

The sodium ion enhancement at Mercury's high latitude magnetosphere during flux transfer event showers: MESSENGER observations

Weijie Sun¹, James A. Slavin¹, Anna Milillo², Stefano Orsini², Xianzhe Jia¹, Jim M. Raines¹, Stefano Livi^{1,3}, Jamie M. Jasinski⁴, Ryan M. Dewey¹, Yoshifumi Saito⁵, Changkun Li¹

wjsun@umich.edu

1, Department of Climate and Space Sciences and Engineering, University of Michigan, Ann Arbor, MI, USA

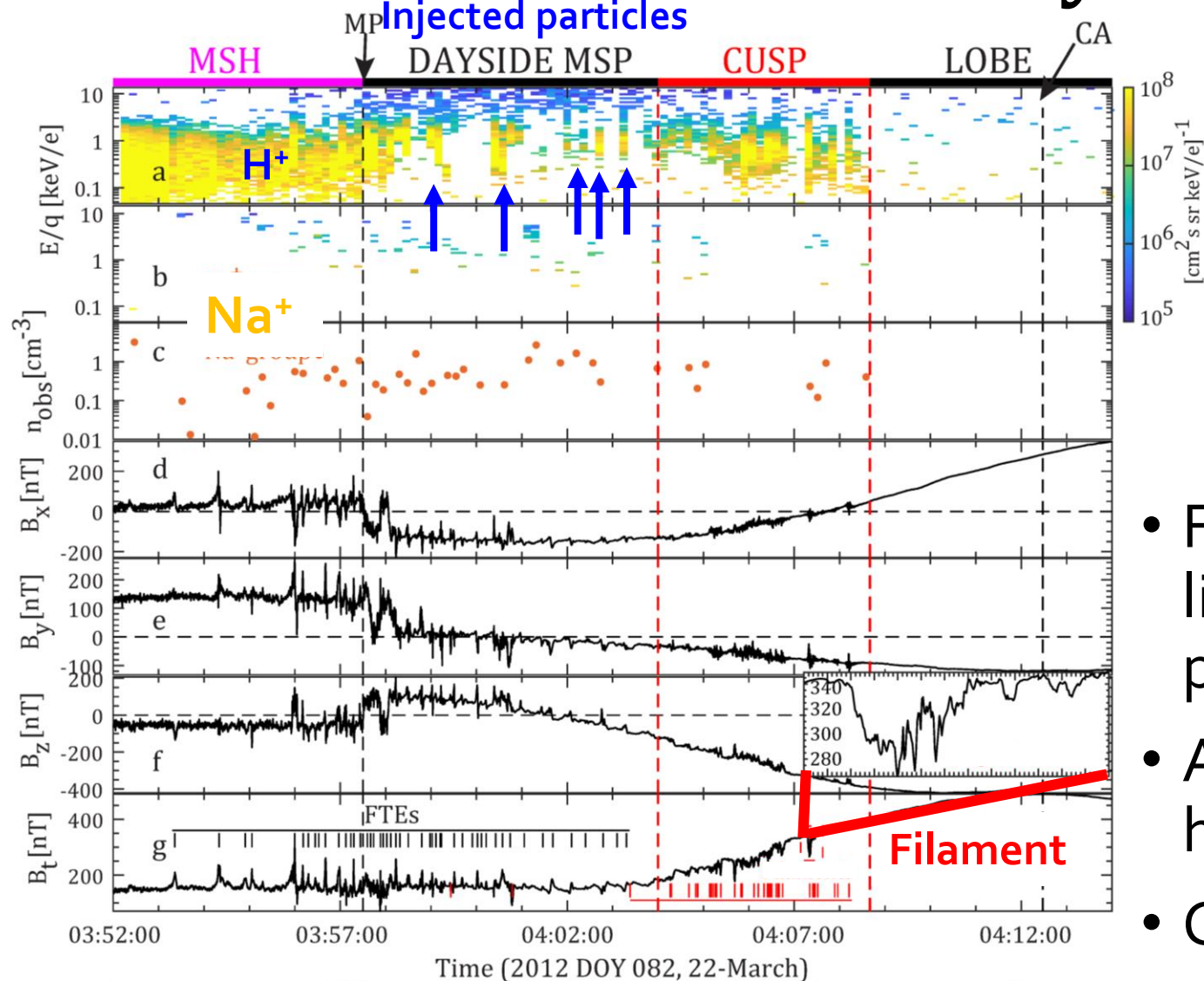
2, Institute of Space Astrophysics and Planetology, INAF, via del Fosso del Cavaliere 100, 00133, Rome, Italy

3, Southwest Research Institute, San Antonio, TX, USA

4, NASA Jet Propulsion Laboratory, California Institute of Technology, Pasadena, CA, USA

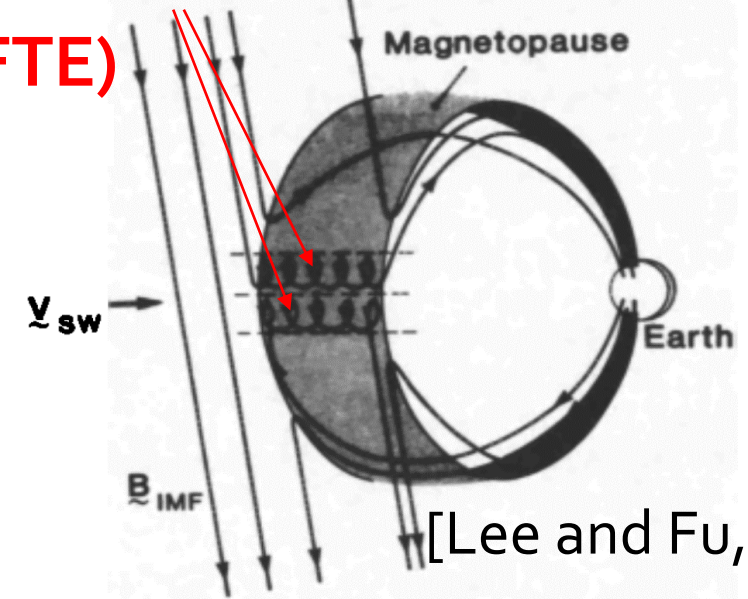
5, Japan Aerospace Exploration Agency, Institute of Space and Astronautical Science, Kanagawa, Japan

FTE Shower at Mercury



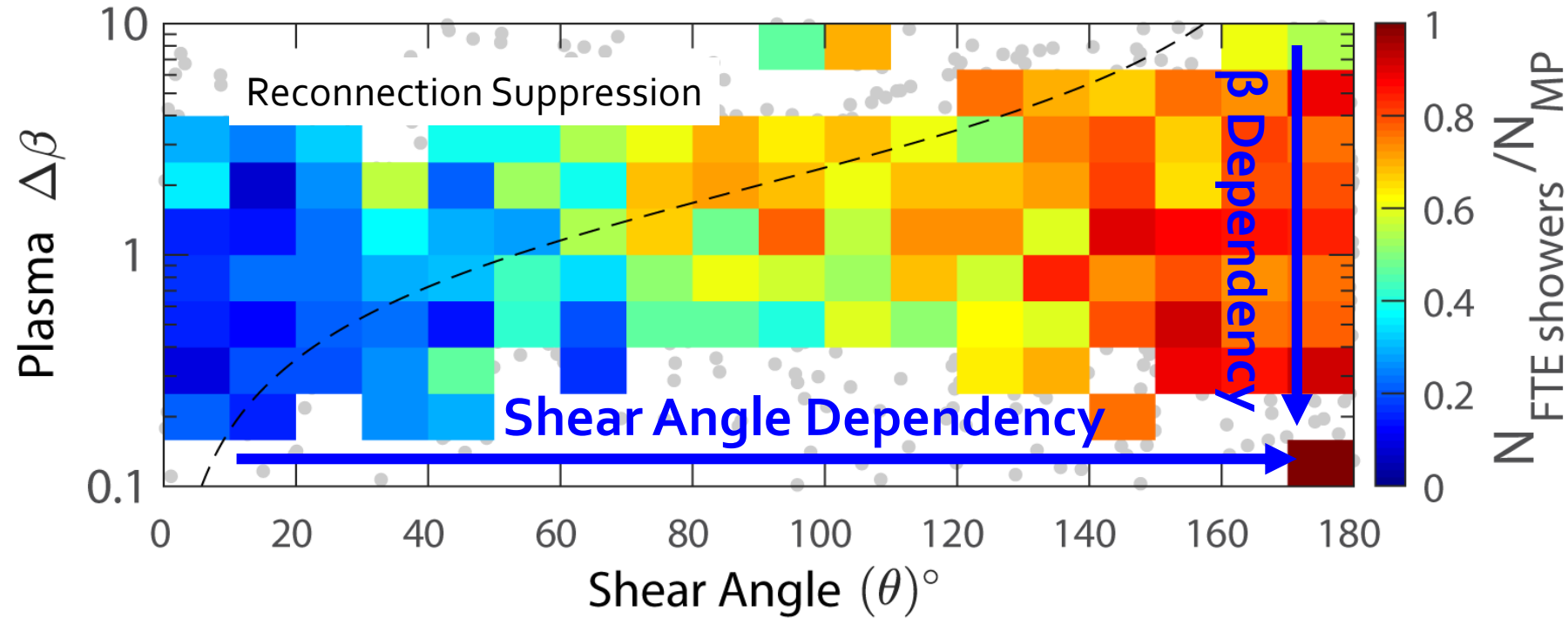
MLAT' [°]	23.9	36.7	54.0	76.0	72.4
LT' [hr]	12.2	12.3	12.6	14.0	22.4
ALT [km]	1470.6	1031.0	649.6	375.4	263.7

Flux transfer event (FTE)



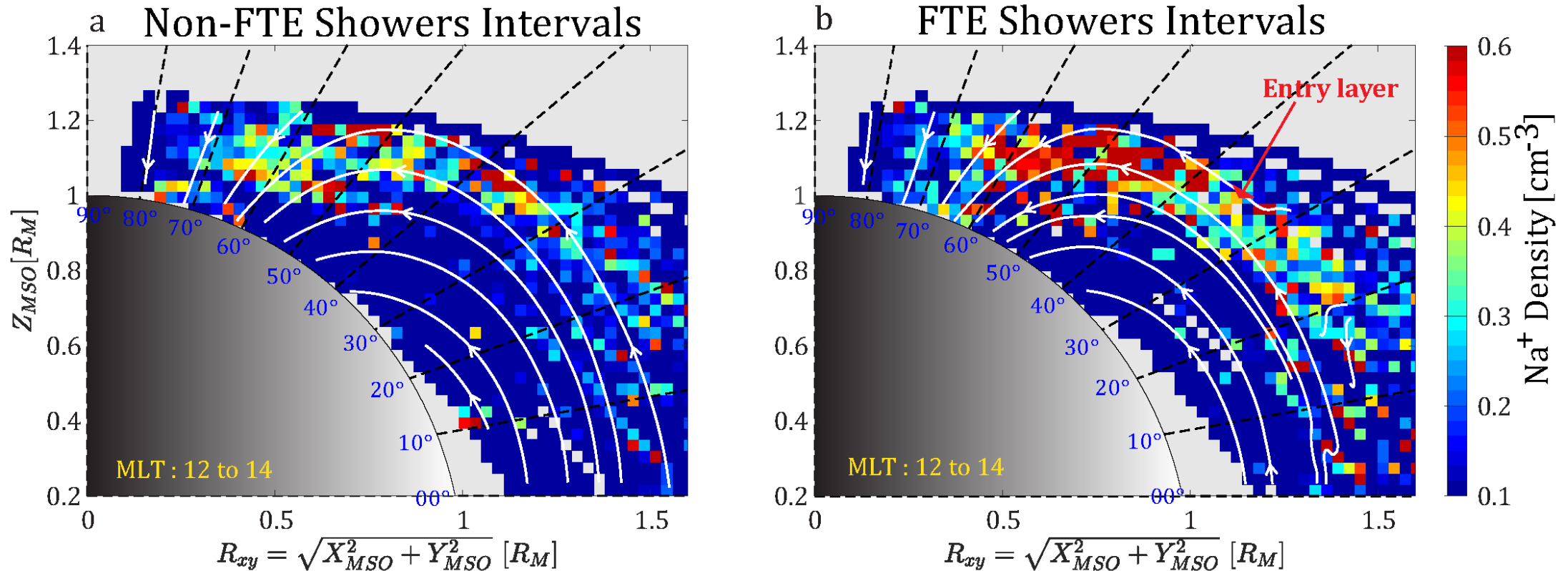
- FTEs contain open magnetic field lines, transport magnetic flux and particle flux
- An FTE "Shower": large-number, high frequent FTEs
- Cusp filament, diamagnetic decrease due to injected particles

Dependence on Plasma β and Magnetic Shear



- Sun et al. [2020, doi:10.1029/2020GL089784]: 1953 (~ 52%) of 3748 dayside magnetopause from MESSENGER correspond to FTE showers.
- The occurrence of FTE showers depends on plasma beta and magnetic shear angle

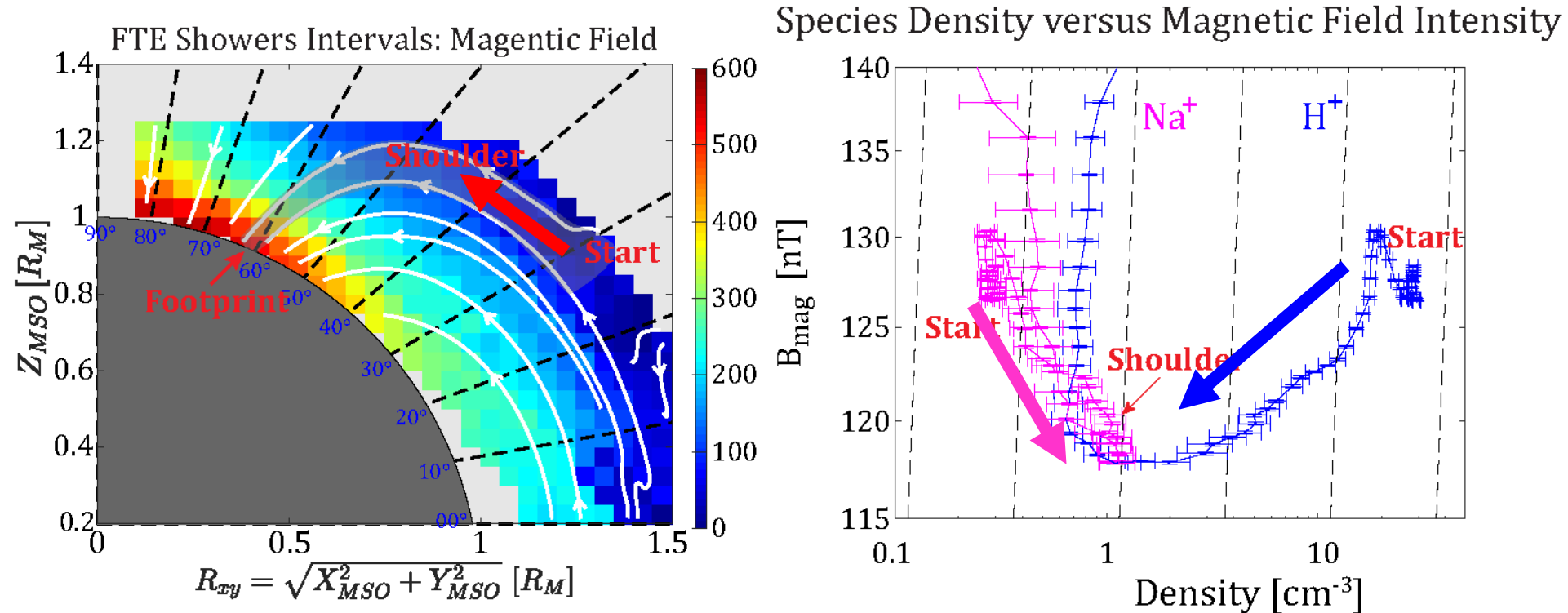
FTE Shower Influences Sodium Ions



FTE shower interval:

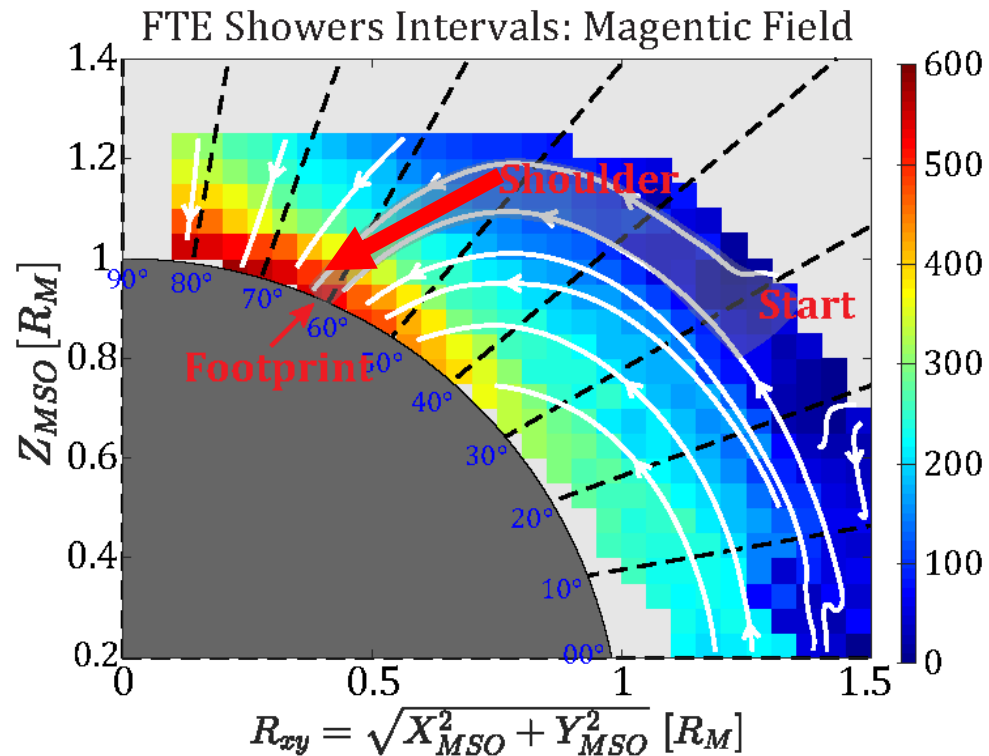
- Sodium-group ions (Na⁺) significantly enhance on the high latitude magnetosphere above the northern cusp.
- Entry Layer: the newly opened magnetic field lines.

Trace the Entry Layer: Start to Shoulder

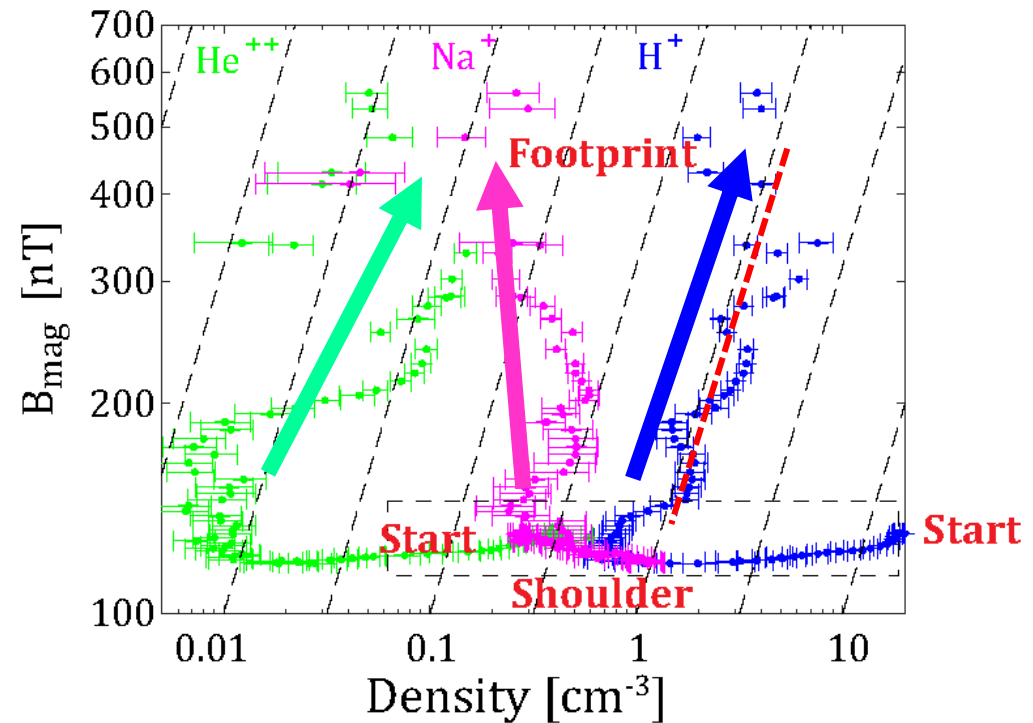


- H⁺ transported poleward. Na⁺-group accumulate from "start" to "shoulder".
- Solar wind particles lose while the Na⁺ increase, indicating the Na⁺ are continuously generated.

Trace the Entry Layer: Shoulder to Footprint

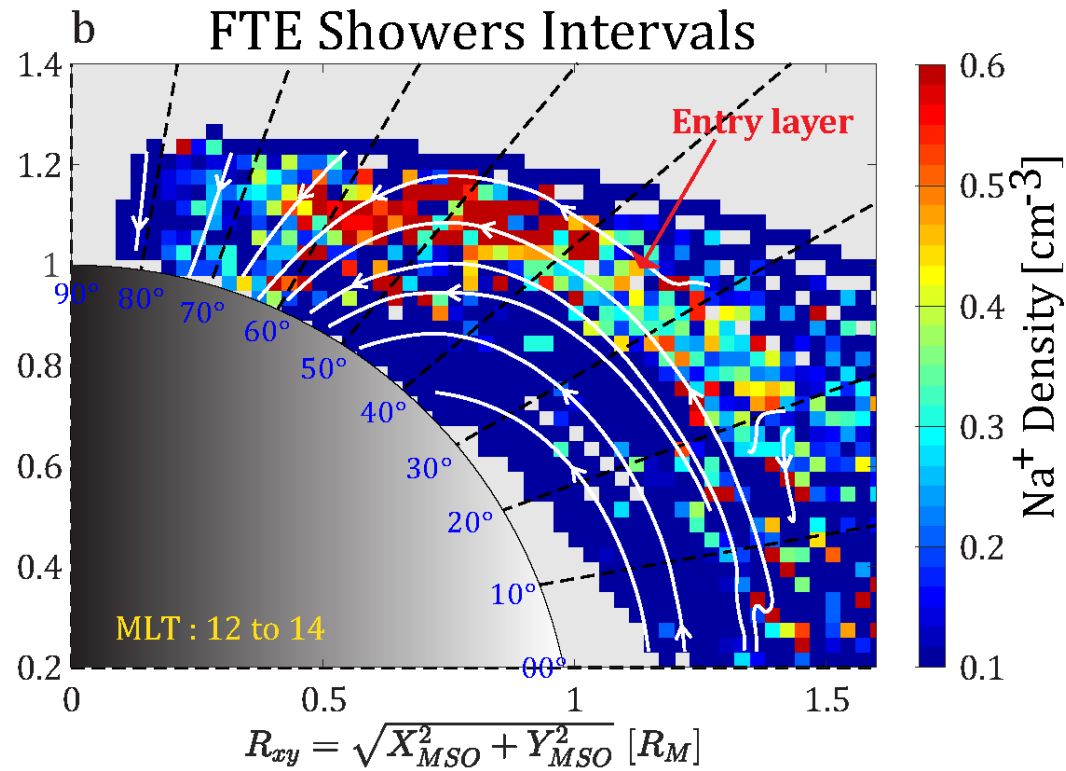


Species Density versus Magnetic Field Intensity

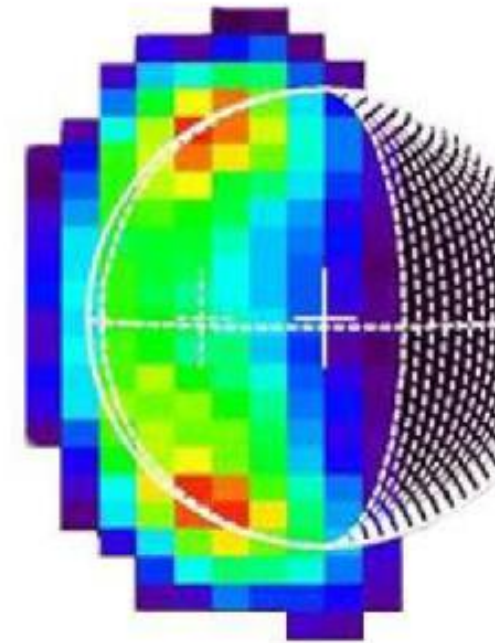


- $n_{\text{density}} / B_{\text{mag}} = \text{Const.}$ The plasma content is conserved along the same flux tube.
- The precipitation rate is $1.64 \times 10^{25} \text{ s}^{-1}$ for H⁺, and is $1.0 \times 10^{24} \text{ s}^{-1}$ for He⁺⁺

Sputtering Influence Exospheric dynamics



Double Peak (2P)



Driven by FTE shower?

The sputtered Na⁺: $6.0 \times 10^{23} \text{ s}^{-1}$ (a lower limit). Na emission patterns identified in

The neutral Na surface density: 10^9 m^{-3} .

Mercury's exosphere

[Mangano et al. 2015]

Then the sputtered column density: 10^{14} m^{-2} .

The background neutral Na is 10^{15} m^{-2} [Cassidy et al., 2015].

Conclusions

- FTE showers form a solar wind entry layer converging toward the cusp, which leads to a high estimated solar wind charged particle precipitation rate on the Mercury's surface ($> 10^{25} \text{ s}^{-1}$).
- The Na^+ is significantly enhanced in the entry layer during the intervals of FTE showers in ~ 10 minutes.
- FTE showers enhanced the precipitation rate and Na^+ in the high latitude region on time scales of ~ 10 minutes, which is comparable to the time scales of rapid variations in Mercury's exosphere near the magnetospheric cusps observed by ground-based telescopes.
- The exospheric enhancement response time due to FTE showers (~ 10 minutes) is much faster than what is expected for exospheric temporal variations due to stimulated desorption (\sim hours) or impact vaporization.