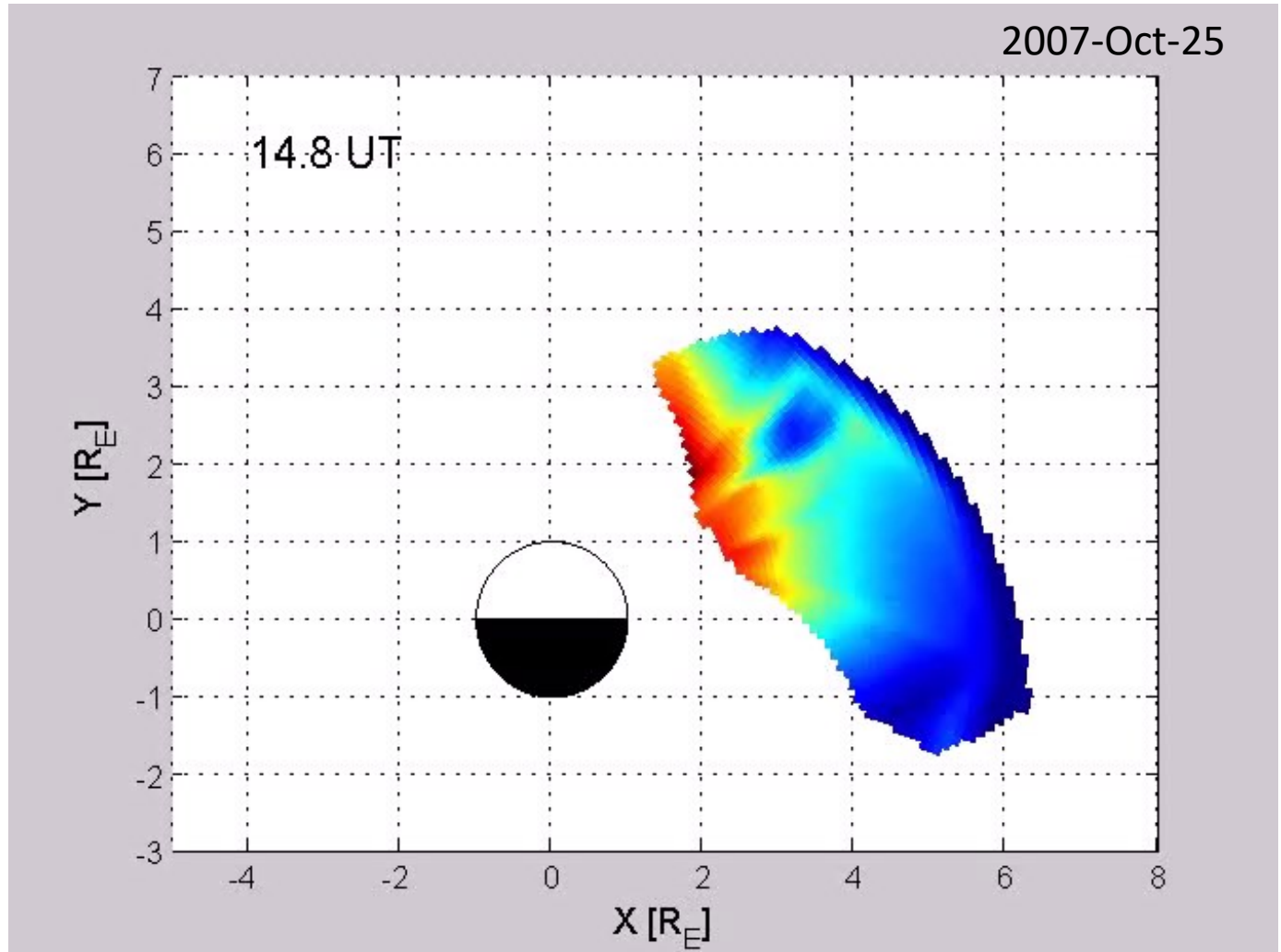
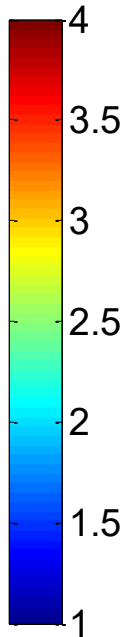
Thermosphere (neutral density)





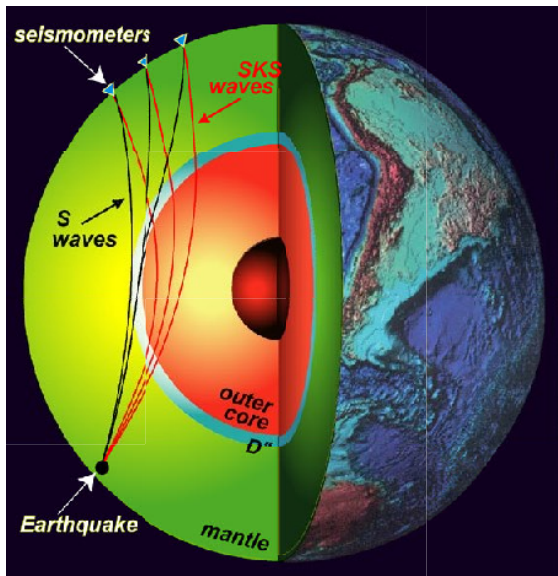
Monitoring Plasmaspheric Density by Ground-based Magnetometers (An example of

$\log_{10}(\rho_m)$
[a.m.u. cm^{-3}]



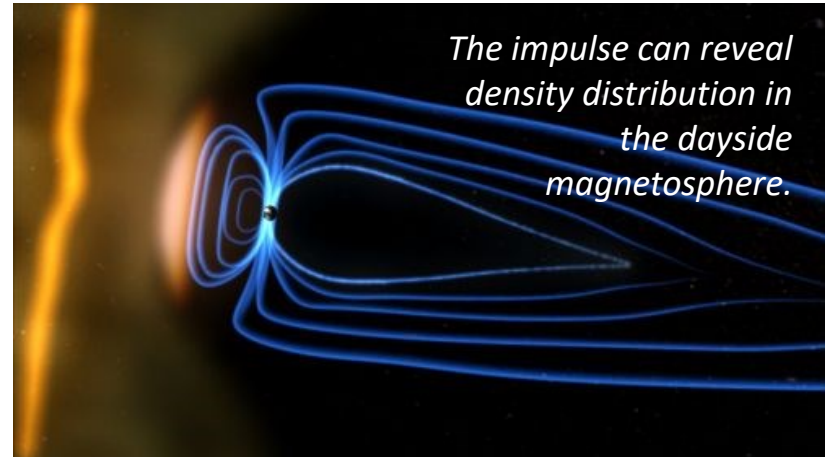
Travel-time Magneto-seismology

Earthquakes and seismic waves

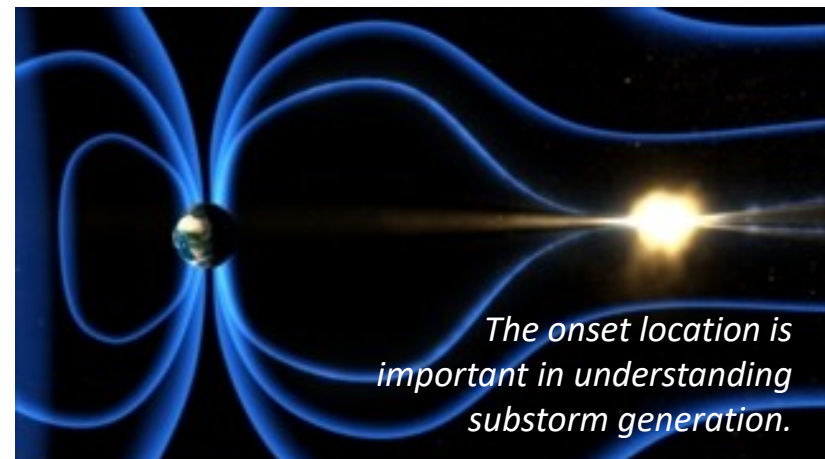


1. Interior structure of Earth
2. Center of Earthquake

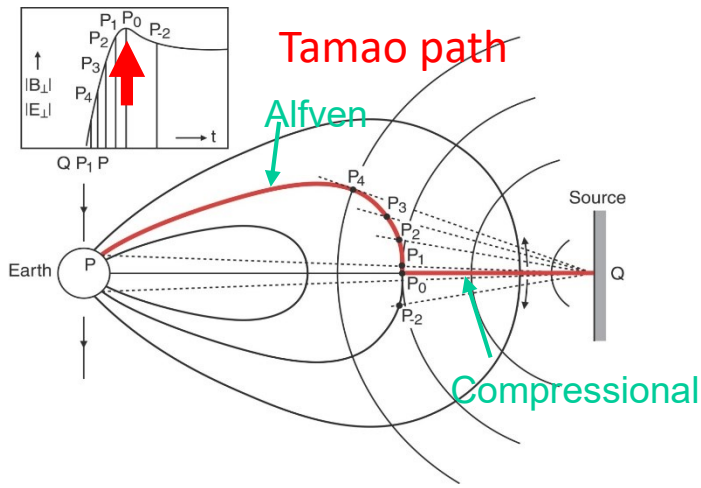
Sudden impulses



Substorm onsets

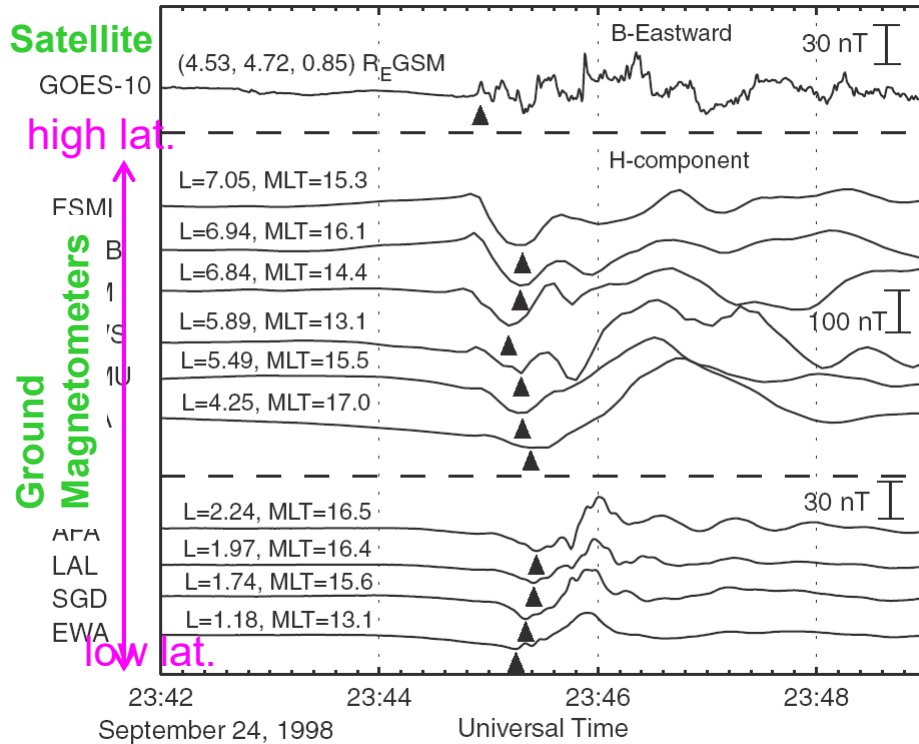


Travel-time Analysis for Sudden Impulses (Dayside)

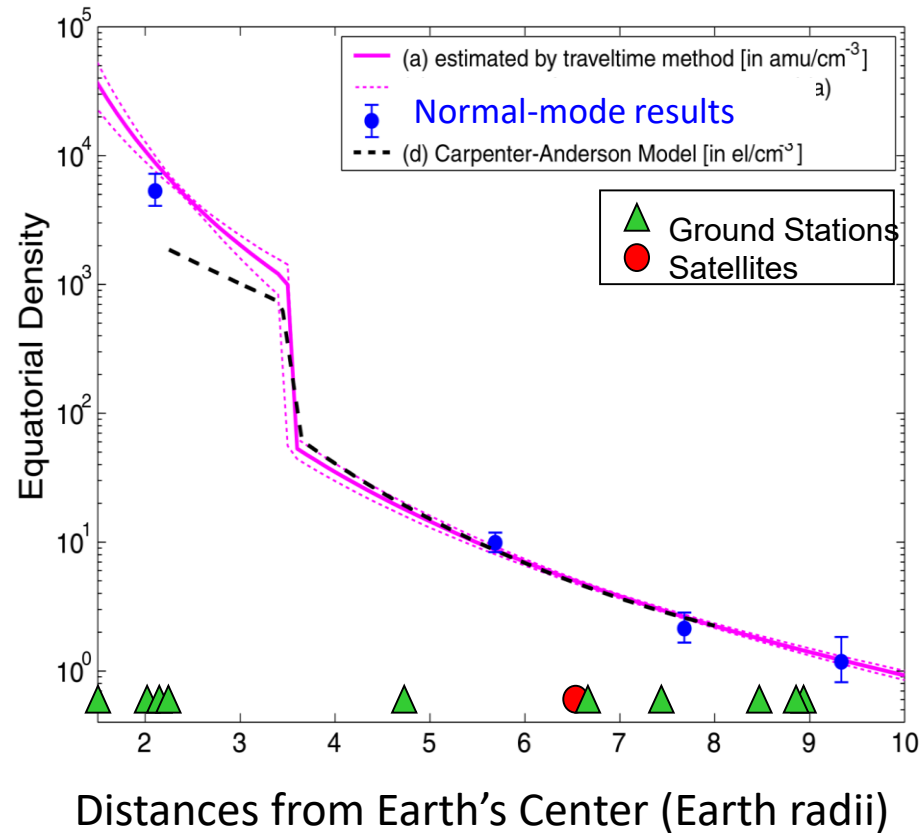


Inversion Results

Observations

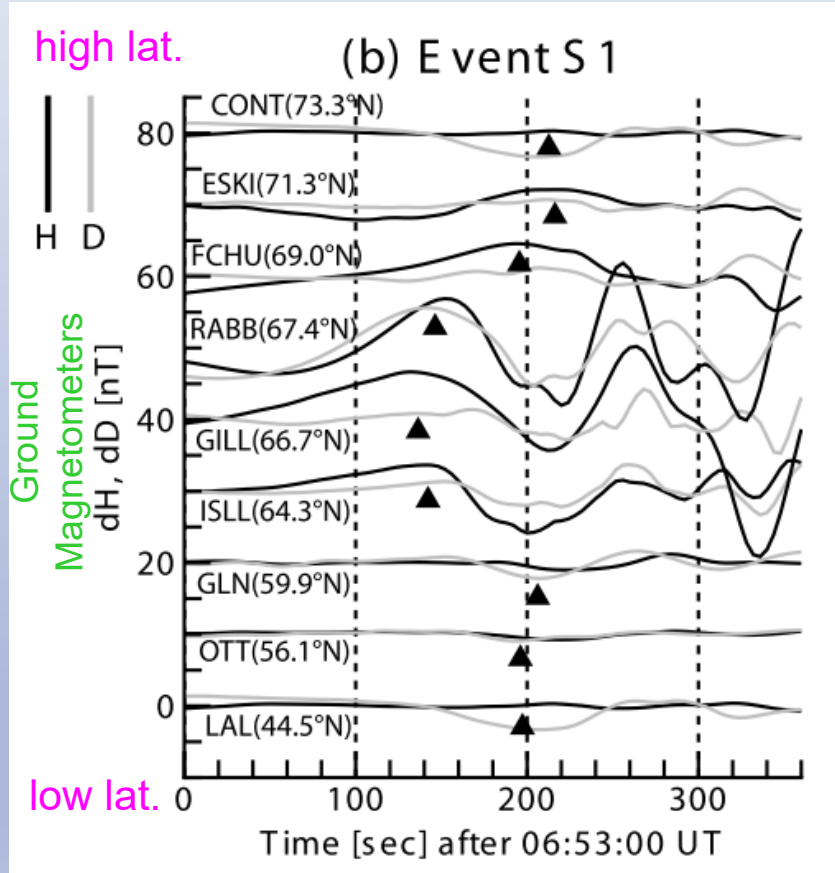


▲ Arrival Time

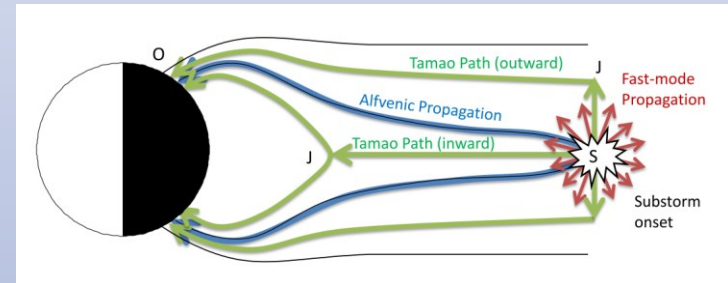
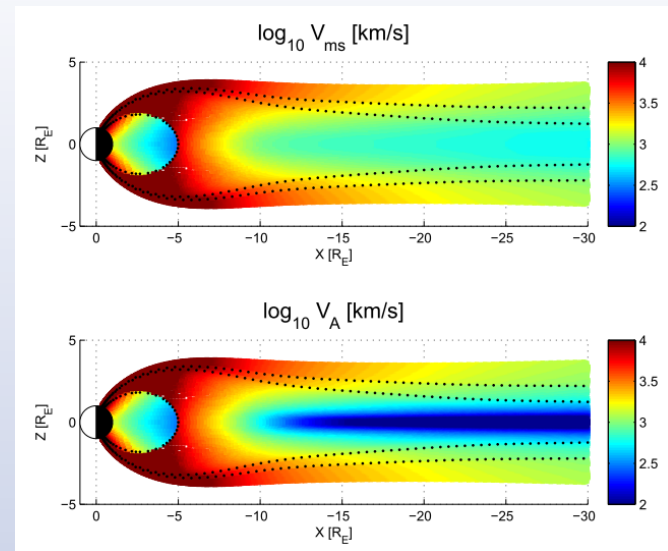


Travel-time Analysis for Substorm Onsets (Nightside)

Observations



▲ Arrival Time

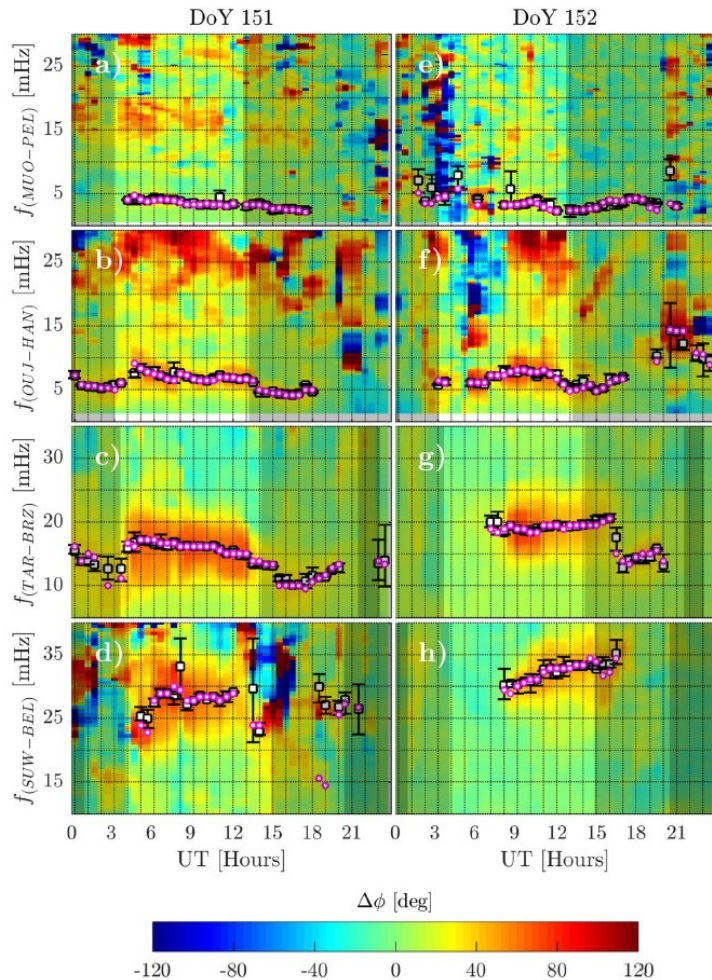


Inversion Results

- Locations of substorm onsets: **18-20 Earth radii** in the tail direction
- Substorm onset starts in the magnetotail **~ 100 --200 sec** before the first auroral brightening.

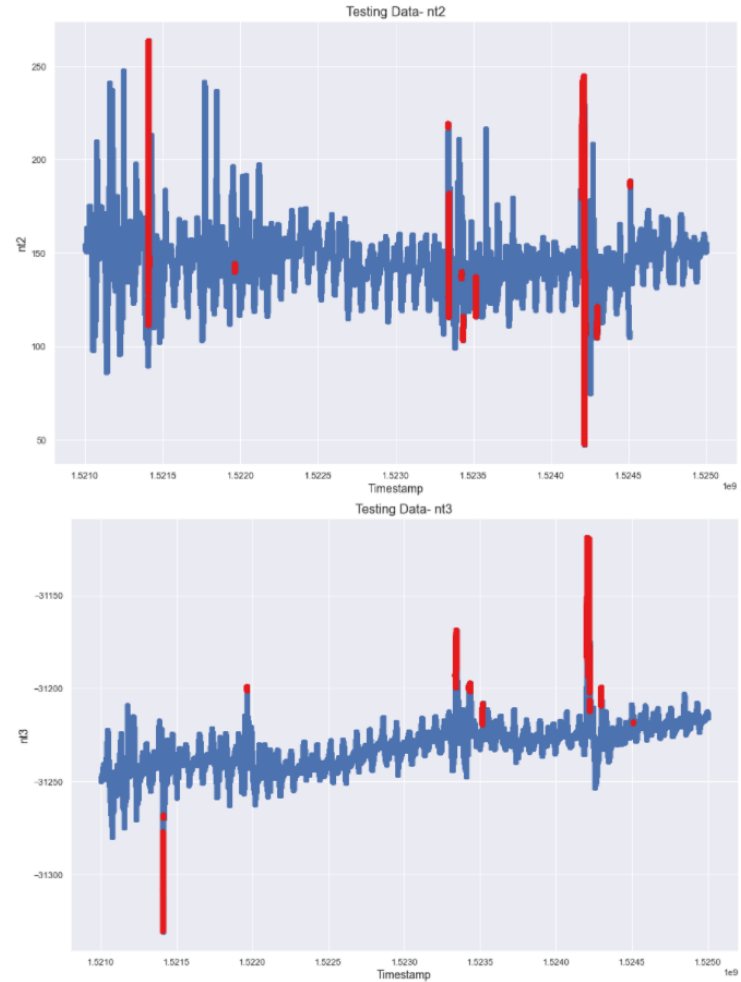
Machine Learning to Facilitate Big Data Analysis

Normal-mode magnetoseismology (FLR)



Foldes et al., JGR, 2021

Travel-time magnetoseismology



← 46 days →

Zhao and Chi, AGU presentation, 2021

Summary

- Magnetoseismology has become a well-developed discipline for new ways to investigate the plasma structures and dynamics of the magnetosphere.
- The normal-mode (FLR) method requires a dense 2-D magnetometer network in the region where magnetospheric field lines are closed.
- The travel-time method needs data from the continuous operation of global magnetometer networks.
- The success in using machine learning methods is enabling many future opportunities.