

RAMAN ANALYSES OF THERMOLABILE SAMPLES IN VARIOUS MARTIAN CONDITIONS: IMPLICATIONS FOR EXOMARS. G. Lopez-Reyes¹, R. Navarro¹, J.A. Manrique¹, A.Sanz¹ and F. Rull¹,
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Introduction: Raman spectroscopy is a very powerful technique for the analysis of all types of materials due to its versatility. The fact that for the past missions in planetary exploration none of them had included any kind of Raman device as part of the payload, makes Raman Laser Spectroscopy, RLS, a novel technique for planetary exploration that needs a deep evaluation in every kind of condition prior to its inclusion in future missions as MARS 2020 (SHERLOC, SuperCam) or EXOMARS (RLS instrument). The use of Raman spectroscopy for the identification and detection of traces of astrobiological interest has been demonstrated [1-3], showing the relevance of the use of this technique for planetary exploration missions such as ExoMars. Despite of the fact that Raman spectroscopy is considered a non-destructive technique, it is important to evaluate the damages that thermolabile samples can suffer due to the local heat caused by the excitation laser, and how the environmental conditions can introduce variations in the damage threshold of a sample.

The RLS instrument onboard the ExoMars rover: The Raman Laser Spectrometer instrument (RLS) onboard the ExoMars rover will fly to Mars in 2018. This rover will provide powdered samples to a suite of instruments inside the analytical laboratory drawer of the rover, including the RLS instrument. RLS will analyze up to 30 randomly selected points along a flat surface of the powdered samples. RLS will work autonomously by self-regulating the Raman acquisition parameters (integration time, number of accumulations) as a function of the sample under analysis. However, not all parameters can be adjusted during the operation phase, namely, the laser power and irradiance level on the sample. So, this parameter has to be fixed by design prior to flight. It is well known that thermolabile samples can be damaged (which means a reduction of the signal to noise ratio of the resulting spectra) if the energy provided by the laser cannot be thermally dissipated by the sample. Furthermore, this effect is worsened in low-pressure conditions, where the convective dissipation of the atmosphere is greatly reduced.

Analysis of thermolabile samples: In order to study the performance of the instrument with different laser irradiance levels, we have performed a series of analysis on several thermolabile sulfates and oxides, including hematite, jarosite, coquimbite, and also organic carbon. These tests were performed at different temperatures ranging from -25°C to -5°C, in a 6mbar CO₂ atmosphere, which will be the operational conditions in which the samples will be found inside the Ex-

oMars rover. A customized vacuum chamber was used to allow analyzing the samples under these conditions.

Though the behavior of the samples can be theoretically established, there are many factors that cannot be easily modeled in the theoretical calculations. Thus, this work is mainly aimed at experimentally observing the resilience of the samples with increasing laser irradiance on the sample surface, and the influence of the sample temperature on these results. The analysis is performed by evaluating the SNR values obtained with each different irradiance level. For these tests, 5 different points of each sample at 3 different temperature levels and 8 different irradiance levels were studied. This provided thousands of spectra that have been analyzed to obtain graphs and to assess the behavior of the samples vs the temperature and irradiance applied. Namely, the irradiance levels for which the best SNR values are obtained were calculated and represented. Figure 1 shows an example for a sample of coquimbite. Further analysis with more organic materials will need to be performed in order to assess the stability of these samples with respect to the laser irradiance levels.

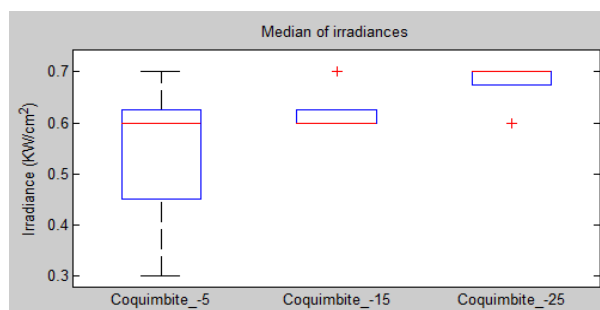


Figure 1. Best SNR irradiance level at different sample temperatures (-5 °C, -15 °C and -25°C)

References: [1] Vitek, P., et al. (2012) *Astrobiology* 12(12), 1095-1099. [2] Edwards, H. G., et al. (2013) *Astrobiology* 13(6), 543-549. [3] Manrique-Martínez, J., et al. (2014) *LPI*, 1783, 5091.