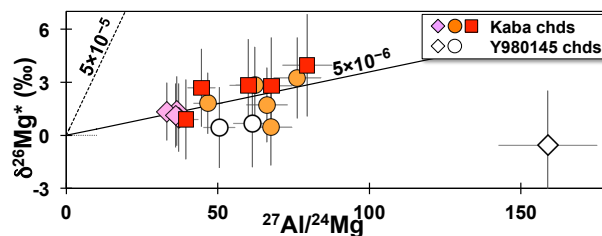


**$^{26}\text{Al}$ - $^{26}\text{Mg}$  SYSTEMATICS IN CHONDRULES FROM KABA AND YAMATO 980145 CV3 CHONDRITES**

K. Nagashima<sup>1</sup>, A. N. Krot<sup>1</sup>, and M. Komatsu<sup>2</sup>. <sup>1</sup>HIGP/SOEST, University of Hawai'i at Mānoa, USA, <sup>2</sup>The Graduate University for Advanced Studies, Japan. E-mail: kazu@higp.hawaii.edu.

**Introduction:** Short-lived radionuclide  $^{26}\text{Al}$  was likely the major heat source of planetesimals formed in the early Solar System. If  $^{26}\text{Al}$  was uniformly distributed in the protoplanetary disk, the initial abundance of  $^{26}\text{Al}$  in chondrules could constrain the accretion ages and thermal history of the chondrite parent bodies. Unlike chondrules from the least metamorphosed chondrites, chondrules from CV chondrites often show disturbed Al-Mg systematics [e.g., 1,2] and initial abundance of  $^{26}\text{Al}$  in the CV parent body is not well known. Here we report preliminary results of Al-Mg systematics in chondrules from two of the least metamorphosed CV chondrites, Kaba and Y-980145 [3,4].

**Results:** Among chondrules in 2 sections of Kaba and 1 section of Y-980145, we found several chondrules that have clean, unaltered plagioclase grains with high Al/Mg ratios. Their plagioclase compositions are typically An<sub>80-90</sub>. Al-Mg systematics of plagioclase in 3 and 2 type-I porphyritic chondrules from Kaba and Y-980145, respectively, were measured with the UH ims 1280 SIMS using the procedure described in [5]. As shown in a figure below, none of plagioclases in Y-980145 chondrules show resolvable excesses in  $\delta^{26}\text{Mg}^*$ . In contrast, most measurements in Kaba chondrules show elevated  $\delta^{26}\text{Mg}^*$  and about a half of them are resolvable from the terrestrial  $\delta^{26}\text{Mg}$ . These excesses are generally correlated with their Al/Mg ratios, indicative of *in situ* decay of  $^{26}\text{Al}$ . With the terrestrial  $\delta^{26}\text{Mg}$  as an intercept, model isochrons for the 3 chondrules from Kaba correspond to initial  $^{26}\text{Al}/^{27}\text{Al}$  ratios of  $(4.4\pm 2.1)\times 10^{-6}$ ,  $(4.8\pm 3.7)\times 10^{-6}$ , and  $(6.4\pm 2.8)\times 10^{-6}$ . Assuming uniform distribution of  $^{26}\text{Al}$  in the disk at the canonical level,  $(^{26}\text{Al}/^{27}\text{Al})_0 \sim 5.2\times 10^{-5}$  [6], these Kaba chondrules formed ~2–2.5 Myr after CV CAIs [7]. These formation ages are comparable to or slightly older than the model accretion age of 2.4–2.6 Myr that is required for CV parent body to reach its peak metamorphic temperature of 600°C [8]. The short time interval requires rapid accretion of chondrules into their parent body after their formation. No  $\delta^{26}\text{Mg}^*$  excesses in Y-980145 chondrules could be due to disturbance of Mg-isotopes in plagioclases due to thermal metamorphism [e.g., 9].



**References:** [1] Kita N. and Ushikubo T. 2012. *Meteorit. Planet. Sci.* 47:1108-1119. [2] Hutcheon I.D. et al. 2009. *Geochim. Cosmochim. Acta* 73:5080-5099. [3] Bonal L. et al. 2006. *Geochim. Cosmochim. Acta* 70:1849-1863. [4] Komatsu M. et al. 2014. Abstract #00370. The 5th Symposium on Polar Science. [5] Nagashima K. et al. 2014. *Geochem. J.* 48:561-570. [6] Jacobsen B. et al. 2008. *Earth Planet. Sci. Lett.* 272:353-364. [7] Connelly J. N. et al., 2012. *Science* 338:651-655. [8] Doyle P. M. et al. 2015. *Nat. Comm.* in press. [9] LaTourrette T. and Wasserburg G. J. 1998. *Earth Planet. Sci. Lett.* 158:91-108.