

## DELIVERY OF BODIES TO THE EARTH AND THE MOON FROM THE ZONE OF THE OUTER ASTEROID BELT

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**Introduction:** Delivery of material to the Earth from different distances from the Sun was studied by different scientists. In [1] I presented the probabilities of collisions with the Earth for bodies migrated from distances from the Sun from 5 to 40 AU. Migration of bodies-planetesimals with initial semi-major axes between 3 and 5 AU is considered below.

**Initial data:** The motion of bodies under the gravitational influence of 7 planets (from Venus to Neptune) was studied with the use of the symplectic code from [2]. In each variant of the calculations, the initial values of semi-major axes of orbits of 250 bodies-planetesimals varied from  $a_{\min}$  to  $a_{\min}+0.1$  AU, their initial eccentricities equaled to  $e_o=0.02$  or to  $e_o=0.15$ , and the initial inclinations equaled to  $e_o/2$  rad. The number of planetesimals with a semi-major axis  $a$  was proportional to  $a^{1/2}$ . The values of  $a_{\min}$  varied from 3 to 4.9 AU with a step of 0.1 AU. Based on the obtained arrays of orbital elements of migrated bodies, I calculated the probability  $p_E$  of a collision of a body-planetesimal with the Earth during time interval  $T$  (up to 5 Gyr in some variants). The calculations of  $p_E$  were made similar to the calculations presented in [3-5]. In each calculation variant the value of  $p_E$  is the ratio of the sum of the probabilities of collisions of 250 bodies with the Earth to 250.

**Results of calculations:** The value of  $p_E$  could vary by a factor more than a hundred for different calculation variants with 250 bodies and the same values of  $a_{\min}$  and  $e_o$ . Such difference was earlier found for calculations of migration of Jupiter-crossing objects [3-4]. One among hundreds or thousands of such objects moved in an Earth-crossing orbit during millions or even tens of millions of years, though the mean time of motion of a former Jupiter-crossing object in an Earth-crossing orbit was about 30 Kyr.

At  $3.0 \leq a_{\min} \leq 3.6$  AU or  $a_{\min}=4.2$  AU and  $e_o=0.02$ , and also at  $3.0 \leq a_{\min} \leq 3.1$  AU and  $e_o=0.15$ , more than a half of bodies still moved in an elliptical orbit after  $T=100$  Myr. At  $a_{\min}=4.2$  AU bodies were close to the Hilda family asteroids. At  $a_{\min} \geq 4.2$  AU and  $e_o=0.02$ , the values of  $p_E$  mainly were in the range from  $10^{-6}$  and  $10^{-5}$ , as for many calculations with  $a_{\min} \geq 5$  AU considered in [1]. At some other values of  $a_{\min}$  and  $e_o$ , the values of  $p_E$  could be much greater – up to the values of the order of  $10^{-3}$  at  $T=100$  Myr and of 0.01 at  $T=1000$  Myr. Though  $p_E=0$  at  $a_{\min}=3$  AU and  $e_o=0.02$ . It is not clear how much material was at distances from 3 to 4 AU from the Sun, compared to that in the zone of the giant planets. If we suppose that the density of a protoplanetary disk is proportional to  $R^{-0.5}$ , then the ratio of the mass of material with a distance  $R$  from the Sun between 4 and 15 AU is greater by a factor of  $\approx 30$  than that with  $R$  between 3 and 4 AU. For such a model, the amount of material delivered to the Earth from the zone of the outer asteroid belt could be comparable with the amount of material delivered from the zone of Jupiter and Saturn.

Initially Jupiter-crossing bodies that have come to the Earth's orbit did it mostly within the first million years. Most of collisions with the Earth of bodies, originally located at a distance from 4 to 5 AU from the Sun, occurred during the first 10 million years. At  $3 \leq a_{\min} \leq 3.5$  AU and  $e_o \leq 0.15$ , some bodies could fall onto the Earth in a few billion years. For example, for  $a_{\min}=3.3$  AU and  $e_o=0.02$ ,  $p_E=4 \times 10^{-5}$  at  $0.5 \leq t \leq 0.8$  Myr and  $p_E=6 \times 10^{-6}$  at  $2 \leq t \leq 2.5$  Myr. For  $a_{\min}=3.2$  AU and  $e_o=0.15$ ,  $p_E=0.015$  at  $0.5 \leq t \leq 1$  Myr, and  $p_E=6 \times 10^{-4}$  at  $1 \leq t \leq 2$  Myr. The zone of the outer asteroid belt can be one of the sources of the late heavy bombardment.

At  $a_{\min} > 3$  AU, the ratio of the number of bodies colliding with the Earth to that with the Moon was mainly in the interval from 16.4 to 17.4. This ratio varied mainly from 20 to 40 for planetesimals from the feeding zone of the terrestrial planets [5]. So more planetesimals per mass of a celestial body collided with the Moon than with the Earth. However, at collisions of planetesimals with the Moon the fraction of ejected material was greater than that with the Earth. The characteristic velocities of collisions with the Moon and the Earth of bodies in calculations with  $a_{\min}$  from 3 to 15 AU were mainly from 20 to 23 km/s and from 23 to 26 km/s, respectively.

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**References:** [1] Ipatov S.I. (2020) EPSC2020-71, <https://doi.org/10.5194/epsc2020-71>. [2] Levison H.F., Duncan M.J. (1994) *Icarus* 108:18-36. [3] Ipatov S.I., Mather J.C. (2004) *Annals of the NYAS* 1017:46-65. <http://arXiv.org/format/astro-ph/0308448>. [4] Ipatov S.I., Mather J.C. (2004) *Advances in Space Research* 33: 1524-1533. <http://arXiv.org/format/astro-ph/0212177>. [5] Ipatov S.I. (2019) *Solar System Research* 53: 332-361. <http://arxiv.org/abs/2003.11301>.