Regional-Scale Lithospheric Recycling on Venus via Peel-Back Delamination

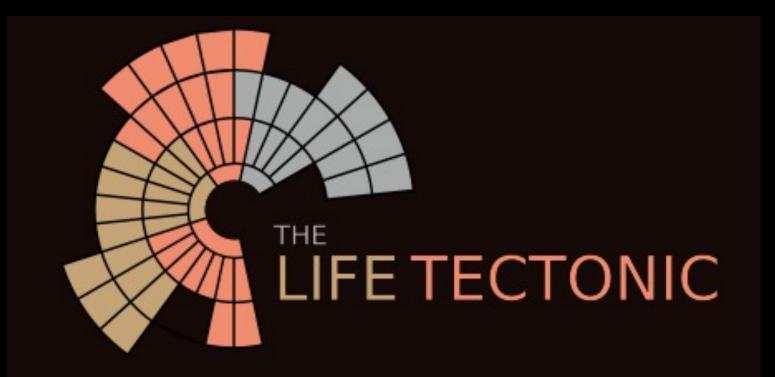
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Introduction

- Over 10,000 km of possible subduction sites have been identified on Venus, which have elastic thicknesses, plate bending moments and bending curvatures comparable to subduction zones on Earth (Schubert & Sandwell, 1995).
- Proposed subduction sites are located near (1) coronae and (2) groupings of rift-zone trenches called chasmata (Fig. 1).
- Previous numerical (Gülcher et al., 2020) and experimental (Davaille et al., 2017) studies have validated the possibility of regional-scale subduction initiation at coronae via plume-lithosphere interactions.
- Until now, no studies have tested the viability of subduction initiation at a chasma rift zone on Venus.
- Here, we present numerical experiments used to determine if and how regional-scale lithospheric recycling could be initiated at a Venusian rift zone.

Dali Chasma Latona Corona Fig. 1

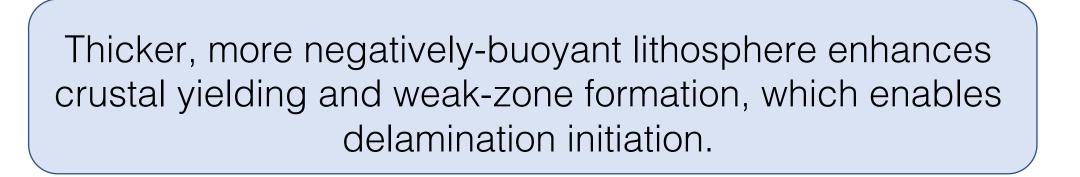
Schubert and Sandwell [1995]

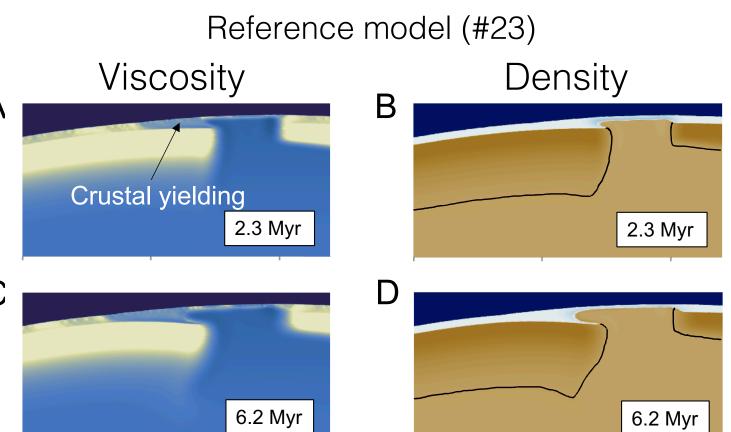
Model Setup

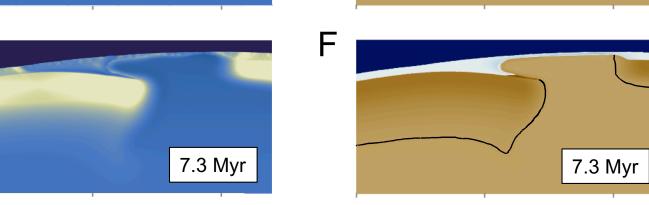
- 2D spherical annulus numerical models using StagYY (Tackley, 2008)
- Parameter space:
- Lithosphere thickness, $h_L = [200, 250, 300]$ km
- Maximum viscosity, $\eta_{max} = [10^{23}, 10^{24}, 10^{25}]$ Pa·s
- Compositional buoyancy of crust (i.e. $B_{crust} = \rho_{0,crust} \rho_{0,mantle}$)
 - $B_{crust} = [-175, -265, -300, -350, -400] \text{ kg/m}^3$
- Earth-like phase transitions were adjusted to shallower depths due to Venus's lower gravity (Ogawa & Yanagisawa, 2014) (Fig. 3) • Post-processing and visualization done using StagLab (Crameri, 2018)

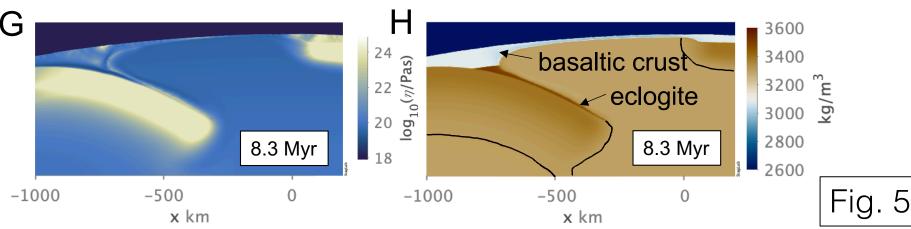
Mechanism for Delamination Initiation

- 30 km-thick layer of crust is relatively weak ($C_0 = 10$ MPa).
- Excess negative buoyancy of the lithospheric mantle results in initial plate bending, which causes the crust to yield (Fig. 5A)
- Crustal yielding reduces its viscosity, creating a weak- c zone delamination-surface to enable decoupling of the crust and lithospheric mantle.
- The length of the weak zone increases with increasing plate bending and crustal yielding (Fig. 5C).
- Eclogite formation at the base of crustal root (~70 km depth) helps sustain slab sinking (Fig. 5F, 5H).



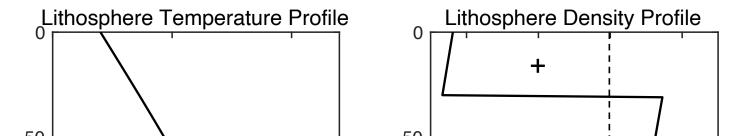






Calculating Plate Buoyancy

Reference model (#23)

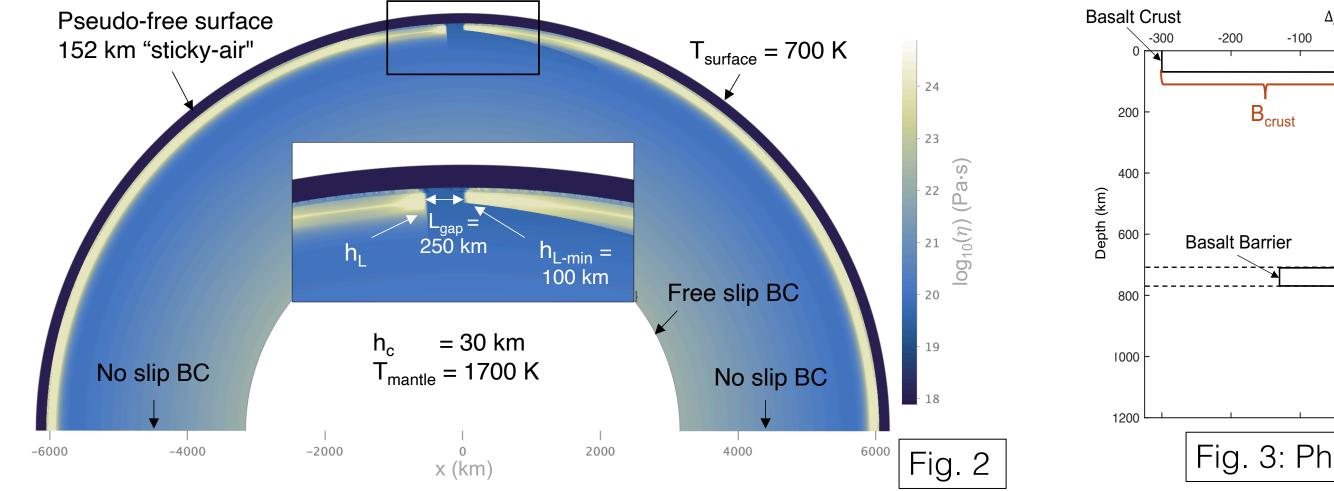


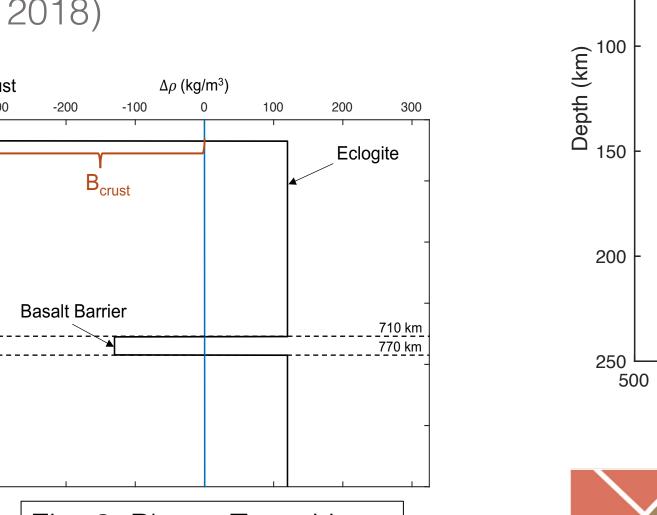
 \sim 100

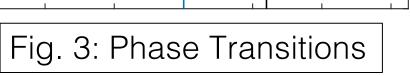
 $^{-150}$

200

Net plate buoyancy was controlled by two of the three variables in our parameter space: lithosphere thickness and crust density.

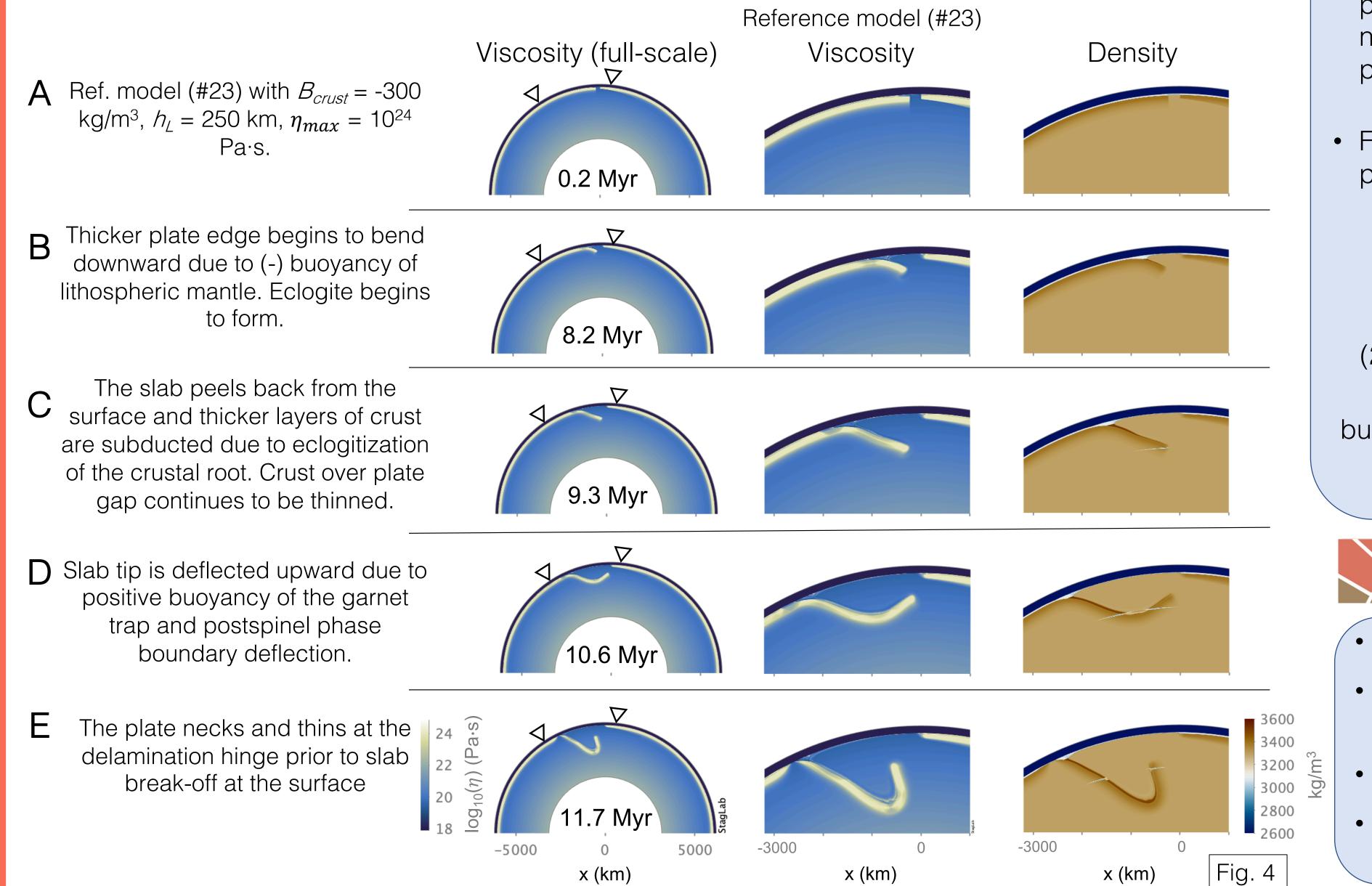






Peel-Back Delamination

- All lithospheric recycling events occurred via the peel-back delamination tectonic regime.
- Peel-back delamination is a form of lithospheric recycling which occurs when dense lithospheric mantle decouples from the crust and peels away, leaving behind a hot, thinned layer of crust at the surface.



- The total density of each plate was calculated as a function of depth, including both thermal and compositional components (Fig. 6). • Integrating the density profiles over depth with respect to the
 - reference mantle density gives a single value, $\Delta \rho_{plate}$, describing the net density contrast between the plate and the underlying mantle:

$$\Delta \rho_{plate} = \int_0^{h_L} (\rho(z) - \rho_0) dz$$

XIIIX Tectonic Regime Diagram

Fig. 6

3100 3200 3300 3400

Density (kg/m³)

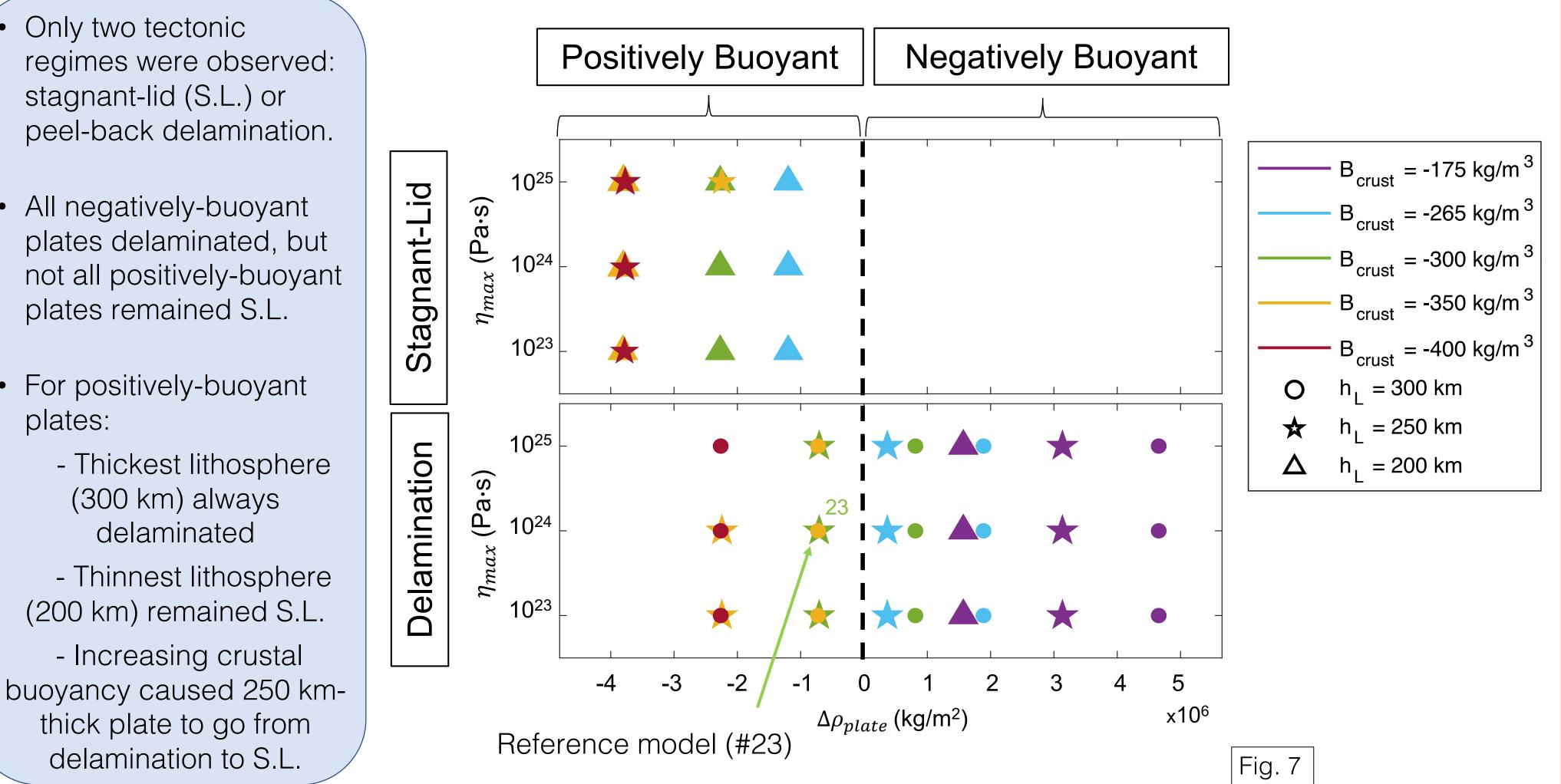
• Only two tectonic regimes were observed: stagnant-lid (S.L.) or peel-back delamination.

1000

Temperature (K)

1500

- All negatively-buoyant plates delaminated, but not all positively-buoyant plates remained S.L.
- For positively-buoyant plates:
 - (300 km) always



Conclusions

- Peel-back delamination is driven by the negative buoyancy of the lithospheric mantle.
- Delamination is resisted by (1) the coupling of the plate across the Moho, (2) the positive compositional buoyancy of the crust, and (3) increasing plate strength.
- Unlike subduction, delamination does not require net-negative plate buoyancy.
- Following a delamination event, the emplacement of hot, buoyant asthenosphere beneath the crust may have consequences for regional-scale volcanism and tectonic deformation.

References

- Crameri, F. "Geodynamic diagnostics, scientific visualization and StagLab 3.0." Geoscientific Model Development 11.6 (2018).
- Davaille, A., Smrekar, S. E., and Tomlinson, S. (2017). Experimental and observational evidence for plume-induced subduction on venus. Nature Geoscience, 10, 349-355.
- Gülcher, A. J., Gerya, T. V., Montési, L. G., and Munch, J. (2020). Corona structures driven by plume-lithosphere interactions and evidence for on-going plume activity on venus. Nature Geoscience, 13, 547-554.
- Schubert, G. and Sandwell, D. T. "A global survey of possible subduction sites on Venus." Icarus 117 (1995):173-196.
- Tackley, Paul J. "Modelling compressible mantle convection with large viscosity contrasts in a three-dimensional spherical shell using the yin-yang grid." Physics of the Earth and Planetary Interiors 171. 1-4 (2008): 7-18.

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