[W455]

## Wednesday, March 22, 2017 VOLCANIC PROCESSES ON SILICATE BODIES 1:30 p.m. Montgomery Ballroom

Chairs: William Garry

**Jeffrey Andrews-Hanna** 

- 1:30 p.m. Seeger C. H. \* Cox R.
  - <u>Geomorphology of Mountains on Io Provides Insight into Mountain-Patera Relationships</u> [#1675] Io's mountains are geomorphologically diverse, recording a range of erosion histories. The most degraded mountains are close to paterae.
- 1:45 p.m. Slezak T. J. \* Radebaugh J. Christiansen E. H.

  <u>Eigenshape Analysis of Planetary Craterforms: Implications for the Formation of Paterae on Io</u> [#2871]

  The shapes of craterforms formed by different processes are quantitatively analyzed and compared to infer possible formation mechanisms for paterae on Io.
- 2:00 p.m. Jawin E. R. \* Head J. W. Kreslavsky M. A. Wilson L.

  Stealth Dome in Central Oceanus Procellarum: Using Detrended Topography Data to Reveal the

  Hidden Past of the Prinz-Harbinger Region [#1184]

  Detrended topography data reveal a large dome in the Prinz-Harbinger region that implicate a multi-stage volcanic history for the region.
- 2:15 p.m. Haruyama J. \* Kaku T. Shinoda R. Miyake W. Kumamoto A. et al.

  \*\*Detection of Lunar Lava Tubes by Lunar Radar Sounder Onboard SELENE (Kaguya) [#1711]

  \*\*An intact lava tube at nearly 100 m depth extends a few tens of kilometers at Marius Hills on the Moon, based on SELENE Lunar Radar Sounder (LRS) echo data.
- 2:30 p.m. Needham D. H. \* Hamilton C. W. Bleacher J. E. Whelley P. L. Young K. E. et al.

  Lava Eruption and Emplacement: Using Clues from Hawai'i and Iceland to Probe the

  Lunar Past [#1177]

  Eruptions in Iceland and Hawai'i were studied to identify flow morphologies linked to pulses in vent activity that can thus be used to calculate eruption flux.
- 2:45 p.m. Garry W. B. \* Hughes S. S. Kobs Nawotniak S. E. Whelley P. L. Lim D. S. S. et al.

  \*\*Planetary Exploration of Lava Tubes with Lidar at Craters of the Moon, Idaho\*\* [#1207]

  We present findings that demonstrate the scientific and operational potential of lidar to explore planetary pits and discuss our results for Indian Tunnel lava tube.
- 3:00 p.m. Young K. E. \* Bleacher J. E. Needham D. H. Evans C. A. Whelley P. L. et al <u>Field Detection of Chemical Assimilation in a Basaltic Lava Flow</u> [#2706] Hawaiian lava / Chemically assimilates / Changing emplacement?
- 3:15 p.m. Bonnefoy L. E. \* Hamilton C. W. Scheidt S. P. Voigt J. Hoskuldsson A. et al. <u>Landscape Evolution After the 2014–2015 Lava Flow at Holuhraun, Iceland</u> [#1652]

  We compare topography before and after the emplacement of the 2014–2015 lava flow field in Holuhraun, Iceland to study associated landscape evolution processes.
- 3:30 p.m. Schaefer E. I. \* Hamilton C. W. Neish C. D. Sori M. M. Bramson A. M. et al. <u>Seeing Pāhoehoe from Orbit (Without Squinting)</u> [#2343]

  We are developing a method to identify morphologic lava flow type (e.g., pāhoehoe) from even coarse (>80 m/pixel) orbital imagery using fractal analysis.

- 3:45 p.m. Whelley P. L. \* Richardson J. A. Hamilton C. W.

  <u>Lava Channel Textures in Tartarus Colles, Elysium Planitia, Mars</u> [#2491]

  Roughness patterns, derived from HiRISE DTMs, are used to infer lava textures and flow emplacement
- dynamics on Mars.
- 4:00 p.m. Christoph J. M. \* Garry W. B.

  Spatial and Temporal Relationships Among Low Shield Volcanoes in the Ceraunius Fossae Region of Tharsis: The Last Gasp of Martian Volcanism [#2798]

  We mapped and dated lava flows from low shield volcanoes at Ceraunius Fossae in Tharsis, finding that these are some of the youngest volcanoes studied on Mars.
- 4:15 p.m. Andrews-Hanna J. C. \* Soto A.

  <u>Climatic Control Over Explosive Volcanism on Mars</u> [#2867]

  Climate models and geophysical models support the possibility that periodic ice deposition on martian volcanoes led to climatic control over explosive volcanism.
- 4:30 p.m. Ojha L. \* Lewis K. Karunatillake S.

  <u>Volcanic Origin of Medusae Fossae Formation from Gravity and Topography Data</u> [#2475]

  We localize the gravity and topography signature of MFF and place a direct constraint on its density.