**AN EXOLIFE FINDER TELESCOPE (ELF) FOR PROX B AND BEYOND.** J. R. Kuhn<sup>1</sup>, <sup>1</sup>Institute for Astronomy, University of Hawaii (34 Ohia Ku St. Pukalani, HI USA 96768, <u>jeff.reykuhn@yahoo.com</u>), S. V. Berdyugina <sup>2</sup>Kiepenheuer-Institut fur Sonnenphysik, (Freiburg, Germany, berdyugina@kis.uni-freiburg.de), D. Halliday <sup>3</sup>Dynamic Structures Ltd. (Vancouver, Canada, <u>starcluster18@gmail.com</u>), G. Moretto <sup>4</sup>CRAL/CNRS Ecole Normale Superieure (Lyon, France, <u>gil.moretto@univ-lyon1.fr</u>), M. Langois <sup>5</sup>CRAL/CNRS Ecole Normale Superieure (Lyon, France, maud.langlois@obs.univ-lyon1.fr)

Introduction: Post Keck-era large optical telescopes will combine elements of interferometry with large, light-weight subapertures[1]. Mass and material stiffness will be achieved in light-weight active structures that include metamaterial reflecting subapertures and fast algorithms that use in-field wavefront references. For small field-of-view observations with sufficiently bright host stars it is possible to build dedicated coronagraphic telescopes optimized for exoplanet studies with 50m diameter or larger effective apertures [2]. During the last two years a group of scientists and engineers from Canada, France, and Hawaii have been developing a 25m telescope concept that would allow photometric light-curve inversion of a Prox-b-like exoplanet to recover surface structure with sufficient resolution to detect Earth-like continents and a variety of atmospheric biomarker signals[3]. It is expected that this instrument could be built for an order of magnitude less than currently planned Keck-era astronomical telescopes of the same aperture that are not optimal for exoplanet direct imaging. This presentation summarizes the "ExoLife Finder" (ELF) and its 5-year timeline to completion.



**References:** [1] Kuhn, J. R. et al. (2014) *SPIE,* 9145, 91451G (8pp), [2] <u>http://www.the-colossus.com/</u>, [3] http://www.planets.life/