## MULTIPLE TECTONIC AND VOLCANIC EVENTS: GINA CRATER AREA, VENUS



Emily Roberts<sup>1</sup> (kroberts.emily@gmail.com), Allan Treiman<sup>1</sup>, Gabriel Eggers<sup>1</sup>, and Justin Filiberto<sup>2</sup>; <sup>1</sup>Lunar and Planetary Institute, Houston, TX, 77058; <sup>2</sup>Code XI, NASA Johnson Space Center, Houston, Texas.

Introduction: A current controversy in the geology of Venus centers on the age(s) of its highlands – the tesserae. One view of Venus' past is that it experienced a global resurfacing event at ~1.5 Ga, now represented mostly by volcanic plains, and that the tesserae represent earlier crust deformed in that event [1,2]. It is also argued that the resurfacing represents multiple volcanic events over long times [3,4]. The ancient age of tesserae has recently come into question [5,6]. Some tesserae include distinct morphologic units that could represent deformed plains material [5,6]; in other cases, it is possible that tesserae are forming today [7]. To address this question, we are mapping a tessera-plains transition around Gina Crater, near Venus' north pole.

Gina is a ~15 km diameter crater at 78.1°N, 76.3°E (Fig. 1), in the Snegurochka (V1) quadrangle [8]. Gina is on the western boundary of the Szél-anya Lineae belt (mapped mostly as tessera [8]), where it abuts a broad area of regional volcanic plains [9]. The area is complex, with evidence for multiple episodes of tectonism and volcanism and was specifically chosen to help constrain the timing of deformation events relative to those of volcanic emplacement.

> Q: Could it be possible that some tesserae formed geologically recently?

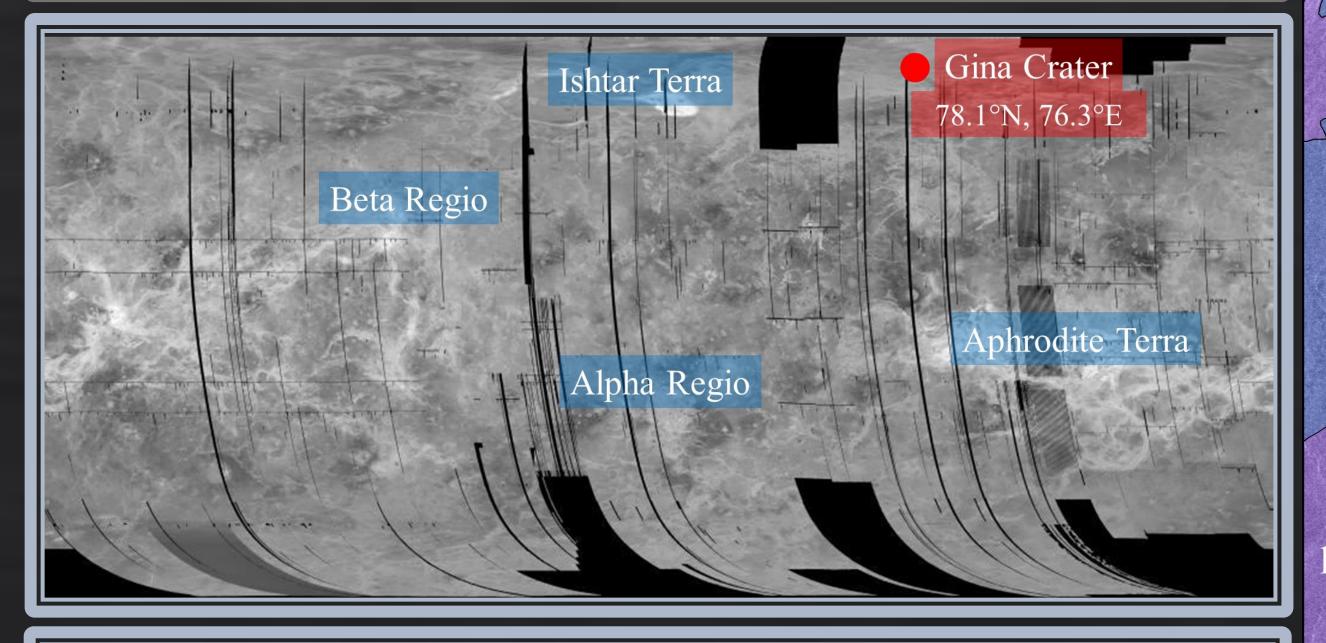
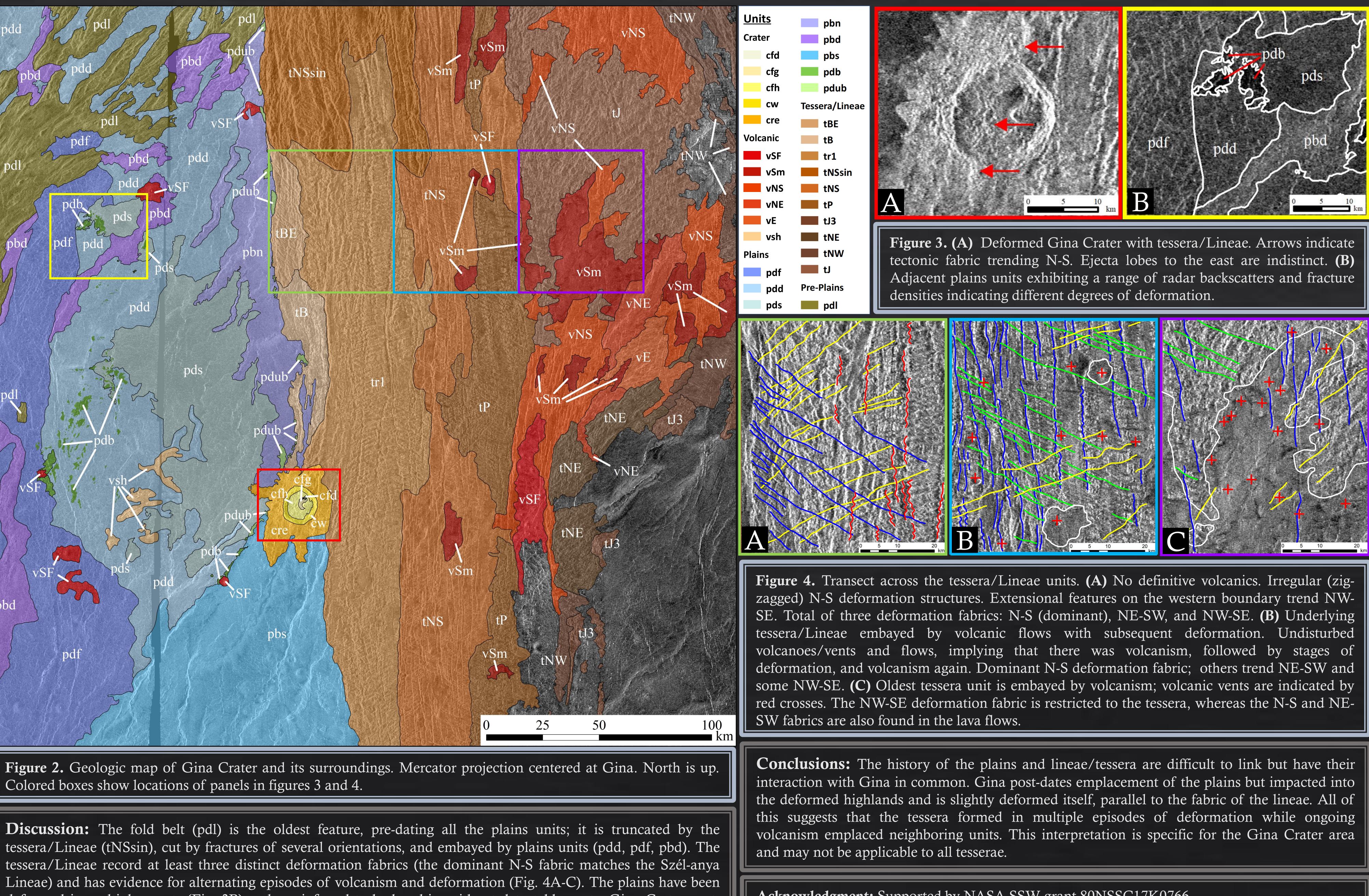


Figure 1. Location of Gina Crater at 78.1°N, 76.3°E.

**Data:** Magellan's SAR left-look global mosaic (~75 m/pix) was used as the basemap. ArcMap 10.6.1 was used for mapping. JMARS was used for general visualizations.

Methodology: Morphologic features are defined by shape, orientation, and SAR backscatter. Map units are defined based on differences in radar brightness, morphology, texture, and stratigraphic relations. Units are generally defined based on the radar brightness of emplaced material, but when material is obscured by deformation, the unit is instead characterized by that deformation. Contacts are defined by embayment relationships, radar brightness, deformation morphology, and the density of fractures.



deformed in multiple events (Fig. 3B) and are inferred to be basaltic, without detectable vents. Gina Crater postdates the emplacement of the plains and pre-dates some of the E-W compression of the tessera/Lineae (Fig. 3A).

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